



EUROPEAN ARC
ALMA Regional Centre || IRAM



ALMA Correlator Cycle 0

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Version 3 07.06.2011



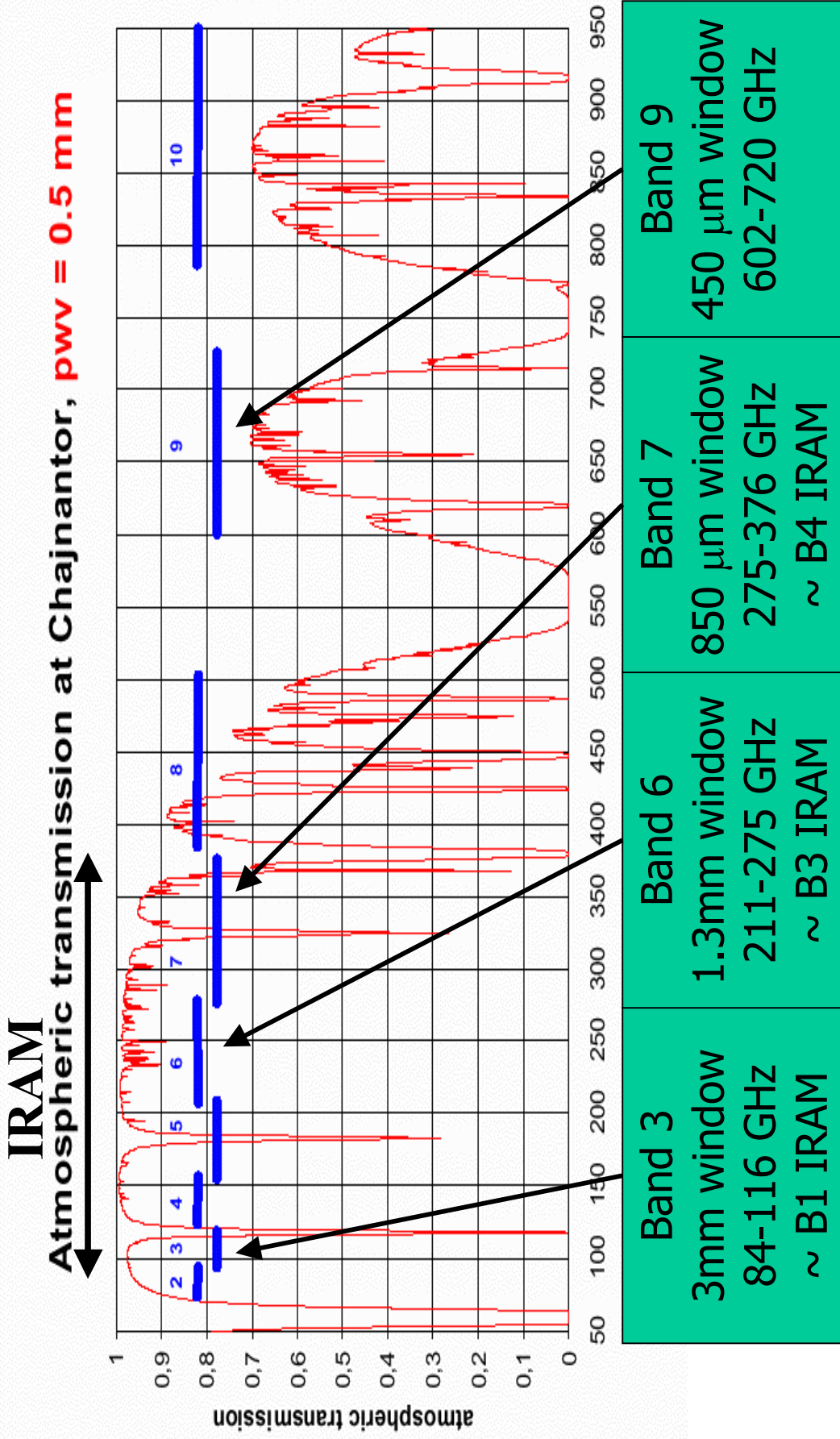


Aim of this document

- This document aims at presenting a user-oriented description of the ALMA correlator modes for Cycle 0.
- It contains a few comparison with the IRAM system (which can be skipped by anybody not familiar with the Bure system).
- This document is **not a manual of the Observing Tool** – how to define a frequency/correlator setup with the OT is not described.



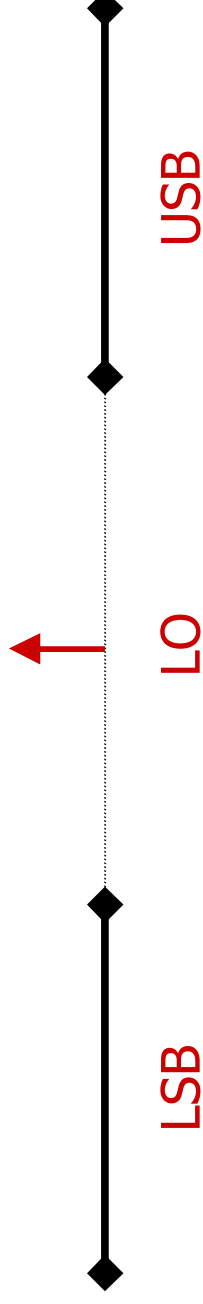
Receiver bands





Receiver bandwidth

Heterodyne receivers are sensitive to Lower Side Band and Upper Side Band

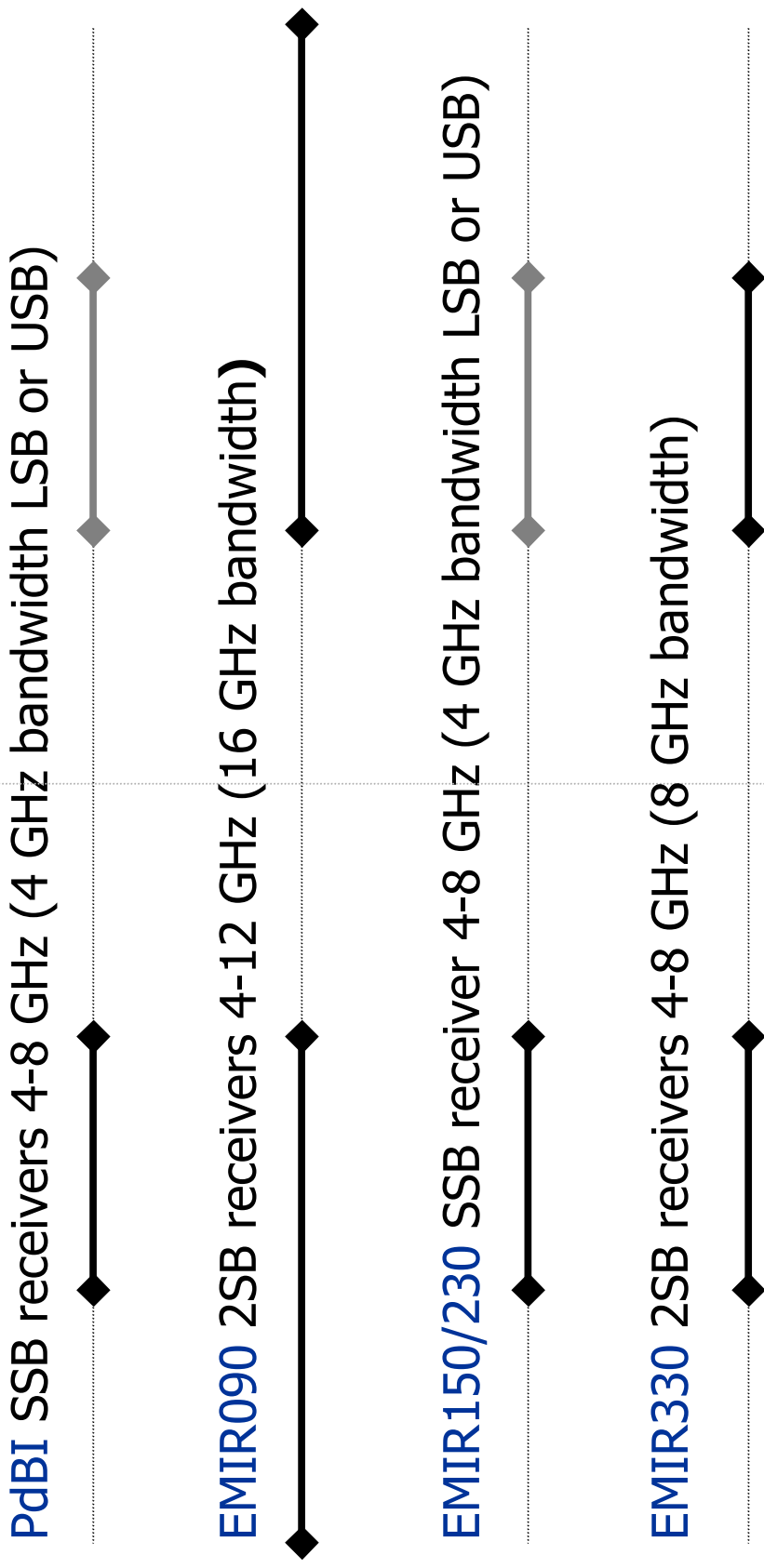


Receivers can be

- **DSB** outputs the sum $LSB + DSB$ → separated in the correlator
- **SSB** outputs **LSB or DSB**
- **2SB** outputs **LSB and DSB** separately



IRAM receivers

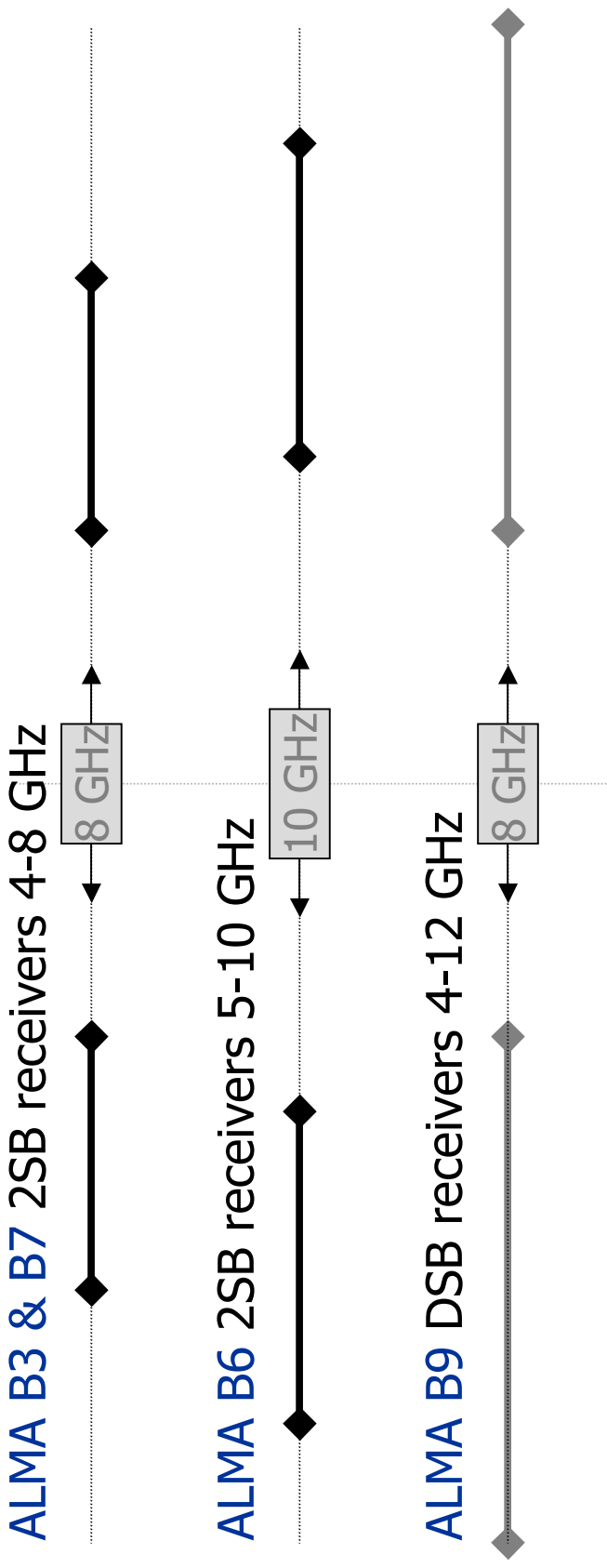


Receivers have 4 to 16 GHz bandwidth

x 2 polarizations



ALMA receivers

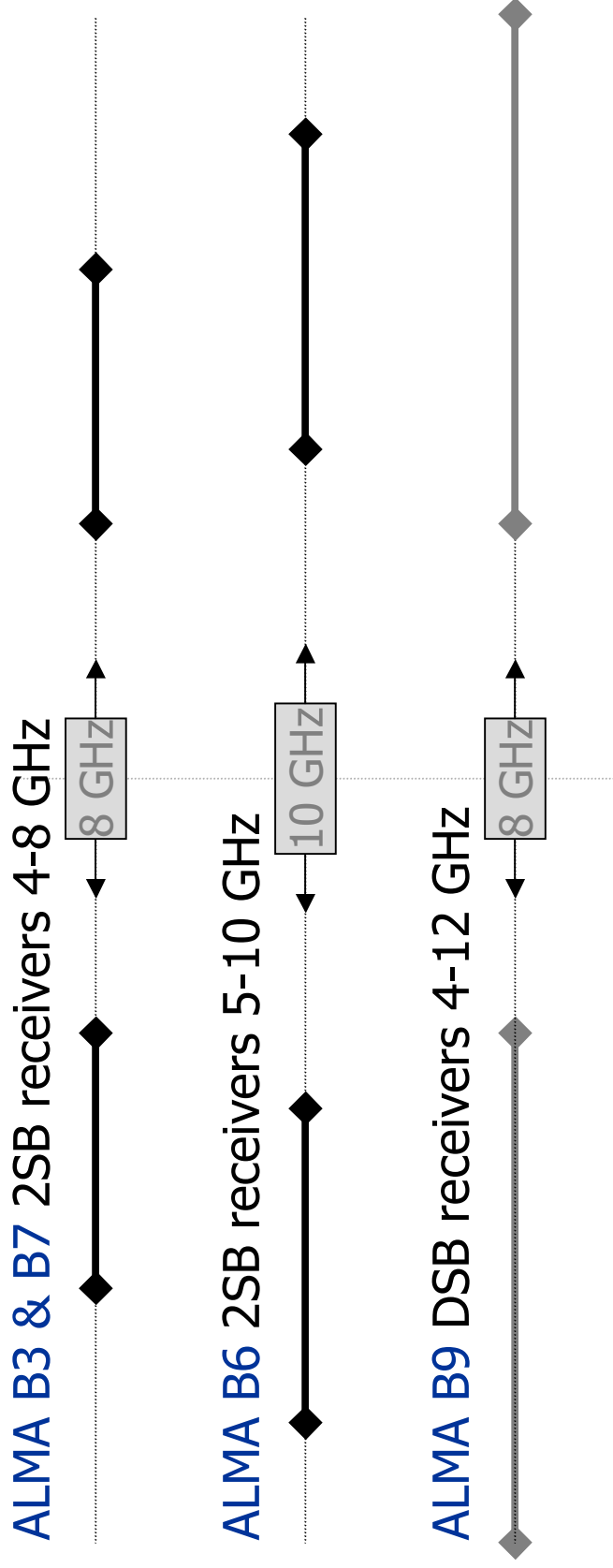


Receivers have bands 8 or 10 GHz bandwidth

x 2 polarizations



ALMA receivers

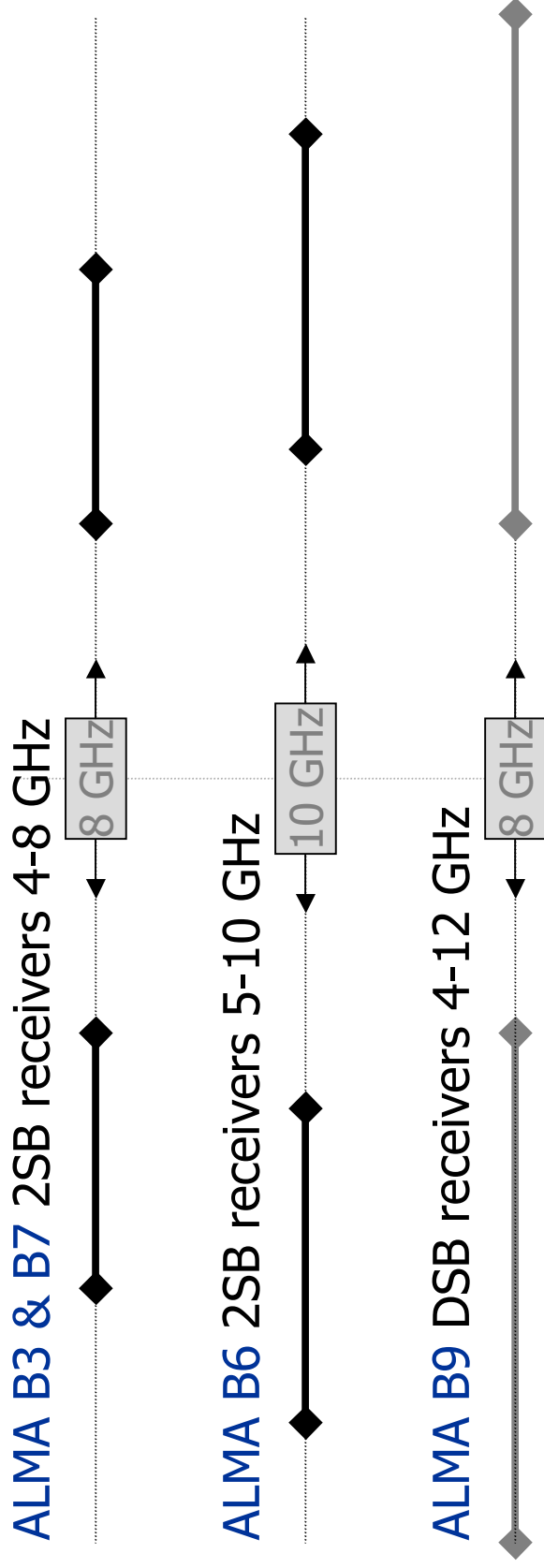


B6 receivers have 10 GHz bandwidth – but only 8 GHz can be processed by the IF and correlator system

→ useful to allow simultaneous observations of multiple lines, esp. ^{12}CO & ^{13}CO (2-1)



ALMA receivers



B9 are double sideband receivers – in Cycle 0, the two bands cannot be separated (Walsh)

... but one of the sideband can be suppressed (LO offsetting)



ALMA correlator

- ALMA correlator = **4 basebands**
- Each baseband processes
 - All antennas
 - 2 polarizations
 - **2 GHz input**

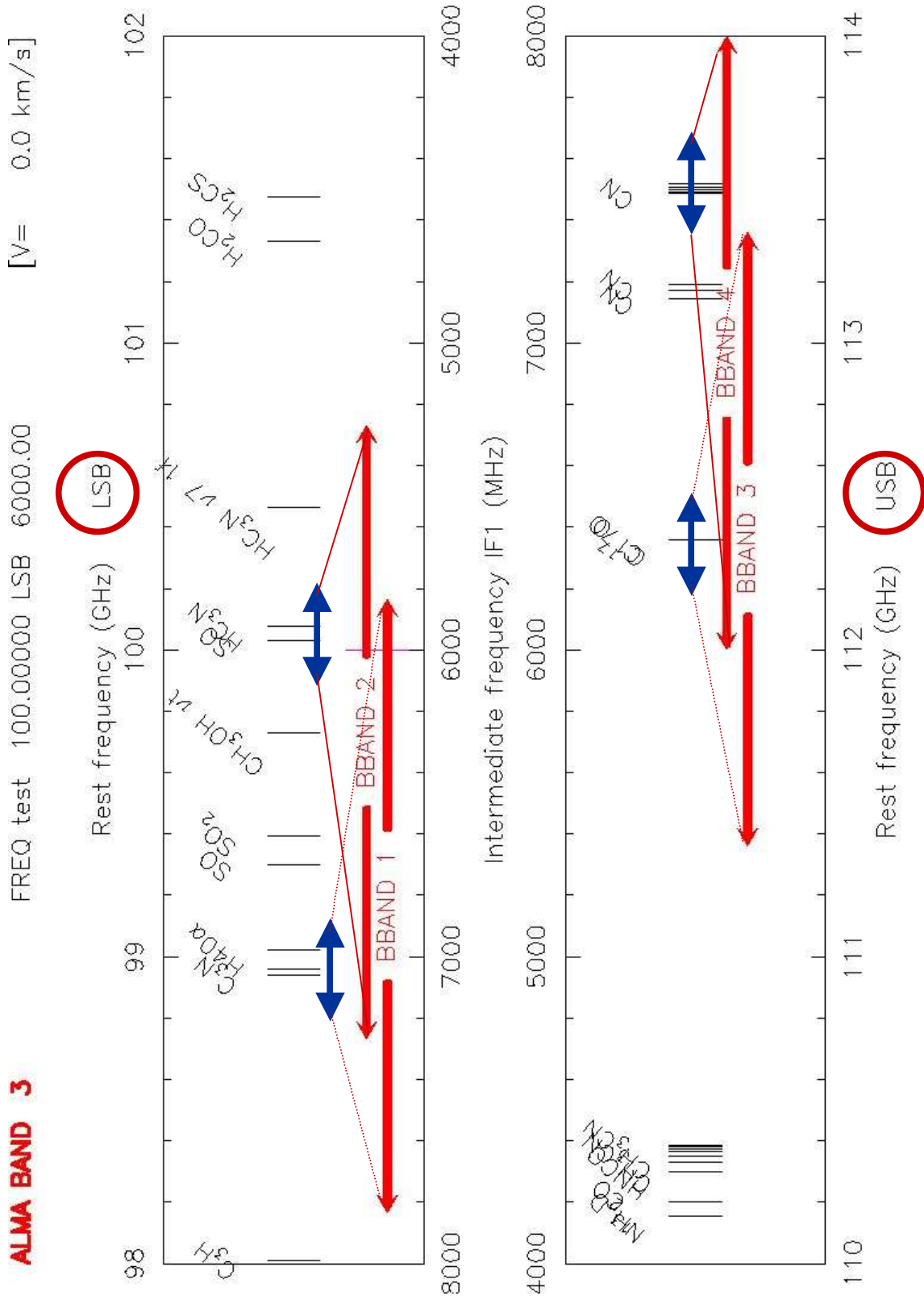
Can process
incoming 8 GHz
x 2 polarizations

- Each baseband can be centered anywhere* in the incoming 8 GHz
- All four basebands can be setup independently : gain on resolution / loss on bandwidth

(* Minor limitations because of LOs finite step)

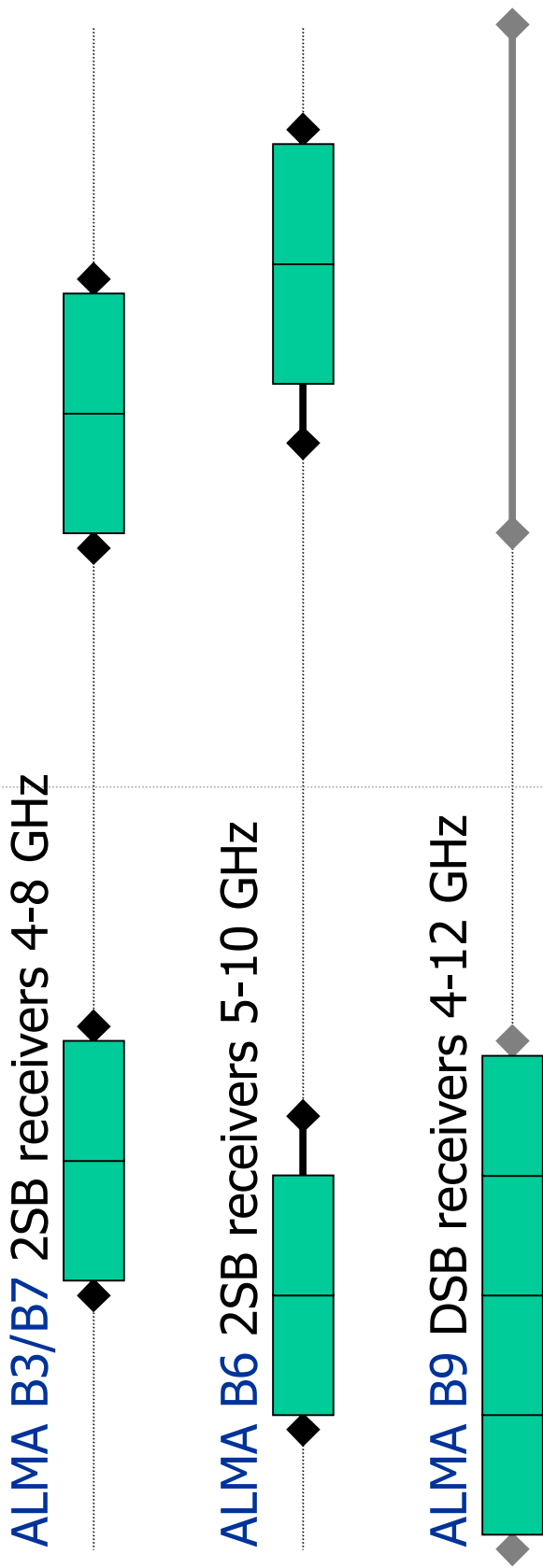
Higher spectral
resolution but
not on the full
8 GHz : spectral
windows

Four overlapping basebands giving four distinct spectral windows





Example



Four basebands covering 8 GHz

Note B6 larger bandwidth (2x5 GHz)

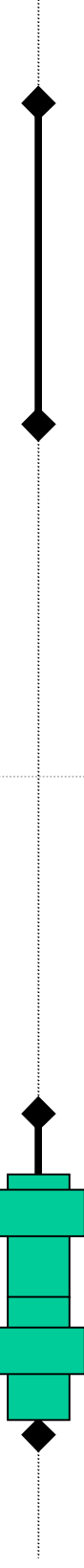


Example

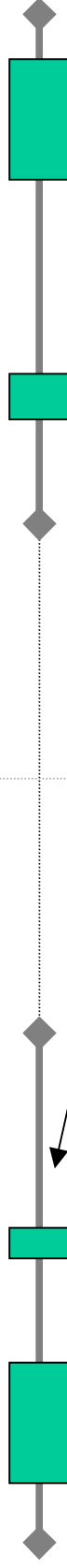
ALMA B3/B7 2SB receivers 4-8 GHz



ALMA B6 2SB receivers 5-10 GHz



ALMA B9 DSB receivers 4-12 GHz



Four basebands with different width/resolution

B9: choice of LSB/USB (sideband suppression) is done for each baseband

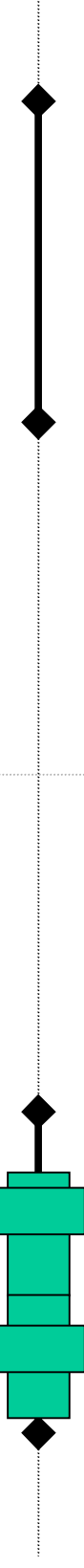


Example

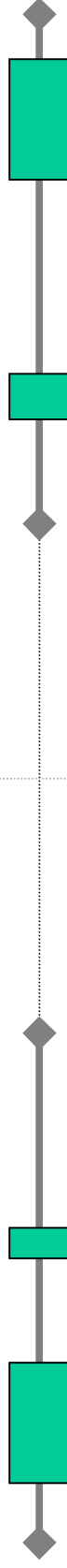
ALMA B3/B7 2SB receivers 4-8 GHz



ALMA B6 2SB receivers 5-10 GHz



ALMA B9 DSB receivers 4-12 GHz



Limitations

- 2 GHz-wide basebands must lie within the receiver IF
- Spectral window must lie within baseband
- Center of spectral window must be at >50 MHz from baseband edge

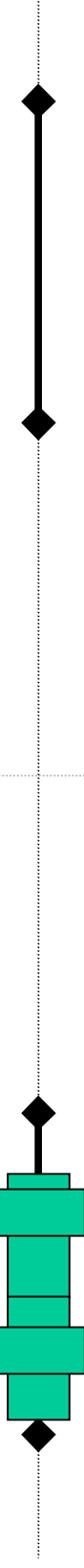


Example

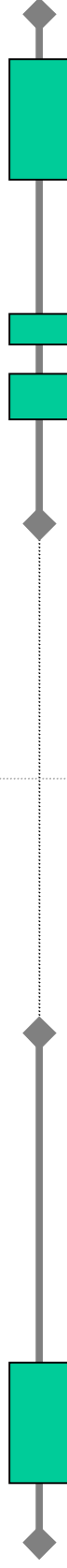
ALMA B3/B7 2SB receivers 4-8 GHz



ALMA B6 2SB receivers 5-10 GHz



ALMA B9 DSB receivers 4-12 GHz



Basebands in LSB and/or USB?

- **2SB receivers = B3/B6/B7 → only 4+0 or 2+2**
- **DSB receivers = B9 → 3+1 also possible**



Basebands modes

1 polarization output (H or V)

- **2 GHz** **8192 channels x 1 Pol** x 244 kHz spacing
 - **1 GHz** 8192 channels x 1 Pol x 122 kHz spacing
 - **500 MHz** 8192 channels x 1 Pol x 61 kHz spacing
 - **250 MHz** 8192 channels x 1 Pol x 30 kHz spacing
 - **125 MHz** 8192 channels x 1 Pol x 15 kHz spacing
 - **64 MHz** 8192 channels x 1 Pol x 7.6 kHz spacing
-
- **2 GHz** 256 ch. x 1 Pol x 7.8 MHz spacing

FDM

TDM



Basebands modes

1 polarization output (H or V)

- 2 GHz 8192 channels x 1 Pol x 244 kHz spacing
- 1 x 122 kHz spacing
- 50 x 61 kHz spacing
- 25 x 30 kHz spacing
- 12.5 x 15 kHz spacing
- 64 MHz 8192 channels x 1 Pol x 7.6 kHz spacing

- 2 GHz 256 ch. x 1 Pol x 7.8 MHz spacing

The spectral resolution is twice the channel spacing (Hanning smoothing – default)



Basebands modes

2 polarization outputs (H and V)

- **2 GHz** **4096 channels x 2 Pol** = 488 kHz spacing
- **1 GHz** 4096 channels x 2 Pol = 244 kHz spacing
- **500 MHz** 4096 channels x 2 Pol = 122 kHz spacing
- **250 MHz** 4096 channels x 2 Pol = 61 kHz spacing
- **125 MHz** 4096 channels x 2 Pol = 30 kHz spacing
- **64 MHz** 4096 channels x 2 Pol = 15 kHz spacing

- **2 GHz** 128 ch. x 2 Pol = 15.6 MHz spacing

Number of channels = 8192/Npolar



Basebands modes

- 4 polarization outputs (HH, VV, HV and VH)
 - 2 GHz x 2048 channels x 30 kHz spacing
 - 1 GHz x 2048 channels x 60 kHz spacing
 - 500 MHz x 2048 channels x 120 kHz spacing
 - 250 MHz x 2048 channels x 244 kHz spacing
 - 125 MHz x 2048 channels x 488 kHz spacing
 - 62.5 MHz x 2048 channels x 976 kHz spacing
 - 31.25 MHz x 2048 channels x 1952 kHz spacing
- 64 ch. x 4 Pol = 31 MHz spacing

Number of channels = 8192/Npolar



Real bandwidths

FDM modes

- Bandwidth edges are not used to avoid aliasing & edge effects
- **Effective bandwidth is 15/16th of the nominal width**
- **Real total number of channels is 7680 instead of 8192**
(remaining channels are dropped within the correlator and not written in the data)

TDM modes

- Filter in the IF chain limits the bandwidth to 1875 MHz
- FDM larger mode is already $2000 \times 15/16 = 1875$ MHz
- TDM: limits usable **bandwidth of continuum mode to 1875 MHz**
(instead of 2000 MHz)



Real bandwidths

TDM modes

- Filter in the IF chain limits the bandwidth to 1875 MHz
- FDM larger mode is already $2000 \times 15/16 = 1875$ MHz
- TDM: limits usable **bandwidth of continuum mode to 1875 MHz** (instead of 2000 MHz)

Consequence: ALMA Continuum sensitivity

- Maximal bandwidth for continuum is $4 \times 1.875 = 7.5$ GHz (per polarization)
- This is taken into account in the OT Cycle 0 **update 1** (released early June 2011)



Correlator Cycle 0

Proposers guide p.36

1 polarization output (H or V) – actual channels numbers

- 1875MHz 7680 channels x 1 Pol = 244 kHz resol.
- 938 MHz 7680 channels x 1 Pol = 122 kHz resol.
- 469 MHz 7680 channels x 1 Pol = 61 kHz resol.
- 234 MHz 7680 channels x 1 Pol = 30 kHz resol.
- 117 MHz 7680 channels x 1 Pol = 15 kHz resol.
- 58.6 MHz 7680 channels x 1 Pol = 7.6 kHz resol.

FDM

- Continuum mode 256 ch. x 1 Pol = 7.8 MHz resol.

TDM



Correlator Cycle 0

Proposers guide p.36

2 polarization outputs (H and V) – actual channels numbers

- **1875MHz** 3840 channels x 2 Pol = 488 kHz resol.
- **938 MHz** 3840 channels x 2 Pol = 244 kHz resol.
- **469 MHz** 3840 channels x 2 Pol = 122 kHz resol.
- **234 MHz** 3840 channels x 2 Pol = 61 kHz resol.
- **117 MHz** 3840 channels x 2 Pol = 30 kHz resol.
- **58.6 MHz** 3840 channels x 2 Pol = 15 kHz resol.

FDM

- Continuum mode 128 ch. x 2 Pol = 15.6 MHz resol.

TDM



Correlator Cycle 0

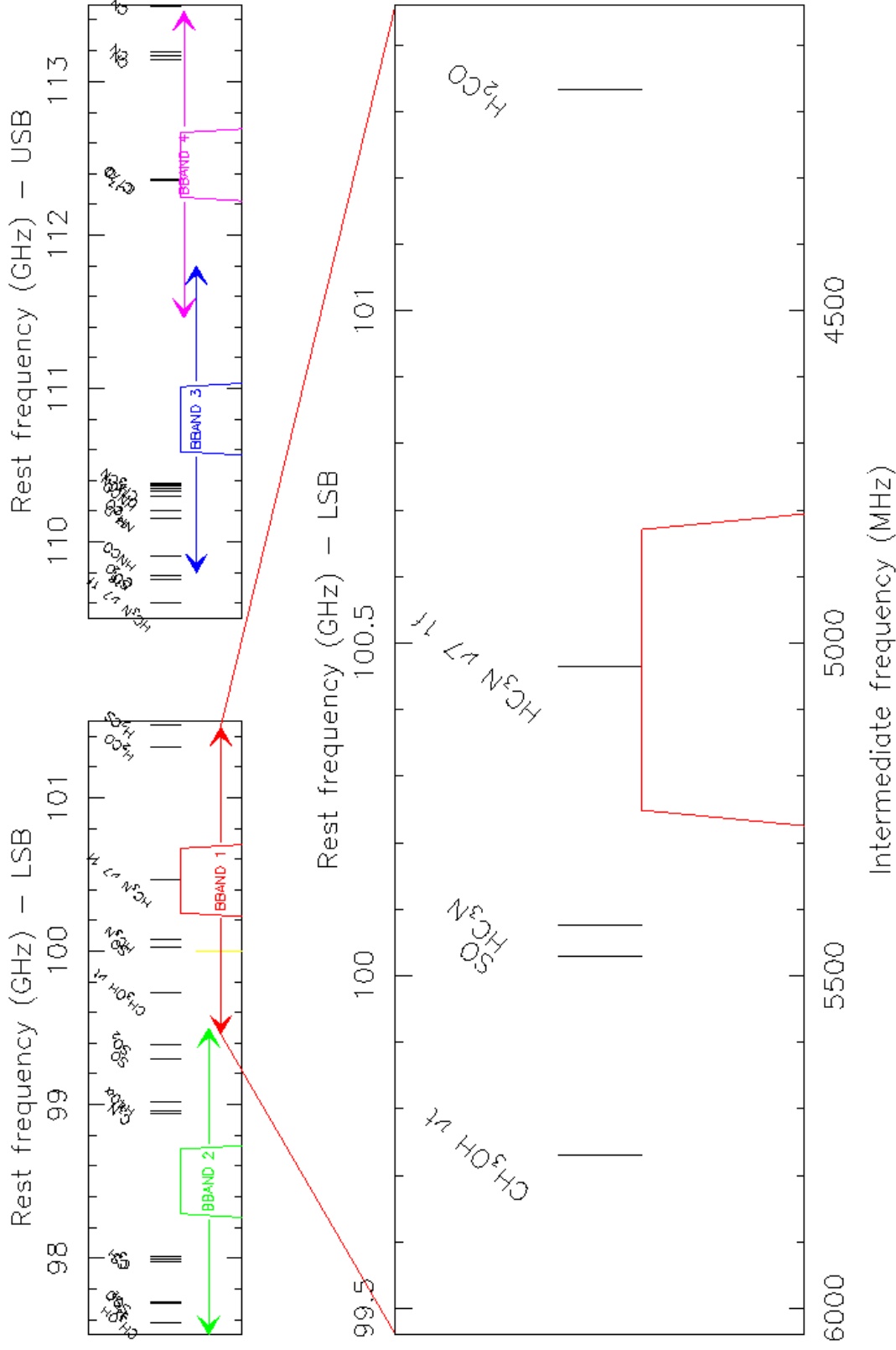
All four spectral windows must share the same mode

- Same **resolution / bandwidth** for the four spectral windows
 - Continuum & line? Must compromise!
- Same number of **polarization** products
 - Not a big issue
- **Same position within the baseband**
 - Making multi-line setups may be tricky if lines are near the edge of the receiver bandwidth

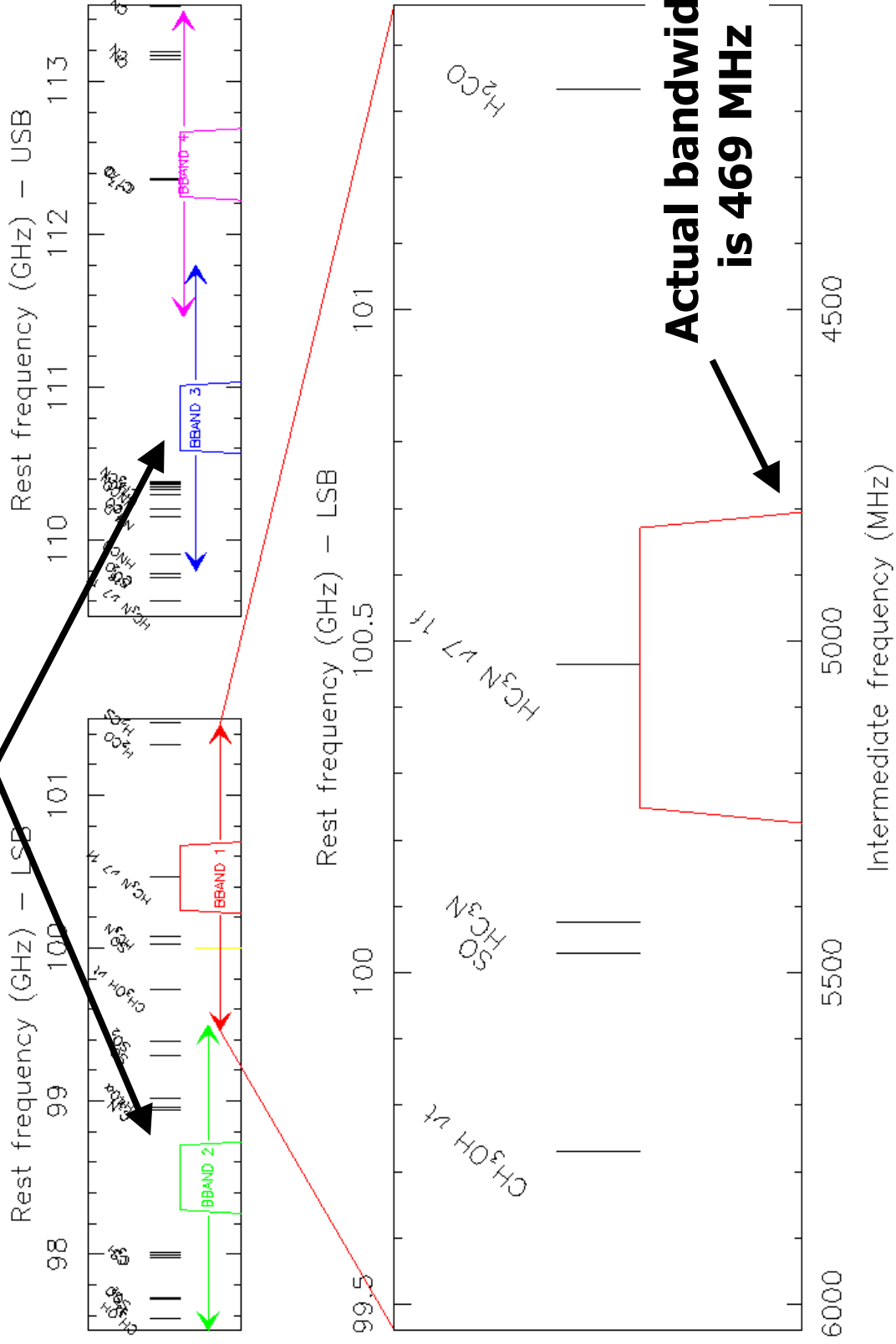
Technical Handbook
table 8 p. 103,
rules 6 & 7

ALMA BAND 3 FREQ test 100.00000 LSB 5500.00 [V= 0.0 km/s]

BASEBAND 1 is centered at IF1 = 5040.00 MHz (LSB) RF = 100.46000 GHz



Spectral windows basebands 2 & 3 are for continuum



Spectral window chosen for baseband 1 → the four spectral windows will be 500 MHz/1 Polar/centered in baseband



Limitation Cycle 0

The four spectral windows share the **same position within the baseband** – is that a serious limitation?

Simple solution: center spectral window in basebands, and center the basebands on the right position (line of interest)

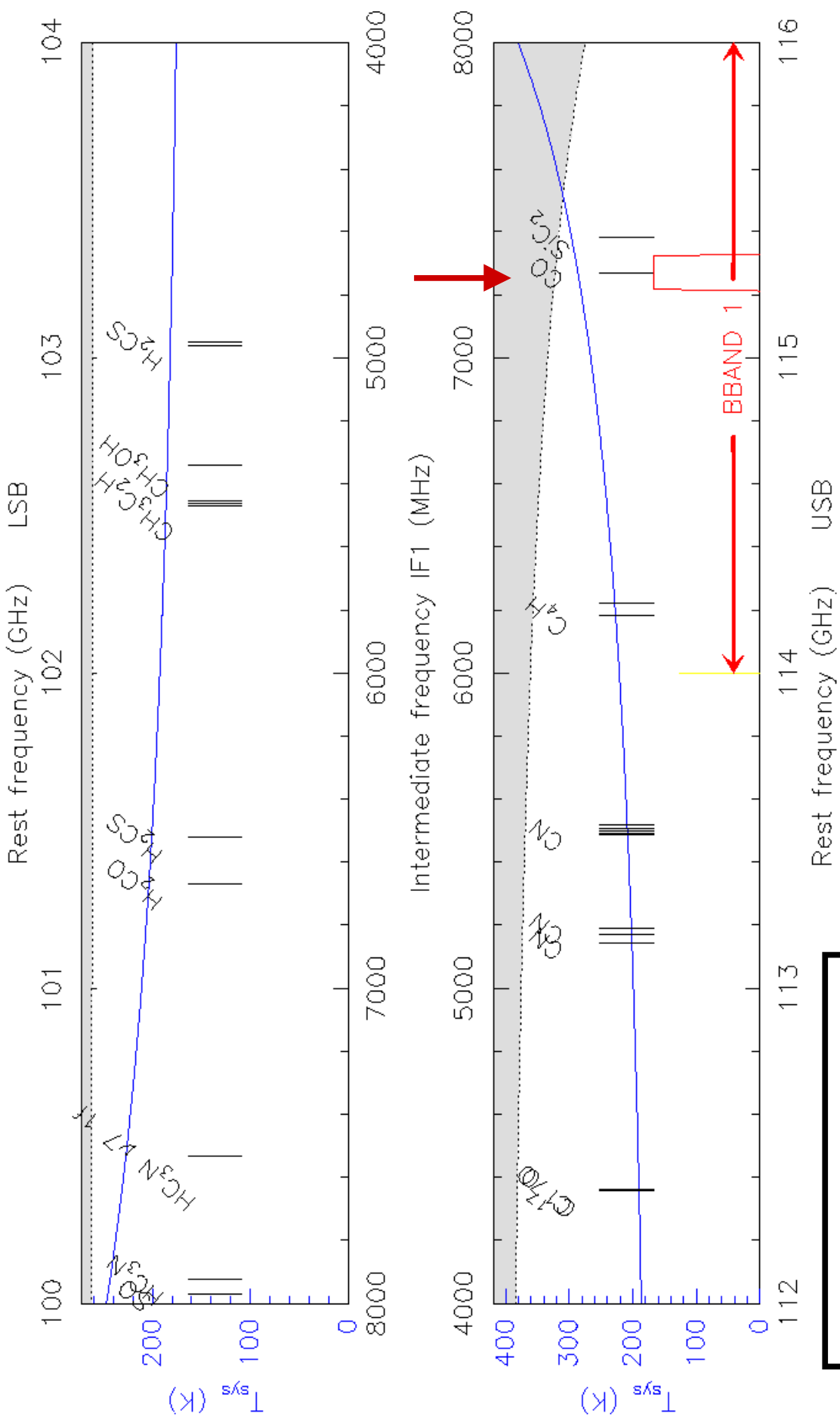
Problem if line to be observed is **near the edge of the IF bandpass** of the receiver

Two important cases: CO(1-0) at 115.3 GHz (B3)
 $^{12}\text{CO}(2-1) + ^{13}\text{CO}(2-1)$ at 230/220 GHz (B6)

ALMA BAND 3

FREQ co(1-0) 114.000000 USB 6000.00

[V= 0.0 km/s]

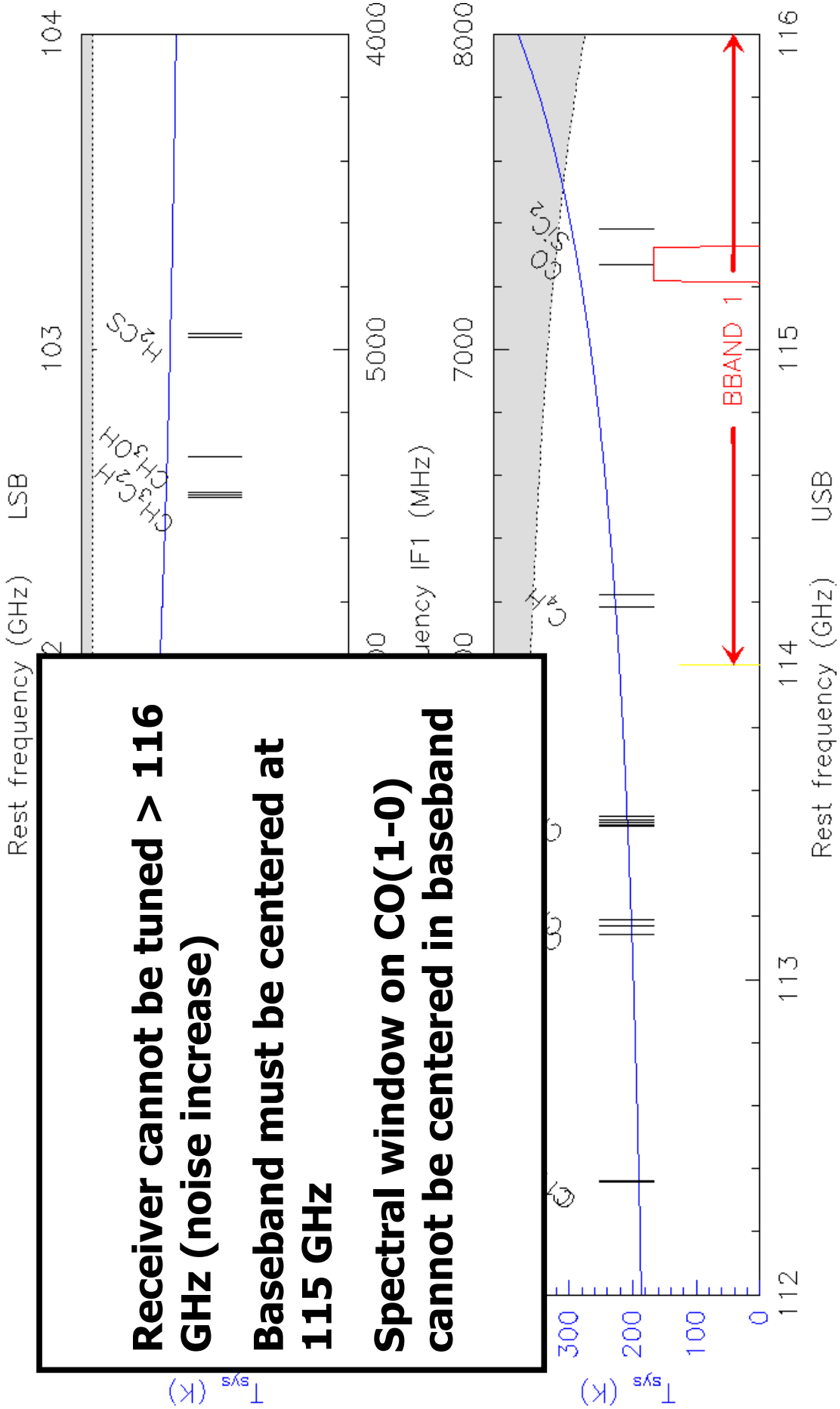


^{12}CO (1-0) with B3

ALMA BAND 3

FREQ co(1-0) 114.00000 USB 6000.00

[V= 0.0 km/s]

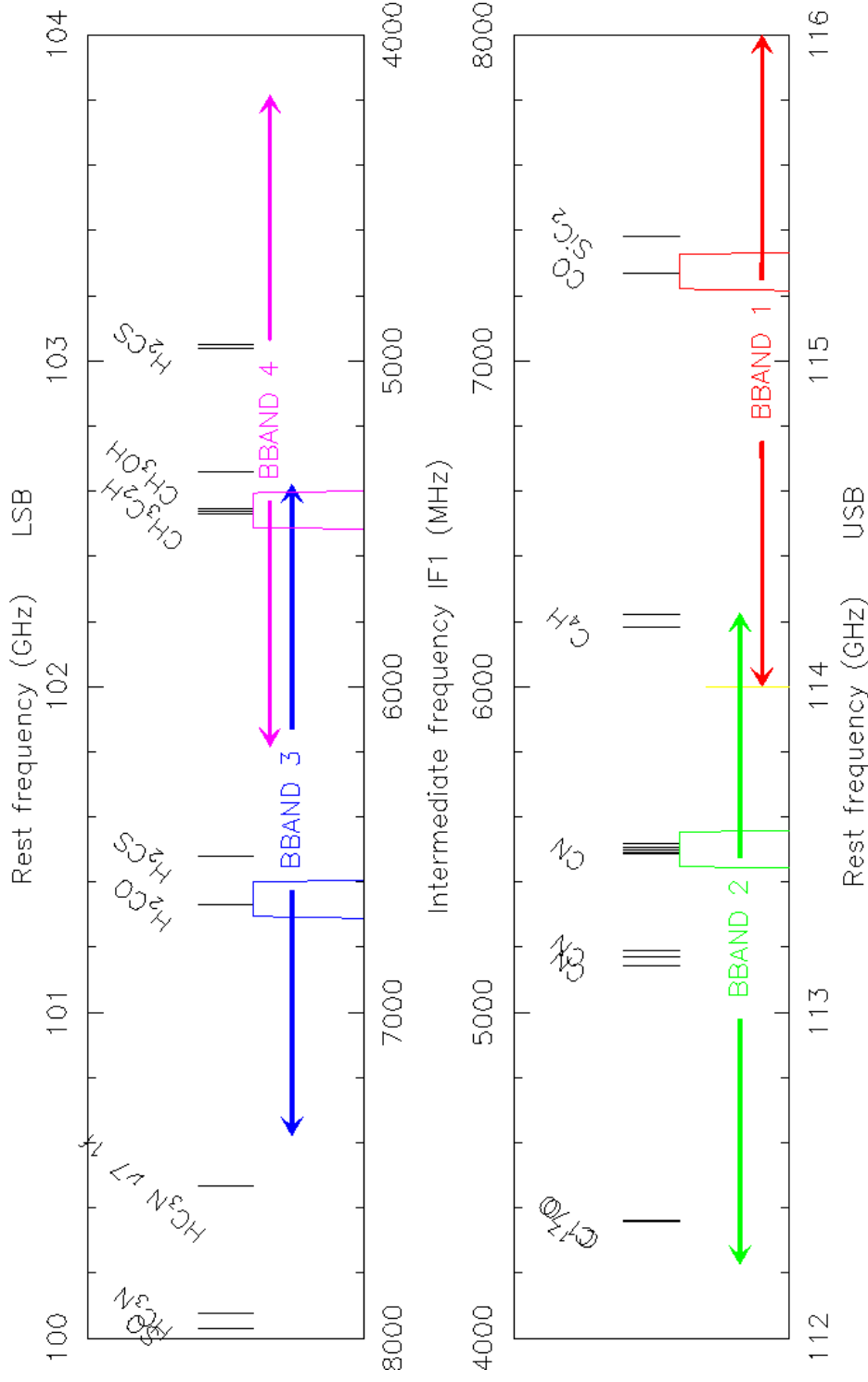


Receiver cannot be tuned > 116 GHz (noise increase)
Baseband must be centered at 115 GHz
Spectral window on CO(1-0) cannot be centered in baseband

ALMA BAND 3

FREQ co(1-0) 114.00000 USB 6000.00

[V= 0.0 km/s]

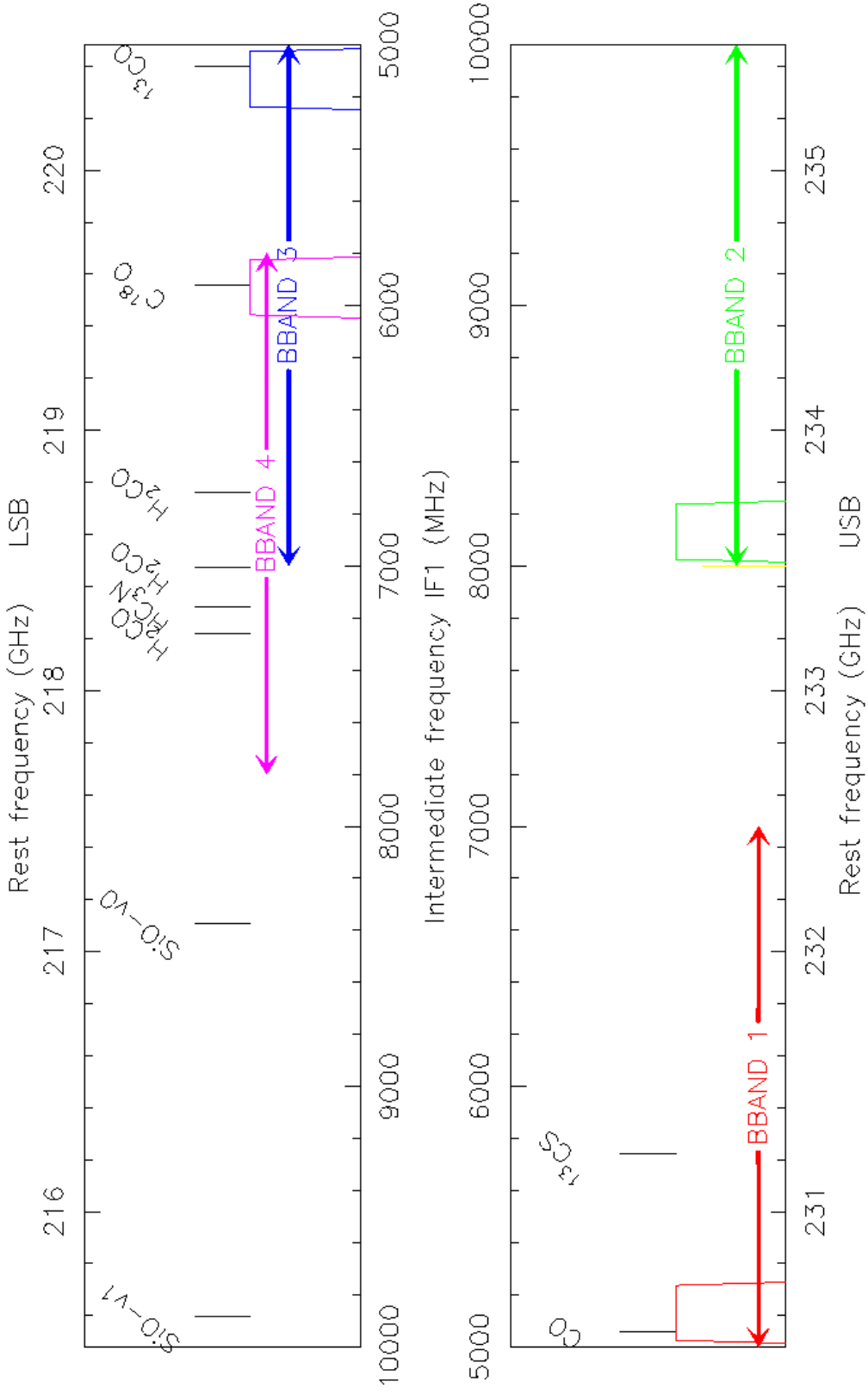


Position of spectral windows in other basebands is fixed
Position of basebands can be moved
But not all lines can be reached

CAUTION: this is not taken into account by the OT

ALMA BAND 6

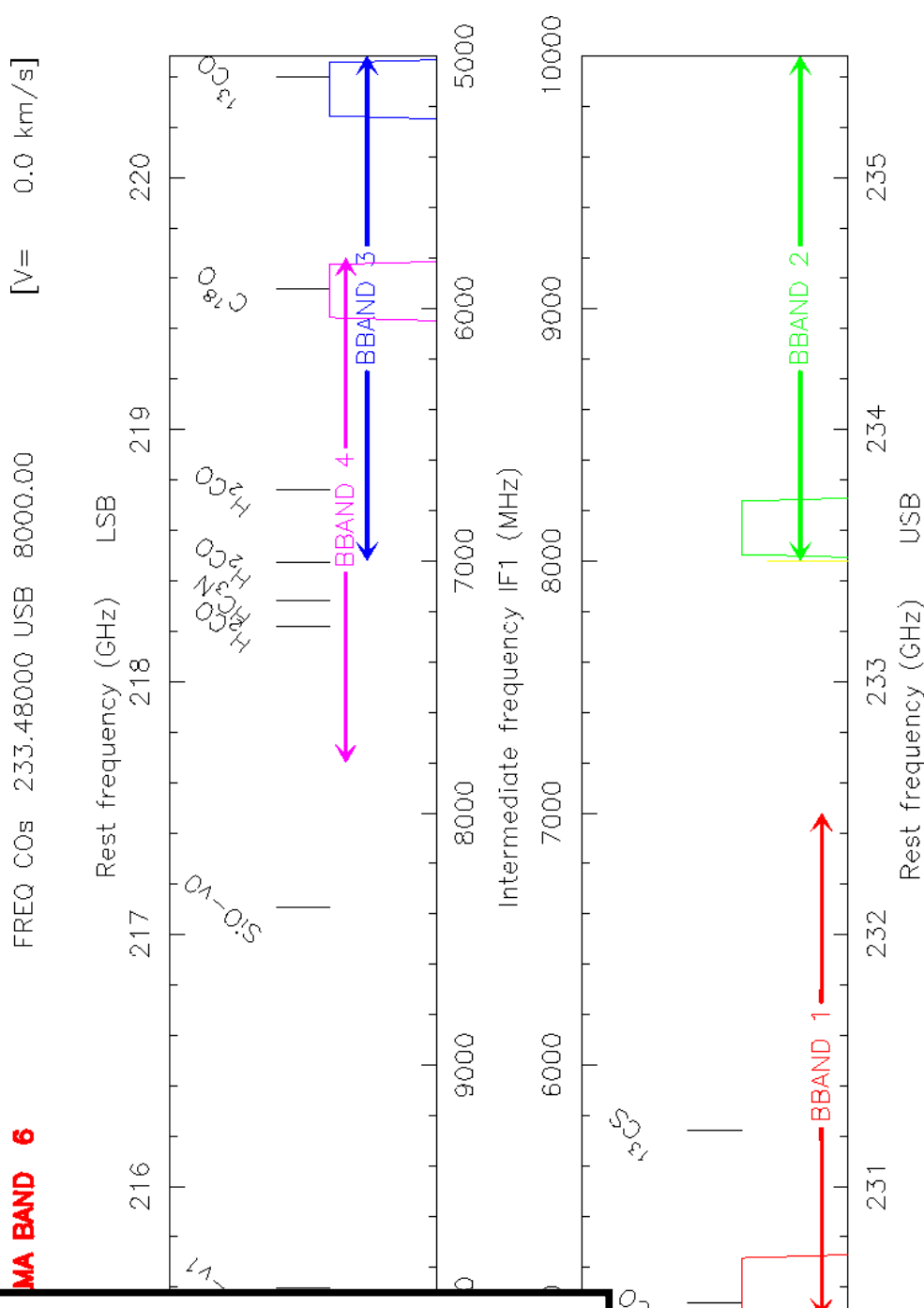
FREQ COs 233.48000 USB 8000.00 [V= 0.0 km/s]



Simultaneous observations of ^{12}CO , ^{13}CO and C^{18}O (2-1) with B6

- ^{12}CO and ^{13}CO are near the IF edge
- Spectral windows must be strongly offset
- Cannot use narrower modes – center would be <50 MHz from edge (forbidden)
- Can observe C^{18}O
- Last baseband has to be in USB (2+2 rule)

CAUTION: this is not taken into account by the OT



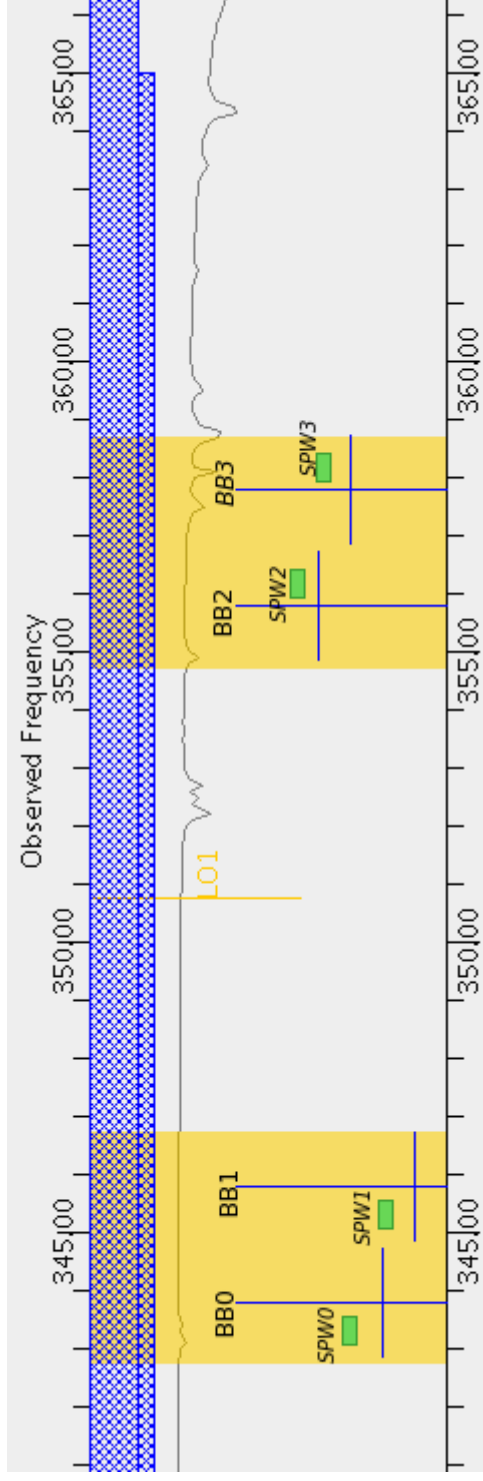


Testing setups with ASTRO

- **GILDAS/ASTRO** can produce Bure-like plots for the ALMA correlator
- The various restrictions are taken into account
- Commands are:
 - **FREQUENCY** to define receiver tuning
 - **BASEBAND** to define position of basebands
 - **PLOT BASE i** to plot baseband *i*
 - **SPWINDOW** to define spectral window in that baseband
 - **PLOT FREQ** or **PLOT BASE i** to produce updated plots (full view or baseband view)



Spectral setup in the OT



Yellow areas = LSB/USB (here Band 7)

BB0 to BB3 = 4 basebands, each 2 GHz wide

SPW0 to SPW3 = 4 spectral windows, each within a baseband

Note similar position of SPWs in each baseband



References

- ALMA Cycle 0 Proposer's Guide
Doc. 0.2 , V1.1, May 2011
- ALMA Cycle 0 Technical Handbook
Doc. 0.3, V1.0, May 2011
- ALMA memo 556
- Thanks to A.Baudry (Bordeaux), R.Hills (JAO),
R.Lucas (JAO), V.Pietu (IRAM), P.Salome (Paris)