

IRAM Annual Report 2020

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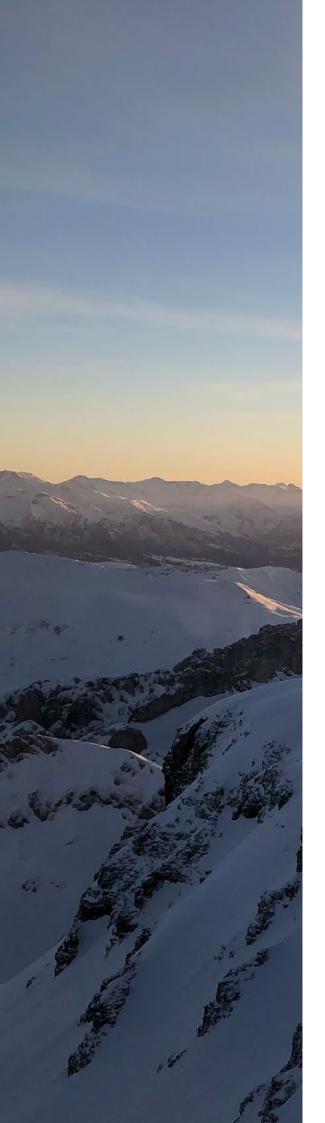
Director of publication Karl-Friedrich Schuster Edited by Frédéric Gueth, Sonja Moreau

With contributions from:

Sébastien Blanchet, Antonio Córdoba, Eduard Driessen, Bertrand Gautier, Olivier Gentaz, Romain Messmer, Santiago Navarro, Roberto Neri, Juan Peñalver, Jérôme Pety, Francesco Pierfederici, Christophe Risacher, Miguel Sánchez Portal, Karl-Friedrich Schuster

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Introduction

The reader of this annual report may easily imagine that the year 2020 was for IRAM totally dominated by the struggle to adapt to the constraints of the Covid-19 pandemic. And indeed, this took a very important part of our energy on all levels: personal, institutional and also societal. Safety for the staff and their families was and is the main concern and this had to be combined with a continuous questioning and re-definition of rules and ways of doing.

It is all the more absolutely remarkable what was achieved over the year. Let me first thank here the entire IRAM staff who knew to stick together, to adapt and to move forward while keeping the spirits high despite the very numerous difficulties in professional and everyday personal live. We have been witnesses of the old truth that people show their great qualities in important crises. Only in the aftermath of the current situation it will become clear how much depended on keeping the everyday business going. On the one side, we wanted to avoid too disruptive situations for the staff and on the other side, it was important to continue best possible service to the scientific user community under the given circumstances.

In fact, as soon as the rapidly degrading situation requested a first general lockdown in March 2020 in France and Spain, a huge effort was made between the safety coordinators as well as the group leaders and the station managers to insure safety at the IRAM observatories and the headquarter including creation of information channels, reorganization of work flows and installation of a vast number of safety measures.

At the observatories, this immense work and the exceptional motivation of the staff allowed to overcome the very difficult situation in the first weeks of the pandemic including a lot of problems generated by the fact that the related ski stations for both observatories went into full stop with no service provided to IRAM. Despite these difficulties to access the observatories, the IRAM 30-meter telescope only had to go into survival mode for four weeks before switching successfully to a totally new operation mode with exclusive remote user support from operators and astronomers on duty. NOEMA even could seamlessly continue science operation and only antenna assembly work was interrupted for eight weeks. With all these challenges, it was nevertheless possible to finish assembly and

commission of the 11th NOEMA antenna in September 2020. At the same time, the French partner CNRS could confirm his commitment to fully finance Antenna 12 with completion of this final NOEMA antenna foreseen for fall 2021.

Likewise, an incredible number of projects made very important progress in the laboratories. Among many other projects and developments, one might mention the first dual-band receiver that was installed at NOEMA showing very promising performance. It is further notable, that the 2nd water vapour radiometer generation went from prototype status into production and that the NOEMA VLBI phasing has seen first very successful tests.

Three outstanding events added onto this in the second half of the year 2020 which I will mention here in temporal order. Beginning of October, our executive council member for more than 30 years, Reinhard Genzel was awarded the Nobel prize for Physics. Throughout the decades, Reinhard Genzel has been a key person to support the progress and evolution of IRAM and we all were extremely happy and proud about this reward. Then in beginning of November came the news that the Spanish science ministry awarded very important structural funds to support the refurbishment and ambitious upgrade of our iconic workhorse, the IRAM 30-meter telescope. And finally, end of November 2020, the IRAM partners signed the extension of the collaboration agreement until 2034. Needless to say, that this is received by the IRAM staff as a confirmation of the high esteem and trust from the partner organisations into the persons at work and in IRAM as a whole.

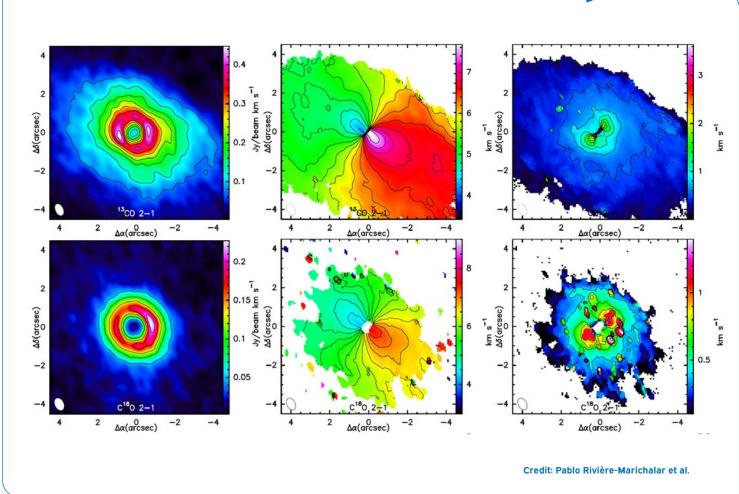
This leaves us with a bittersweet look onto 2020 where so much suffering and problems on all levels are side by side with a large number of achievements and positive events in our institute. My hope is that the health urgency can soon be overcome and we can find ways to duly celebrate together at least part of these achievements.

With best regards

Karl-Friedrich Schuster
Director



Highlights of research with the IRAM telescopes



A protostellar binary system fed by a 10500 AU gas streamer

One of the key challenges in the field of star formation is the formation of binary stars. Among the possible routes for the formation of close-companion stars is the fragmentation of a gravitationally unstable disk around a protostar early in the star formation process. As the disk grows in mass, it eventually becomes gravitationally unstable and fragments, forming one or more new protostars orbiting each other. This scenario is supported by models where the small-scale disk asymmetries observed around young protostars and possibly resulting from an eccentric fragmentation of the large-scale disk environment, are at the origin of the formation of a close binary.

In an attempt to explore the environment of a Class 0 binary system, a team of scientists led by Jaime Pineda (MPE/Garching) used NOEMA to observe, at disk-forming scales, the dynamical properties of the gas in the proto-stellar system Per-emb-2 (IRAS 03292+3039). The high-resolution observations revealed the presence of a large reservoir of fresh material in almost free-fall that narrows down to an elongated gas stream of 10500 AU under the gravitational pull of the innermost part of the core. The stream carries large amounts of gas (0.1 M_{\odot}) with chemicals such as HC₃N, C₂S, ¹³CS, which are channeled at a rate of 7 10⁻⁷ M_{\odot} /yr onto the disk surrounding the binary star system. According to the authors, both the locations and the speed of the gas are well matched by a theoretical model of a stream of material free-falling from large to small scales, and are

thereby confirming that the streamer's dynamics are controlled by the densest central region of the system. From the streamer density, the depletion timescale is estimated to \sim 20000 yr, which is comparable to the free-fall time scale but much shorter than the canonical \sim 100000 yr duration of the Class 0 phase.

The results of this study provide impressive evidence that the large-scale environment around forming stars has an important influence on small-scale disk formation and evolution. Moreover, they show that the non-axisymmetric enrichment of fresh material injected into the disk is a possible trigger for the formation of binary and possibly multiple star systems. Finally, they provide new support for studies claiming that streamer-like features seen on disk scales are the result of strong envelope accretion.

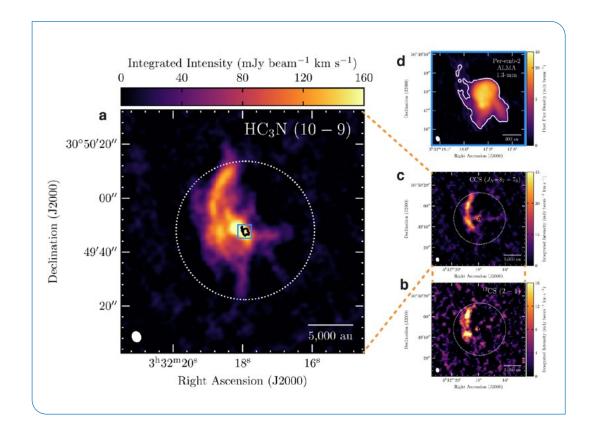


Image of the 'streamer' feeding chemically fresh material from a distance of about 10500 AU to the proto-star at the centre of the image. The three images use different molecules as tracers, indicated in the top right corner, and all show the streamer in action. The colour coding is according to the integrated intensity of the signal. The 1.3mm continuum panel at the top right is from ALMA. Work by Pineda et al. 2020, Nature, 4, 1158

Cold gas accretion mode of star formation in AzTEC2?

Several lines of evidence suggest that the global production of stars in galaxies is mainly regulated by the steady accretion of gas from the intergalactic medium. This process of galaxy evolution is believed to be responsible for the widespread star formation in galactic disks over giga-year timescales. It is known as the "cold gas accretion mode" of star formation, which differs from the rather short-lived but more intensive production of stars often triggered by major/minor galaxy mergers. Whereas both regimes of star formation have been widely explored out to intermediate redshifts $(z\sim2)$, the relative role of the cold gas accretion and merger mode in driving the intense production of stars in galaxies at higher redshifts $(z\sim3)$ remains an open issue.

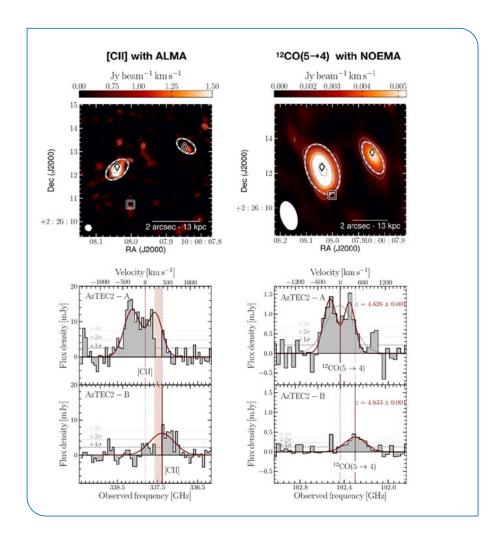
To investigate the role of cold gas accretion in star formation at higher redshift, a team of astronomers led by Jiménez-Andrade (NRAO/Charlottesville) studied AzTEC2, the second brightest SMG in the COSMOS field, which consists of a massive, star-forming disk and a smaller companion galaxy. By combining observations from NOEMA and ALMA, the authors assessed the redshift and confirmed that AzTEC2 is binary galaxy system at a redshift of 4.63, and thereby ruled out previous associations with a galaxy at $z\sim1$.

By using the 12 CO (5–4) line emission and adopting typical SMG-like gas excitation conditions, and a gas mass-to-luminosity conversion factor α_{co} = 2.5, Jiménez-Andrade and collaborators estimated the molecular gas mass to 3.2 10 11 M $_{\odot}$ for AzTEC2-A, and to 0.8 10 11 M $_{\odot}$ for AzTEC2-B. With infrared-derived star formation

Upper panels: velocity-integrated intensity maps of [CII]158µm and ¹²CO (5-4) detected towards AzTEC2-A and AzTEC2-B with ALMA and NOEMA, respectively. Lower panels: [CII]158µm and ¹²CO (5-4) spectra of AzTEC2-A and AzTEC2-B. The red solid line represents the model. The velocities displayed in the spectra are relative to the central frequency of the ¹²CO (5-4) line emission in AzTEC2-A.

Work by Jiménez-Andrade et al.

2020, ApJ, 890, 171

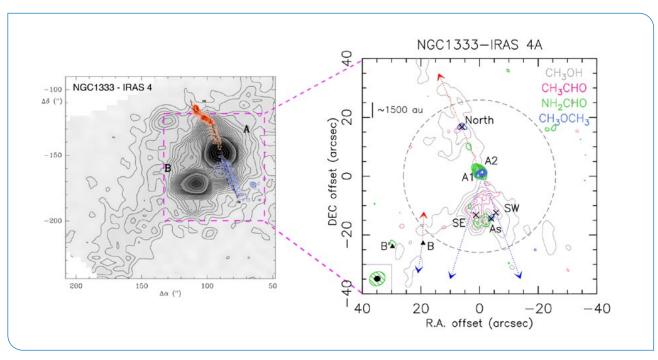


rates of 1920 M_{\odot} /yr and 710 M_{\odot} /yr for AzTEC2-A and AzTEC2-B respectively, their current gas reservoirs are expected to be depleted within 30–200 Myr. Their results indicate that AzTEC2-A hosts a massive, rotation-dominated disk with a de-projected rotation velocity of 660 km s⁻¹ where star formation occurs at intense levels. According to the authors, this indicates that even disk galaxies that harbour vast gas reservoirs could sustain intense star formation activity that resembles that of merger-driven SMGs.

This study supports the emerging consensus whereby the population of single-dish selected SMGs is rather heterogeneous, including both interacting systems and galaxies that form stars through a smoother mode of star formation sustained by cold gas accretion. Further systematic studies are already planned to verify this scenario.

Complex organic molecules in IRAS 4A

The presence of complex organic molecules (iCOMs) around solar-type protostars has raised the question of their role in the chemical origins of biotic molecules and thus in the emergence of life on Earth. While their possible importance for the occurrence of life is undisputed today, iCOMs pose a challenge to astrochemistry, as their synthesis is far from obvious. Nowadays, two main paradigms are proposed, according to which iCOMs are synthesized either on the surfaces of dust grains or in the gas phase via sputtering. Although both pathways require the formation of simple hydrogenated molecules on dust grains during the pre-stellar phase, both pathways can lead to the synthesis of iCOMs. However, which of the two is the more efficient and under what conditions iCOMs form remain open questions.



Overlay (left) of the 1.25 mm continuum emission (grey contours, IRAM 30-meter telescope) and the SiO emission tracing the outflows (colored, VLA). Zoomed-in image (right) of the IRAS 4A system observed ith NOEMA. The dashed blue and red arrows indicate the directions of the blue- and red-shifted outflows from 4A1, 4A2 and 4B. The velocity integrated maps of the iCOMs listed in the top right corner are colored accordingly. Work by De Simone et al. 2020, A&A, 640, A75

One method that turns out to be efficient in disentangling the two formation routes is to compare observations toward low-mass outflow shocks with model predictions. Indeed, shocks that sweep through the outflows drive the chemical evolution of the outflows over time, allowing observed molecular abundances to be compared with model predictions and thereby constraining the formation pathways of iCOMs.

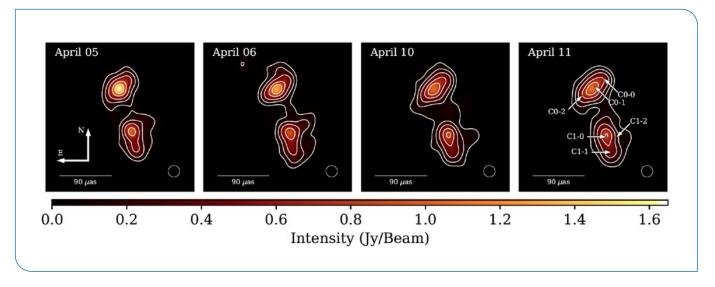
To shed new light on the iCOM chemistry in the outflows of NGC1333-IRAS 4A, a research group led by Marta De Simone (IPAG/Grenoble) reported the detection of four iCOMs with NOEMA: methanol ($\mathrm{CH_3OH}$), acetaldehyde ($\mathrm{CH_3CHO}$), formamide ($\mathrm{NH_2CHO}$), and dimethyl ether ($\mathrm{CH_3OCH_3}$). The researchers were able to observe significant differences in the iCOM abundances between the southeastern outflow driven by IRAS 4A1, and the north-southwestern outflow driven by the hot corino in IRAS 4A2. By using an astrochemical gasphase model (GRAINOBLE), the authors show that the different distribution of some iCOMs can be explained by gas-phase reaction, and in particular for acetaldehyde. Indeed, although laboratory experiments show that acetaldehyde can be formed on the surfaces of dust grains, the comparison of their models with the observed gas distribution provides strong evidence that the gas-phase formation route cannot be neglected for acetaldehyde, and that it depends very little on the details of the modeling.

The authors conclude that gas-phase reactions do indeed appear to be the dominant process in the IRAS 4A outflows, and that the chemical differentiation between the two outflows is presumably related to their different kinematical age. Further observations of outflows at different ages are planned to corroborate these results.

Event Horizon Telescope observations of the blazar 3C279

Relativistic jets in active galactic nuclei (AGN) are believed to originate from the vicinity of a supermassive black hole (SMBH) at the center of the galaxy. Understanding the detailed physical processes that lead to the formation, acceleration, collimation, and the subsequent propagation of a jet is one of the major quests in modern astrophysics. Extensive high-resolution AGN studies have been conducted on these topics over the past decades, but it is only through observations such as those allowed by the Event Horizon Telescope (EHT) Collaboration that a significant breakthrough seems possible.

To resolve the fine-scale structures of jets in the vicinity of SMBHs, a critical precondition to better understand both the accretion phase and the way jets evolve, Jae-Young Kim (MPIfR/Bonn) and collaborators made use of the EHT, and so of the IRAM 30-meter telescope, to observe the 230 GHz continuum emission in the innermost jet regions of 3C279, one of the brightest known blazars. The EHT observations, which were obtained at 20 μas resolution, show that the core region (C0 region) in 3C279 consists of three bright substructures (CO-0, CO-1, CO-2) oriented perpendicular to an elongated structure, and separated by 30-40 μ as, which corresponds to a projected spatial scale of ~3000 R_s for M_{BH} = 810⁸ M_{\odot} The EHT observations were supplemented by quasi-simultaneous VLBA (43 GHz) and GMVA (85 GHz) maps of the large-scale core and jet features. According to the authors the three substructures are associated with either a curved jet or a linear, nodular structure such as a site of large-scale magnetic reconnection or of plasma instabilities. The fact that the three core substructures show proper motions of ~1.4 μas/day and high flux variability suggests that they have something to do with how the jets meet the accretion disc surrounding the black hole. According to the authors, the rather low intrinsic brightness temperature (T \leq 10¹⁰ K) of the core and its morphological complexity also suggest that the core region of 3C279 becomes optically thin at short (mm) wavelengths, or that the innermost jet is dominated by magnetic energy if the synchrotron turn-over frequency were close to 230 GHz.



EHT images of 3C279 on April 5 to 11 generated by averaging the images obtained using three different imaging pipelines. The circular 20 μs restoring beam is shown in the bottom right corner of each panel. The IRAM 30-meter telescope is part of the GMVA and EHT networks.

Work by Kim et al. 2020, A&A, 640, A69

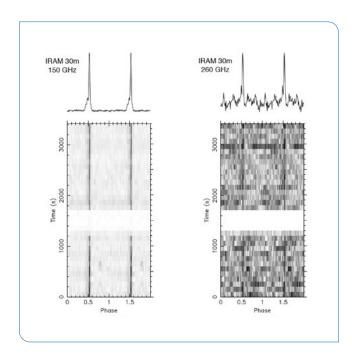
NIKA2 detects strong pulses in the magnetar XTE J1810-197

Pulsars are fast spinning, highly magnetized neutron stars capable of generating beamed radiation that appears as periodic pulses. While the physical mechanism of pulsar radio emission remains poorly understood, observational evidence suggests that it must be due to some coherent radiation mechanism. Unfortunately, neither the exact mechanism nor its origin is clear. Measuring pulsar emission across the electromagnetic spectrum is thus important for constraining theoretical models, and in this regard the poorly explored window between the radio and optical is of particular relevance.

To investigate whether pulsars can be detected in the mm-wavelength range with kinetic inductance detectors (KID), IRAM researcher Pablo Torne and collaborators, have managed to detect broadband pulsations from the X-ray pulsar XTE J1810–197 using the NIKA2 camera installed at the IRAM 30-meter telescope. XTE J1810-197 was inactive for about three years, but it recently showed pulsar emission again, which shows that the source is highly variable.

According to the researchers, this is the first time a pulsar was detected in the millimeter band using receivers based on KID technology, and the first time that synchrotron pulses were detected from XTE J1810–197 in the millimeter range.

Using the complex readout system, in which signals are sampled at a rate of 23.48 Hz, short radiation pulses were observed at 150 and 260 GHz, the two frequency bands of operation of NIKA2. The extremely high brightness of the single pulses, which was estimated to be $T_{\rm B} > 10^{17}$ at 150 GHz, is a strong indication that the emission region is radiating coherently. The authors of the study believe that although the pulses at 260 GHz could not be detected individually despite the high sensitivity of NIKA2, the emission mechanism is likely the same because of the high similarity between the continuum-equivalent flux density and the averaged pulse profile. According to them this result is another step forward in the quest to uncover the pulsar radio emission mechanism.



Average pulse profiles (top) and signal intensities (bottom) of the magnetar XTE J1810–197 observed with NIKA2 at 150 GHz (left) and 260 GHz (right). The pulsar orbital period is ~5.54 s. Integration time bins are 43 ms on the horizontal axis, 1.7 min on the vertical axis. The white gap corresponds to the time when the telescope was not observing the source.

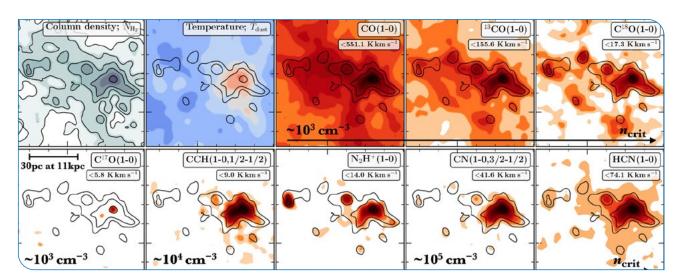
Work by Torne et al. 2020, A&A, 640, L2

A LEGO offspring – a chemical survey of the massive star forming region W49

The process of star formation from dense molecular clouds is still one of the major unsolved problems in astronomy. Many fundamental open questions are still hotly debated, such as understanding the relative balance of turbulence, magnetic field, and gravitational energy, how these properties can be observationally determined, how these properties may vary with the environment, and if they are important for the fraction of dense gas that is eventually converted into stars. Developing a better understanding of the emission properties of dense gas tracers across a broad range of environments is of critical importance to address these questions.

For this purpose, Ashley Barnes (AlfA/Bonn) and collaborators investigated the molecular environment of a 0.01 kpc² region centred on the star-forming region W49, which was observed with the IRAM 30-meter telescope in the frame of the "Line Emission as a Tool for Galaxy Observations" (LEGO) Large Program.

The researchers found that the spatial extent or brightness of the molecular line transitions are not well correlated with their critical densities, highlighting the fact that abundance and optical depth must be considered when estimating line emission characteristics. By exploring how the total emission and line emission vary as a function of the molecular hydrogen column density and dust temperature, they found that there is not a single region of this parameter space responsible for the brightest and most efficiently emitting gas for all species. It turns out for example, that the HCN (1-0) transition shows high emission efficiency at high column densities (10^{22} cm²) and moderate temperatures (35 K), whilst for instance N_2H^+ emits most efficiently towards lower temperatures (10^{22} cm²² and <20 K). The authors determine $X_{co(1-0)} \sim 0.3 \ 10^{20}$ cm²² (K km s²¹)¹¹ and $\alpha_{HCN(1-0)} \sim 30 \ M_{\odot}$ (K km s²¹ pc²)¹¹, which both differ significantly from the commonly adopted values. All in all, these results provide new evidence that when interpreting molecular emission, the typical assumptions that CO probes the bulk gas properties in galaxies, while molecular species such as HCN and HCO⁺ potentially characterize the densest gas environment, should be viewed with caution.



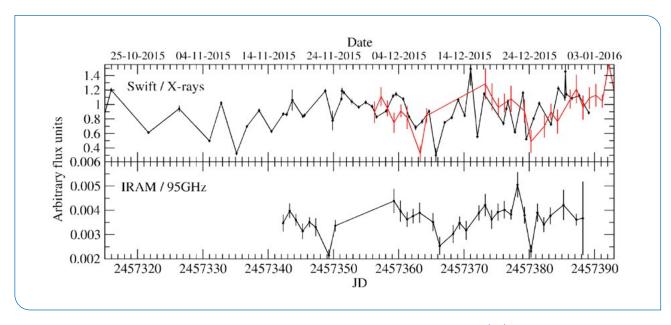
IRAM 30-meter telescope maps towards W49 of the integrated intensity of selected molecular gas tracers. The upper left two panels show the column density and dust temperature determined from data obtained with the Herschel Space Observatory. Overlaid on the upper left panels are contours of NH₂. All maps are smoothed to 60" and ordered by increasing critical density of the molecular transition. Work by Barnes et al. 2020, MNRAS, 497, 1972

Coronal activity in the accretion disk of NGC 7469?

Radio-quiet active galactic nuclei (AGN) differ from radio-loud nuclei in that they exhibit emission that is several orders of magnitude fainter. While the high brightness temperature of these sources suggests the presence of hot, non-thermal electrons, the actual physical origin and location of the radio emission source in radio-quiet AGN remains controversial. Current understanding is that the radio emission from radio-quiet AGN may be due to weak jets, nuclear star formation, or even coronal activity on the surface of the accretion disk.

In an effort to improve the understanding of the origin of radio-quiet AGN emission, Ehud Behar (Technion Physics Department/Haifa) and collaborators report on the flux monitoring over several weeks of the Seyfert galaxy NGC 7469 around 95 and 143 GHz with the IRAM 30-meter telescope, and in the optical, UV and X-ray with the Swift, XMM-Newton and WISE observatories. According to the authors it is the first time that light curves were measured simultaneously at millimetre-wave and in X-rays towards this source. While the flux density in both millimetre bands showed changes of about 50% over the 36-day observing period, consistent with the variability observed in previous observing campaigns, the X-ray light curve showed much larger flux variations of up to a factor of 5 over the same period. The high degree of variability in both millimetre bands and in the X-ray suggests that both sources of radiation originate from the same physical component of the AGN, and very likely from the corona of the accretion disk. Indeed, a tentative correlation between the millimetre and X-ray light curves suggests the presence of a ~14-day lag of the X-ray emission. If real, the lag could imply that magnetically heated electrons produce the millimetre waves, and ultimately lose most of their energy by emitting X-rays. Simultaneous monitoring in eight UV/ optical bands shows much less variability than in the millimetre bands and X-ray, implying this source of radiation is originating from a different AGN component, likely the accretion disc itself.

More precise millimetre-band measurements of a sample of X-ray-variable AGN are now needed, preferably also on time-scales of less than a day where X-rays vary dramatically, in order to properly test the physical connection between the two bands. Such campaigns require higher sensitivity than that available with the IRAM 30-meter telescope and would possibly make use of facilities such as NOEMA and ALMA.



Light curves of NGC 7469 in X-rays (top) and at 95 GHz (bottom). A tentative 14-day lag is overlaid in the top panel (red curve) that shows the 95 GHz light curve shifted forward by 14 days.

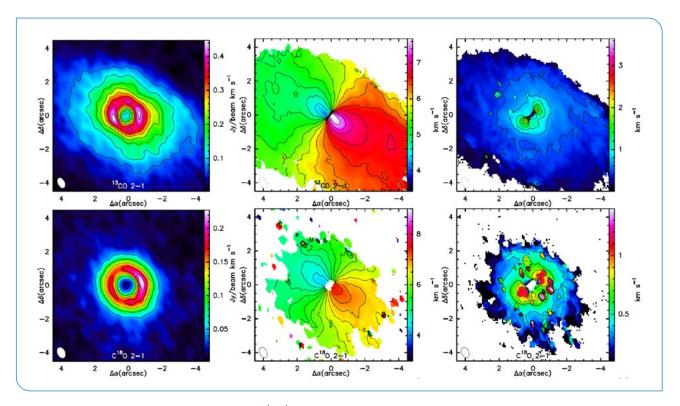
Work by Behar et al. 2020, MNRAS, 491, 3523

AB Aur, a Rosetta stone for studies of planet formation

Planets are known to form in circumstellar systems of gas and dust called protoplanetary disks, but the exact mechanism leading to planet formation remains unclear. While the gas in these systems is a key factor in the dynamical evolution of the protoplanetary disk, and in particular the dispersion of the gas appears to set the timescale for the formation of giant planets, key parameters such as the gas-to-dust ratio remain largely unknown. Therefore, a thorough characterization of the physical conditions and chemical composition of the gas is of utmost importance for understanding the disk evolution.

In an effort to characterize the gaseous disk around the Herbig Ae star AB Aur, a team of researchers led by Pablo Rivière-Marichalar (OAN/Madrid) performed a thorough spectroscopic study to determine the physical and chemical conditions in the disk using NOEMA. A total of 23 transitions from different species were observed, including 12 CO, 13 CO, 13 CO, 13 CO, 13 CO, 13 CO, SO, SiO, DCO+, HC $_3$ N, OCS, CCS, 13 CN, 13 CS and CO+. The authors used the integrated intensity maps and stacked spectra to derive reliable estimates of disk temperature and calculated the gas-to-dust ratio throughout the disk by combining the 13 CO and C 18 O data. Column density maps for the different species were also derived and used to calculate abundance maps. Finally, the observational results were compared with a series of astrochemical models to gain insight into the properties of the disk. The molecules observed show different spatial distributions, and the peaks of the distributions were found to not correlate with the molecules' binding energy. Using H_2 CO and SO lines, a mean disk temperature of 39 K was derived. The gas-to-dust ratio ranges from 10 to 40 across the disk, and abundance of SO (2 10^{-4}) with respect to 13 CO is found to be almost one order of magnitude larger than the value derived for H_2 CO ($1.6 \cdot 10^{-5}$).

The authors of the study thus show that AB Aur hosts a peculiar transition disk characterized by a high gas and dust temperature, a low gas-to-dust ratio, and large sulfur depletion. By determining for the first time the gas temperature and the gas-to-dust ratio across the disk, the researchers provide new information to further constraining hydrodynamical simulations.



NOEMA integrated intensity maps (left), intensity-weighted velocity maps (center) and velocity dispersion maps (right). Species shown are ¹³CO (2-1) (top), C¹⁸O (2-1) (bottom). The white ellipses in the bottom left corner of each map shows the synthesized beam at each wavelength.

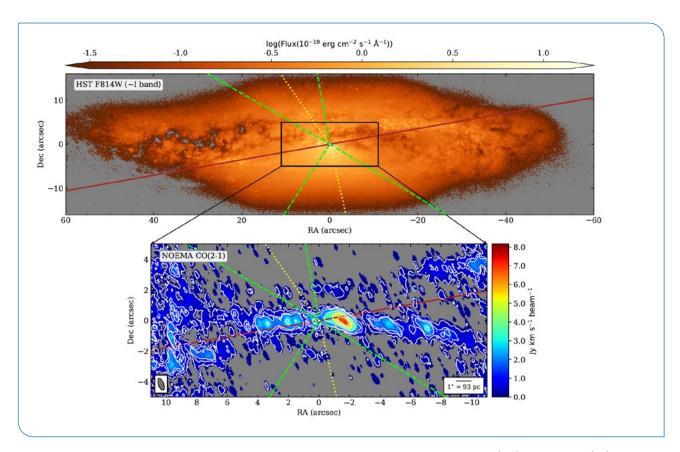
Work by Rivière-Marichalar et al. 2020, A&A, 642, 32

A search for outflows in the nuclear regions of Seyfert galaxies

Star-AGN coexistence makes galaxies excellent laboratories for the study of stellar and AGN outflows and their feedback. Outflows provide negative feedback when redistributing gas on a large scale in the interstellar and intergalactic environment, but can also provide positive feedback when associated with vigorous star formation. AGN-driven outflows are believed to play an important role in regulating the growth of galaxies, mostly via negative feedback. However, their effects on their hosts are far from clear, especially for low and moderate luminosity Seyfert galaxies.

To investigate this issue, Antonio Domínguez-Fernández (OAN/Madrid and IPARCOS-UCM/Madrid) and collaborators have obtained cold molecular gas observations, traced by the 12 CO (2-1) transition, using the NOEMA interferometer of five nearby Seyfert galaxies. The spatial resolution (\sim 30-100 pc) and field of view (\sim 5000 pc) of NOEMA allowed the researchers to investigate the CO (2-1) morphology and kinematics from the outer edges of the galaxy disks down to the nuclear regions.

CO (2-1) emission was detected in all five galaxies with disk- or circumnuclear ring like morphologies. Cold molecular gas masses were derived on nuclear (100 pc) and circumnuclear (650 pc) scales in the range from 10^6 to 10^7 M $_\odot$ and from 10^7 to 10^8 M $_\odot$, respectively. It appears that although non-circular motions are also present, most of the gas rotates in the plane of the galaxies. In NGC 4253, NGC 4388 and NGC 7465, the streaming motions are related to the presence of a large-scale bar. In Mrk 1066 and NGC 4388, the non-circular motions in the nuclear regions are explained as outflowing material due to the interaction of the AGN wind with molecular gas in the galaxy disk.



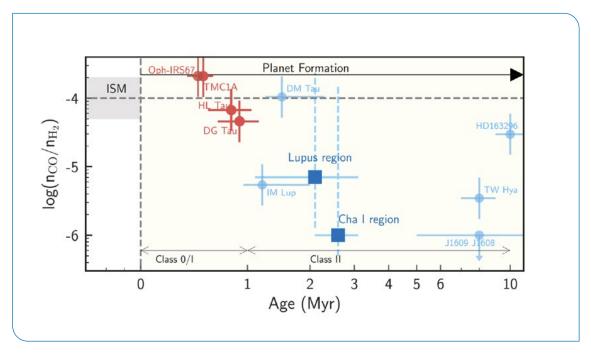
HST/WFC3 I-band image of NGC 4388 (top), NOEMA ¹²CO (2-1) integrated intensity map (bottom). The star marks the assumed location of the AGN, which corresponds to the 21-cm continuum peak from VLBI observations. Masked pixels are shown in grey. The green dashed lines represent the edges of the ionization cone and the dashed yellow line the cone axis. The brown solid lines are the position angle of the major axis of the main bars.

Work by Domínguez-Fernández et al. 2020, A&A, 643, 127

Although the spatial resolution may have an impact on the results achieved, the study clearly shows that it is only in cases of favourable geometry that it will be possible to detect molecular outflows. The authors conclude that the presence of galactic bars, present and past AGN-wind disk interactions, interactions with companion galaxies, and information about the kinematics of the ionized gas are necessary for an unambiguous and accurate interpretation of the kinematics of the cold molecular gas.

Evolution of volatile CO from the protostellar to the protoplanetary stage

Recent observations show that the CO gas abundance, relative to H_{2^l} in many 1–10 Myr (Class II) old protoplanetary disks may be heavily depleted by a factor of 10–100 compared to the canonical interstellar medium (ISM) value of 10^{-4} . When and how this depletion happens can significantly affect the compositions of planetesimals and the atmospheres of giant planets. It is therefore important to constrain whether the depletion already occurs at the earliest protostellar disk stage.



CO gas abundance measurements in protostellar and protoplanetary disks. The protostellar disks (<1 Myr) and protoplanetary disks (1–10 Myr) are in red and blue, respectively. The filled circles are for individual disks and the squares for averaged values of star formation regions. The TMC1A, HL Tau, and DG Tau data are from this work. Work by Zhang et al. 2020, ApJ, 891, L17

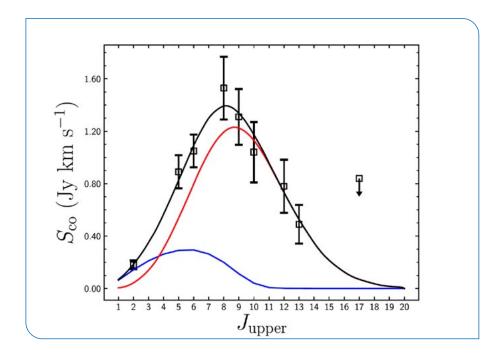
Zhang Ke (University of Michigan/Ann Arbor) and collaborators used NOEMA to obtain spatially resolved observations of C¹⁸O, C¹⁷O and ¹³C¹⁸O J=2–1 lines in three protostellar disks. They show that the C¹⁸O line emits from both the disk and the inner envelope, while the C¹⁷O and ¹³C¹⁸O lines are consistent with a disk origin. The line ratios indicate that both C¹⁸O and C¹⁷O lines are optically thick in the disk region, and only the ¹³C¹⁸O line is optically thin. The line profiles of the ¹³C¹⁸O emissions are best reproduced by Keplerian gaseous disks at similar sizes as their mm-continuum emission suggests small radial separations between the gas and mm-sized grains in these disks, in contrast to the large separation commonly seen in protoplanetary disks. Assuming a gas-to-dust ratio of 100, they find that the CO gas abundances in these protostellar disks are consistent with the ISM abundance within a factor of 2, nearly one order of magnitude higher than the average value of 1–10 Myr

old disks. If these CO abundances are representative of protostellar disks < 1 Myr, then there are two important implications: first, the CO gas depletion seen in Class II disks are results of processes that occurred inside disks rather than the infalling envelopes; second, the process is extremely efficient, as it depletes volatile CO gas by a factor of 10 within 1 Myr. The timescale of the CO depletion process therefore is comparable with the general timescale of planetesimal and planet formation. Whatever mechanisms drive the CO depletion, it can significantly affect the final compositions of planets. These results suggest that there is a fast, ~1 Myr, evolution of the abundance of CO gas from the protostellar disk stage to the protoplanetary disk stage.

Probing the SLED of a quasar-starburst system at z=6

Quasars at z~6 are first representatives of the coeval evolution of supermassive black holes (SMBHs) and their host galaxies. A large fraction of these early systems shows host galaxies with dynamical masses of $10^{10}-10^{11}M_{\odot}$. The huge amounts of dust and gas they contain are warmed up by massive star formation activity ($10^2-10^3M_{\odot}/yr$) and traced by their bright ^{12}CO emission. Observations of this molecule appear to be crucial to understand the physical and chemical conditions, the kinematics of the multiphase interstellar medium, and to study SMBHs and their host galaxies in the earliest stages of coevolution.

To investigate the molecular environment of a quasar-starburst galaxy at high redshift, a team of researchers led by Jianan Li (Peking University, Bejing) combined NOEMA observations of 12 CO (5–4), (6–5), (12–11), and (13–12) and ALMA observations of 12 CO (8–7), (9–8), H_2 O ($2_{0.2}$ – $1_{1.1}$), and OH+ (1_1 – 0_1) of the z = 6.003 quasar SDSS J231038.88+185519.7. They constructed the CO spectral line energy distribution (SLED), which they analyzed with the radiative transfer code MOLPOP-CEP. The best-fitting result suggests that two gas components, a warm (T_{kin} = 306 K) and a cold one (T_{kin} = 50 K), are necessary to explain the CO spectral distribution. The molecular gas density of the warm (cold) component is estimated to $\log(n(H_2)/\text{cm}^{-3})$ = 5.25 (4.20), and the corresponding column density to $\log(N(\text{CO})/\text{cm}^{-2})$ = 15.5 (18.0). Compared to samples of local ULIRGs, starburst galaxies, and high-redshift submillimeter galaxies, J231038.88+185519.7 exhibits higher CO excitation at $J \ge 8$, like other high-redshift quasars. The authors conclude that the high CO excitation, together with the enhanced L_{H2O}/L_{IRV} L_{H2O}/L_{CO} and L_{OH+}/L_{H2O} ratios, suggest that besides the UV radiation from young massive stars, other mechanisms such as shocks, cosmic-rays, and X-rays might also be responsible for the heating and ionization of the molecular gas. In the nuclear region any of these mechanisms might be present due to the powerful quasar and the starburst activity.



¹²CO SLED of J2310+1855 fitted with a two-component model. The black squares are the CO fluxes measured with NOEMA and ALMA. The CO (5-4), (6-5), (8-7), (9-8), (12-11) and (13-12) data are from this work. The blue and red lines are the best-fitted cold and warm gas components. Work by Li et al. 2020, ApJ, 889, 162

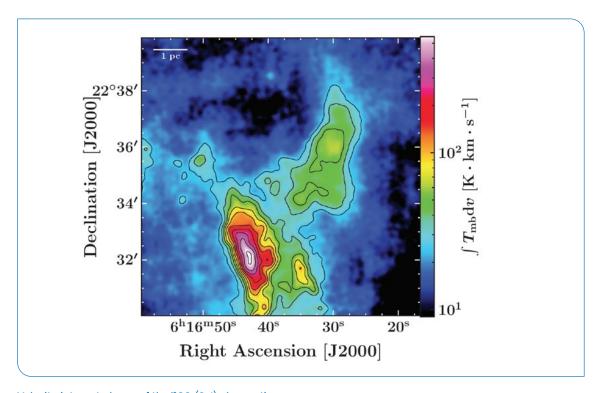
Interstellar anatomy the G region in the IC443 supernova remnant

Supernova remnants (SNRs) represent a major feedback source from stars in the interstellar medium of galaxies. During the late stages of supernova explosions, shock waves produced by the initial blast modify the chemistry of gas and dust, inject kinetic energy into the surroundings that gradually decays in turbulence and thereby alters star formation characteristics. Cosmic rays (CRs) accelerated in the earlier stages of the explosion and trapped in the shock fronts, interact with the dense medium and can generate strong *y*-ray emission.

To characterize the mechanisms of energy injection by an evolved SNR, its effects on local star formation, and identify all possible sources of ongoing acceleration of CRs by the SNR, a group of researchers led by Pierre Dell'Ova (ENS/Paris and Observatoire de Paris) started to investigate the stellar and interstellar contents of IC443, an evolved shell-type SNR at a distance of 1.9 kpc with an estimated age of 30 kyr. The researchers aimed at measuring the mass of the gas and characterize the nature of infrared point sources within the extended G region, which corresponds to the peak of γ -ray emission. The region was mapped with the IRAM 30-meter telescope in 12 CO and 13 CO, and complemented them with C^{18} O observations at the APEX telescope.

The observations reveal four molecular structures: a shocked molecular clump associated with emission lines, a quiescent, dark cloudlet, a narrow ring-like structure, and a shocked knot. Masses of the four structures were estimated to ~230, ~90, ~210, and ~4 M_{\odot} , respectively, and a total mass of ~1100 M_{\odot} was estimated throughout the rest of the field of observations. candidates. The results indicate that the shocked clump holds only a fraction of the total molecular mass in the extended G region, which shows that the mass associated with the ring-like structure and the cloudlet cannot be neglected when quantifying the interaction of CRs with the dense local medium.

The team finds 144 protostars in the G region, and concludes that they might provide a fresh source of CRs, which must also be taken into account in the interpretation of the strong γ -ray observations in the region.



Velocity-integrated map of the ¹²CO (2-1) observations carried out with the IRAM 30-meter telescope over the extended G region. The colour scale is logarithmic to enhance the dynamic range and emphasize the fainter molecular cloudlet.

Work by Dell'Ova et al. 2020, A&A, 644, 64







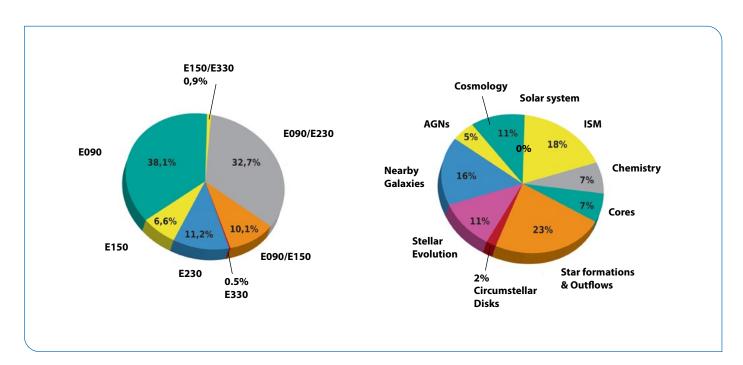
The Covid-19 pandemic has obviously strongly impacted the activity of the 30-meter observatory and Granada offices. The lockdown period and the strict safety measures imposed changes in the work patterns. The Granada premises were kept closed until early May and the staff worked exclusively in home office mode. Since mid-May, the offices started to be intermittently populated with strict protective measures. These measures are still in place early 2021. The attendance was limited to 50% of the nominal staffing. At the telescope site, the work patterns also changed, trying to set the occupancy of the observatory to a strict minimum for safe and efficient operation. The most noticeable change is that on-site visiting observers have no longer been allowed since March 2020 and are replaced by a fully remote visitor mode.

2020 was also marked by some very good news: after a very busy period of preparation, IRAM was eventually able to initiate the long-sought upgrade of the 30-meter telescope, namely the full renewal of the servo control system for the telescope mount, sub-reflector and wobbler by state-of-the-art hardware and software components, and the improvement of the surface of the primary reflector for better efficiency at any elevation and in day and night times. A thorough proposal for EU co-financing through the European Regional Development Funds (ERDF) programme was submitted in the last days of March, 2020 and has been awarded funding in July, for a total amount of 4.8 million Euro. The upgrade project will be hopefully completed by early 2023, ensuring that the 30-meter will be kept on the leading edge of research in the coming years.

Astronomical projects

Despite the Covid-19, pandemic, science observations at the 30-meter telescope continued almost without interruptions. Only during a short period of a couple of weeks in March/April, the telescope had to be stopped due to the lockdown in Spain. In May, telescope operation fully resumed with no visiting astronomers anymore and most observations done remotely. During the year, a total of 191 projects were observed. This number includes 9 large programs with EMIR or NIKA2, 7 Director's time projects, and 13 VLBI projects. About 21% of these projects were scheduled in pool weeks. The percentage of remote observations increased to 69% averaged over the full year. Galactic topics were addressed by two-thirds of the scheduled projects, while a third of the projects were devoted to nearby galaxies and more distant objects. EMIR was used during almost 81% of the observing hours, while HERA was used for less than 1% of the time, and NIKA2 usage increased to 19%.

During the scheduling year, astronomers could only visit the 30-meter telescope during the first months. In total, only 72 astronomers visited the telescope to support projects, 42 of which came to support the observing pools. Two groups of master students and their tutors visited the telescope to observe short projects as part of their training courses just in time before the lockdown started.



Usage of EMIR bands in 2020.

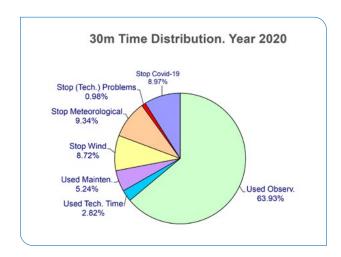
Time distribution of scientific categories observed in 2020.

Observatory operation

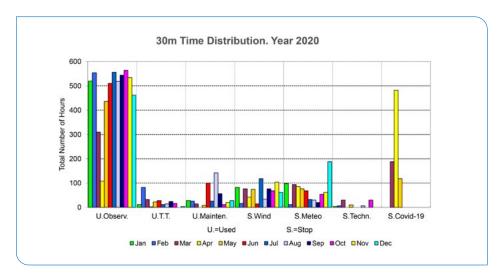
With the start of the first Covid-19 pandemic, on-site personnel at the observatory was been reduced to an absolute minimum, with just one transport per week. The company operating the Sierra Nevada ski resort completely (and suddenly) stopped operations, and the cable-car was no longer operational thereafter. Telescope operations were stopped for 3 weeks and were reduced to 12h night shifts due to the lack of personnel for another 4 weeks. However, since May 12 normal 24h shifts were resumed. All AOD shifts were done remotely until approximately mid-September. The visits of external observers to the observatory were stopped and replaced by the remote visitor mode. When remote visitor observations are not possible (e.g. inexperienced observers), they are carried out in service mode.

In spite of the difficult situation described above, the observing efficiency has been outstanding: around 64% of the total available time was allocated to scientific observations with either EMIR, HERA or NIKA2. The fraction of time lost due to the Covid-19 crisis was nearly 9%, and was partially compensated by the reduction of the time loss due to poor weather conditions (around 18%, versus 23% in 2019). The time fraction spent in maintenance and technical time was around 8% (this includes the telescope stop period due to the replacement of the tower cladding). Finally, the time lost due to technical problems remains extremely low, less than 1%.

In spite of the many restrictions in the activities of contractors imposed by the pandemic, several important refurbishment activities were started or resumed.



Usage of the total time at the 30-meter telescope in 2020. Note the nearly 9% time lost due to the Covid-19 pandemic.



Monthly time distribution of observing time, technical projects (T.T.), maintenance, time lost due to weather conditions (wind, snow etc.) and technical problems. Notice the time lost due to Covid-19 in March, April and May.

The most important and visible action was the replacement of the antenna pedestal cladding: the original plates were made of fibre cement with a substantial fraction of asbestos, a material proven to be very toxic. The refurbishment was carried out in two stages: first, the fibre-cement plates were removed taking all safety measures for workers; the wooden frame supporting the old plates and the insulating material were also removed. In a second stage, the new metallic supporting frame and sandwich panels made of steel sheets with rock wool isolation were installed. The new cladding is behaving very well: on the one hand, the day-time

thermal behavior is similar to the original, and on the other, the new cladding shows a good resilience to the impact of ice pieces falling from the upper telescope structure that usually caused breakage of the fragile fibre-cement plates every Winter.

A second important upgrade in the telescope area was the replacement of the air conditioning unit in the room hosting the spectrometers and computers. The old Liebert machine has been replaced by two new Stulz units of enhanced cooling capacity, each one with 22.2 kW frigorific power.





Removal of original fibre-cement plate.

Installation of the new asbestos-free cladding.



The 30-meter telescope with its new facelift.

Instruments

Radioelectric protection

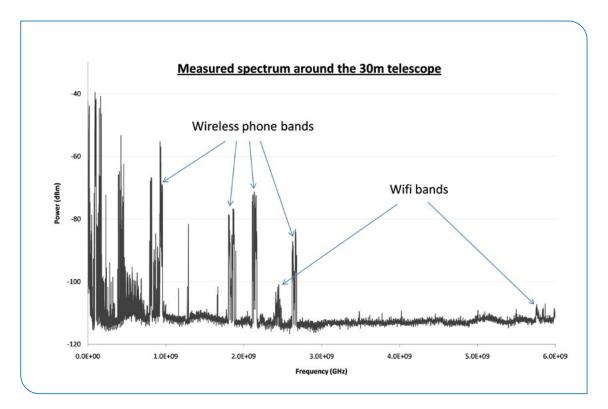
As the car collision radars, mobile phone technology and other wireless applications evolve, the spectrum around the 30-meter telescope is likely to get contaminated by radio-electric interferences. A recent event was the installation of a communication pylon in close proximity of the telescope. The tower, planned to be used as a private communication system for the ski resort, includes several low frequency (<1 GHz) sector antennas and a parabolic dish working over 10 GHz. After explaining the risks of EM interference from such transmitters on the observatory operation, these antennas were finally moved to a much further distance from the telescope.



In order to ascertain the situation concerning possible sources of EM interference around the observatory, a survey on the 0 – 20 GHz band was performed. Two very broad band omnidirectional disc-cone antennas, a low-frequency commercial one and a homemade smaller design for the higher bands, were used for the signal acquisition . Despite the high sensitivity of the instrumentation, no signal was detected above 6 GHz.

The data were acquired during the month of June 2020, a very quiet period due to the inactivity of the ski station and the current pandemic situation. The spectrum can thus be considered as very close to the best scenario and used to compare with future RFI surveys.

Setup used for the measurements of radiofrequency interferences in the IF bandwidth of the receivers.



Measured spectrum in the 0-6 GHz frequency range.

Backend

A significant VESPA maintenance effort was undertaken to ensure its availability for the next few years, fixing mostly the sampler boards whose stock had dwindled down to a critical level. A new version of the firmware and software for the broadband continuum detectors was prepared, which will enable faster sampling for pulsar observations and improved analysis of the power fluctuations of the heterodyne receivers. After validation of the prototype, a set of 6 barycentric references was built. The integration in the system still depends on further work on the interface and control software.

VLBI

In January 2020, a general EHT dress rehearsal was carried out at 1.3 mm, successfully obtaining fringes in several baselines involving the 30-meter telescope. However, the restrictions derived from the pandemic forced the cancellation of both the EHT (1 mm) and GMVA (3 mm) Spring sessions. Eventually, in October 2020 the global GMVA (3 mm) session took place; fringes were successfully obtained on the baselines joining the 30-meter and the other European telescopes. In December 2020, the first test at 3 mm between the phased-NOEMA (pNOEMA) array, comprising 8 antennas, the IRAM 30-meter and the OAN 40-meter Yebes telescopes took place. Fringes were successfully obtained on all baselines. To ease future VLBI sessions, in particular the increased need for flexibility in their scheduling, an automatic, remote switching between VLBI and normal operations has been fully implemented.

An important recognition was received in February 2020: several IRAM scientists and engineers involved in the production of the first image of the shadow of the supermassive black hole of M87 were awarded by the Andalusian Regional Government (Junta de Andalucía) with the "Bandera de Andalucía 2020" for the province of Granada in the category of Research, Science and Health.

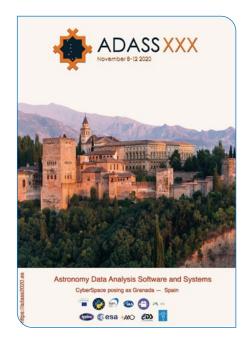
Computers & software

Regular operations changed dramatically in 2020 because of the Covid-19 pandemic. The IRAM staff switched to home-office, while the 30% fraction of remote observations rapidly increased to, and continues to be, 100%. To adapt to this new situation, the computer group strengthened the remote access infrastructure and logistics for both internal and external users. The OpenVPN server worked and is working exceptionally well. The VoIP system for telephones also allows calls to be answered on a regular computer with the use of commonly available, open-source software. In addition, several video-conference tools were set-up

A number of important improvements were implemented: first, the archive storage system was upgraded (both the master and the copy) increasing its capacity to close to 900 TB of usable space. We are currently at 55% capacity. Second, two critical machines (mrt-lx1 and mrt-lx3) were eventually virtualized allowing for



New layout of the telescope operators' (left) and astronomers' (right) working areas. greater resilience in case of failures and easy rollback of changes. The two virtual machines are currently working flawlessly. The operator and observer workstations were upgraded to faster computers with four 27" screens each in a 2x2 layout.



The Granada data center work was eventually finished. The new data center has almost twice usable space, better electrical installation and cooling, and three racks for servers, communications and storage respectively. The floor was also sealed from the parking space underneath. The second copy of the archive system was relocated to the refurbished Granada data center.

In the area of the operational software, PaKo has been adapted to the GFortran compiler, in order not to depend anymore on the extensive Intel Fortran Compiler. It is hoped this will allow IRAM users to install PaKo more easily at their home institute, in order to test scripts before starting their observing campaigns. The IMBFits for NIKA2 has been further developed to include support for polarization and also for the calibration source. Soon after the observations, the IMBFits are written to the archives in Granada and then mirrored in Grenoble.

A highlight of the year was the organization, as an on-line event, of the ADASS XXX conference. This multi-wavelengths conference focusses on "Astronomy Data Analysis Software and Systems" and gathers every year hundreds of experts in the field. It was a very successful event with almost 600 participants from all over the world and five full days of tutorials talks and BoFs ("birds of a feather", i.e. discussion groups). All contributions were recorded and are available online.

Facilities

An important highlight regarding the improvement of life conditions at the observatory was the installation of a new digital TV system, available in each and every bedroom. Spanish, but also French, English, and German channels are now available to all residents and visitors, making their stay more enjoyable.

Work was also performed to improve the security at entrances of the telescope and main building: electromagnetic locks and video-intercoms at the entrances have been installed. This responds to the evolving situation: since the inauguration of the 30-meter telescope, the access to its premises is much easier through the road or ski station, hence increasing the risk of intrusion.



The IRAM 30-meter cooks' team attended a course of vegetarian cooking.





The highlight this year was the addition of the 11th antenna to the NOEMA array. First fringes were observed with Antenna 11 on August 19, and after a successful science commissioning phase, NOEMA started regular science observations with all eleven antennas. In parallel, the observatory continued to be engaged in the next milestones: the transformation of NOEMA into a very-long-baseline interferometry (VLBI) station, and the assembly of Antenna 12. While work on these projects progressed relatively well, their schedule could not be fully met in 2020 due to the Covid-19 pandemic.

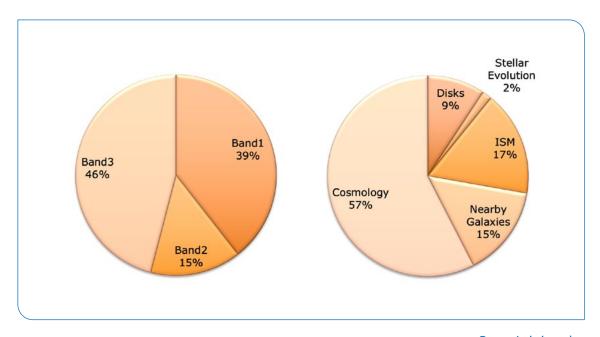
As in previous years, the NOEMA observatory continued its efficient operation, characterized by a high availability of antennas and instruments and a low technical downtime. The antennas, receivers and the correlator all performed well throughout the year. The combination of high system reliability and structured organization of work at the observatory helped to keep the scientific productivity at the highest level. To take advantage of the best weather periods and the most extended configuration (A) of NOEMA, technical activities were kept to the strict minimum during the winter period. All scientific observations were performed in service observation mode only.

To fully exploit the winter conditions at the beginning of the year, the interferometer reached its most extended configuration (A) at the end of January, operating with all ten available antennas. The antennas

were moved into the intermediate configuration (C) at the end of February and into the most compact configuration (D) in early April. Despite the complicated lockdown situation in spring, the observatory continued to observe, and thanks to exceptionally good weather conditions, even obtained a significantly higher scientific yield (by 14%) in the March-May period than in the two previous years over the same period. Despite difficult Covid-19 related working conditions, the NOEMA observatory was able to deliver Antenna 11, retrofit three first-generation antennas, perform the annual maintenance on the NOEMA antennas, pursue commissioning activities, and maintain full scientific productivity with the 10-antenna NOEMA array through the end of the summer semester. By the end of November, NOEMA was ready to start regular astronomical observations and to enter the winter semester 2020/2021 with all eleven antennas.

The program committee met twice during the year, around four weeks after the deadlines for the submission of proposals. It reviewed 219 regular proposals and recommended 142 of them (15% more than 2019). Over the year, NOEMA also received and accepted 7 Director's Discretionary Time (DDT) proposals. Including the backlog of projects from 2019, science goals from 115 proposals were scheduled in 2020 at the NOEMA observatory, including science from 2 Large Programs, 5 Max-Planck-IRAM Observatory Programs (MIOP), and 10 DDT proposals. This corresponds to 291 individual sub-projects (same number as in 2019) that received time on the interferometer.

As in previous years, NOEMA continued to provide unique and exciting scientific results and to demonstrate its effectiveness at exploring the interstellar medium in our Galaxy and in the high-redshift universe. As in previous years, the observing time requested to carry out galactic research was less than the time requested for extragalactic science. This testifies to the enduring and widespread interest of the extragalactic community, which has been persistently growing over the past years. The largest amount of observing time was invested in the compact and intermediate configurations of the interferometer between spring and autumn. The percentage of observing time invested in 2020 on science programs for the user community was on average 49% of the total time, or equivalently, 178 days. An additional 17% were invested on technical operations and developments: software work, commissioning, technical verifications and installation of new equipment (11%), array reconfigurations (4%), and receiver tunings for user projects (2%). The remaining 34% were lost due to weather constraints.



Requested observing time by receiver band and science category.

Ongoing work and activities



Science observations with Antenna 11.

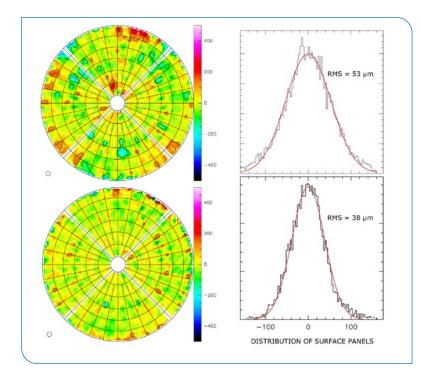
Antenna 11 commissioning

The integration of Antenna 11 into the NOEMA array was the most significant event and among the most exciting activities of the commissioning and science operations (NCSO) team in 2020. The start of the NCSO phase of Antenna 11 began on August 17, 2020, with the delivery of the antenna to the commissioning team. The goal of the NCSO was to take the antenna from the stage reached at the end of construction to an instrument that meets scientific requirements. This stage, which is aimed at checking that all parameters are within the specifications, provides first quantitative information on the antenna performance in terms of sensitivity, image quality and accuracy. To achieve this objective, the NCSO team worked in close collaboration with the computing, engineering and construction teams, and used specific test procedures already applied to the previous antennas. The commissioning activities included antenna safety checks, pointing, focus and tracking accuracy verifications, improvements of the surfacy accuracy using phase-coherent holography, and the assessment of a number of key performance indicators such as the antenna interferometric efficiency. First results from the NCSO of Antenna 11 helped to improve operational reliability and to identify areas where further work had to be done. For instance, due to an error in the manufacturing process, the panels of rings 1 to 4 were found not to fully meet the specifications and will have to be replaced in 2021. Ultimately, after the successful commissioning campaign, which demonstrated NOEMA's once again increased performance, Antenna 11 joined the array and began routine scientific operations.

Antenna upgrade project

The antenna upgrade project is aimed at bringing first generation antennas (1-6) to the same technological and operational standard as the new NOEMA antennas (7-12). In the frame of this project, another important milestone was achieved in the fall of 2020 with the completion of the electro-technical refurbishment phase on Antennas 2, 3 and 6. To ensure that technical modifications were in line with science requirements, the NCSO team ran a series of test procedures. At the end of November, the antennas were declared technically ready for science operations and made available for routine science observations. Work on the retrofitting of the remaining first-generation antennas will continue in 2021.

As in previous years, the surface quality of all antennas was verified by means of holographic measurements and readjusted during the maintenance period, and iteratively improved when deemed necessary. All in all, the primary surfaces of all six antennas show excellent stability and a median accuracy of 36µm RMS.



Surface panel distributions of Antenna 11 relative to the ideal paraboloidal surface (left) and associated surface RMS (right). On the top, holographic results of the antenna surface after delivery to astronomical commissioning; on the bottom, improved antenna surface after three iterations of panel adjustments.

Commissioning of the 14-channel radiometer

After the initial testing period in 2019, the new generation 14-channel radiometer underwent an optimization phase in the Grenoble frontend laboratory and was then installed again on Antenna 8. In particular, internal temperature stability and homogeneity were significantly improved and the load table mechanism was revised. The radiometer was integrated into the real-time observing and data recording system, and has been correcting the atmospheric phase turbulence on the baselines of Antenna 8 since December 2020. Although the radiometer is currently operating in an emulation mode that uses only 4 of its 14 channels, its performance is already comparable to the best first-generation radiometers. Work was underway at the end of the year to use the data from all 14 channels for even more sensitive atmospheric phase corrections. Current plans are to equip two more antennas with a 14-channel radiometer by the end of 2021, and to have all antennas equipped with a radiometer in 2022.

Large IF bandwidth observing mode

As part of the development of new observing modes for NOEMA, the feasibility of covering frequency ranges larger than the IF bandwidth by cyclically switching between two LO frequencies separated by 7.744 GHz was investigated. Such mode would offer the possibility of performing spectral scans over contiguous bandwidths of 32 GHz, under practically identical observing conditions. First very promising tests conducted in late 2020 with a subset of antennas have shown that switching LO tunings between two selected frequencies shows repeatable and stable performance within specifications. Should further investigations in 2021 validate the feasibility and efficiency of the observing mode, first science demonstration observations could then be conducted during the year.

User support

The NOEMA Science Operations Group (SOG) has been largely successful in meeting the goals and overcoming the organizational and technical development challenges in the NOEMA project despite the constraints of the pandemic.

The central mission of the SOG is to ensure that the NOEMA Observatory provides users with the means to conduct state-of-the-art research. The SOG astronomers work towards optimizing the scientific performance,

efficiency and return of the interferometer, either on site or remotely from the Grenoble headquarters. They provide technical support and expertise on NOEMA to researchers and visiting astronomers who have questions about the instruments, observing procedures, data reduction and calibration, pipeline data processing, and archiving of NOEMA data. Providing the best science data is at the core of the SOG mission.

Internally, the SOG interacts with the scientific software group on developments that may impact the long-term future of the interferometer, performs the technical review of scientific proposals, works with the technical groups to ensure that operational requirements are met, keeps user documentation up to date, and performs operational testing and commissioning tasks. Among the most important activities in 2020 were the commissioning of Antenna 11, performance assessments of the retrofitted first-generation antennas, of the 14-channel prototype water vapor radiometer installed on Antenna 8, and of the dual-band receiver installed on Antenna 9, and support for two VLBI test campaigns.

In normal times, IRAM headquarters hosts a regular stream of visiting astronomers from around the world who stay at the institute for periods ranging from a few days to a few months.

Starting March 2020, advice and assistance to calibrate and analyze 56 NOEMA projects remotely from their home institutes were provided to 63 astronomers from Europe and overseas for a total of 326 days. Prior to the lockdown period, face-to-face assistance was also given to 8 investigators visiting IRAM Grenoble. In total, the SOG invested 357 days (27% more compared to 2019) to reduce and analyze data from the interferometer. 63 science projects received support and advice. Compared to previous years and as a result of the global Covid-19 measures, the overall level of remote user support increased drastically. IRAM astronomers have been leading or collaborating on 50 projects in which they were directly involved. Given the increasing demands for support and supervision of external investigators, the computer system and network infrastructure at IRAM headquarters were reorganized to provide full computing power to individual users when necessary.



Members of the NOEMA Science Operations Group (SOG) at a Zoom meeting.

Data Archive

The data headers of observations carried out with the NOEMA and the former Plateau de Bure Interferometer are archived at the Centre de Données astronomiques de Strasbourg (CDS), and are available for viewing via the CDS search tools. In 2020, the archive contained coordinates, on-source integration time, frequencies, observing modes, array configurations, project identification codes, etc. for observations carried out in the period from December 1991 to September 2019. The archive is updated at the CDS every 6 months (May and October) and with a delay of 12 months from the end of a scheduling semester in which a project was observed in order to keep some of the information confidential until that time.

Access to the science data is initially limited to the principal investigators of the observing programs and their delegates. While the proprietary period of Large Programs is set to end 18 months after the end of the last scheduling semester in which the program was observed, the proprietary period of science data from standard NOEMA observing programs is set to terminate 36 months after the end of the last scheduling semester.

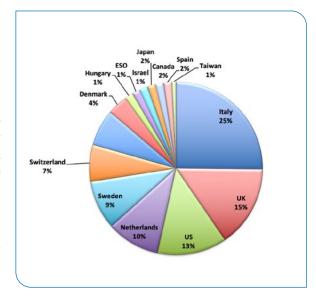
The IRAM Data Archive is the collection point for research carried out at the IRAM observatories in the framework of a Large Program and of regular science projects. The goal is to provide open access to calibrated images, data cubes, and visibility data from NOEMA, the former Plateau de Bure interferometer, and the 30-meter telescope. During the last year we have received more and more requests for data from the archives, and as we expect this number to grow in the coming years, developments are planned to facilitate access to the archives via the web in the future.

The science-ready products of Large Programs are made available to the astronomical community at the end of the regular data proprietary period. The archive is the result of a joint effort between IRAM, the principal investigators of the Large Programs and their collaborators.

Radionet transnational access

The interest of the European community in using the NOEMA facility has remained high during the 2017 to 2020 period covered by the Horizon 2020 initiative. NOEMA received more than 800 observing proposals from more than 1500 unique users originating from 45 countries. In the same period, NOEMA counted over 300 RadioNet eligible proposals submitted by researchers affiliated to scientific institutions from 10

European countries (IT, IE, UK, SE, CH, NL, DK, HU, DE, ES), one intergovernmental organization (ESO) and 4 non-European countries (TW, US, JP, CA) and organizations. Observing time was granted to 68 RadioNet eligible proposals. The observing capabilities offered to RadioNet eligible groups benefitted more than 130 researchers with a well-balanced distribution between senior scientists (66%) and younger PhD students and post-doctoral researchers (34%), and with an increasingly balanced distribution between female (44%) and male scientists (56%). Seven eligible teams benefited from financial support of their travel costs to visit IRAM headquarters in Grenoble and reduce their data with the help of IRAM experts. Despite the fact that RadioNet has come to an end in 2020, this successful Europe-wide initiative will continue in the coming years under the Opticon-RadioNet Pilot (ORP) program, still within the Horizon 2020 research framework program.



RadioNet PI User Distribution, 2017-2020.

VLBI

NOEMA participated in the Event Horizon Telescope (EHT) dress rehearsal on January 23 and 24, 2020, operating in single dish mode. However due to the Covid-19 pandemic and its impact on the EHT operations around the world, the EHT session scheduled for March-April 2020 was cancelled. As a result, NOEMA also did not participate in the Global Millimeter VLBI Array (GMVA) spring campaign. First 3 mm VLBI observations, which validated the full phased capabilities of the interferometer, were successfully performed in December together with the IRAM 30-meter telescope and the Yebes 40-meter telescope. While the fringe analysis by MPIfR Bonn indicated that work remained to be done to cope with long observing sessions, the December test demonstrated the excellent sensitivity of the phased NOEMA array for upcoming VLBI observing campaigns.



Compared to the previous phased array mode that was dismantled in 2017 with the arrival of PolyFiX, progress is considerable: where the old system worked for a maximum of 6 phased antennas with a maximum data rate of 1 GB/sec corresponding to a maximum bandwidth of 256 MHz on the sky, the present system is ready for 64 GB/s data rates and 12 phased antennas, with the possibility to go to 128 GB/s if the recording capabilities are upgraded. The new active beamformer allows to optimize the antenna phasing in real-time, improving efficiency over all observed bands.

While further tests with the interferometer are still planned in 2021, NOEMA is expected to participate in the spring 2021 science observing campaigns. NOEMA will join the IRAM 30-meter telescope as key elements of the EHT and GMVA networks and strengthen the role of European observatories in these global networks.

VLBI Mark 6 recorders installed at NOEMA.

Observatory operations

Following the outbreak of Covid-19 in France and the lockdown imposed mid-March, the activity of the NOEMA observatory has been deeply impacted, but has not stopped.

At first, a strict protocol was put in place where only the 3 teams operating the instrument (7 persons) were allowed on site. These 3 teams took turns to ensure the site and instrument security, and operate the observatory in a nominal way until the end of April. Science observations were not stopped. During that period, other NOEMA staff members have worked in home-office or on tasks to be carried out off-site.

Taking advantage of this period, a sanitary protocol was designed to allow as many activities as possible to be carried out on the site. These rules were implemented at the beginning of May, allowing to resume the construction of the antennas and the maintenance operations, in addition to the science observation activities.

It was however decided to strictly limit or postpone any other activity which was not essential to the good operation of the site.

Maintenance, construction and retrofit of antennas

For the eighth consecutive year, maintenance was carried out simultaneously with the construction of new NOEMA antennas. However, and contrary to previous years, in order to accelerate the delivery of the last antennas, the assembly of Antenna 12 started while Antenna 11 was still under construction in the hall, which had the direct consequence of occupying the whole maintenance & construction hall. Hence, the maintenance

of the antennas in operation was not possible in the hall until Antenna 11 was moved outside on August 10 for its commissioning.

Due to these organizational constraints, only the first-generation antennas (Antennas 1 to 6) were maintained in the hall, from mid-August until early December. Antennas 7 to 10 have been inspected outside, and the most urgent corrective maintenance operations have been performed on-site.

In parallel to the usual maintenance operations, the upgrade of the azimuth and elevation drive systems, which started the previous year on Antenna 1, was successfully continued on Antennas 2, 3 and 6. Each of these operations required 4 weeks of work, carried out in parallel with the usual maintenance operations in order to minimize the impact on the schedule.

The construction of Antenna 12, the last antenna in the frame of the NOEMA project, was on schedule at the end of 2020. It is planned that it will be available for scientific observations in the winter semester 2021/2022.



Construction of Antennas 11 and 12.

Site Maintenance

Among the tasks achieved despite the constraints imposed by the Covid pandemic are important maintenance works on the cable-car. One of the 43-tons carrying rope was successfully replaced, and the old cable was cut into a dozen sections before being removed. During this shutdown phase, the carriage and the suspension of the passenger cabin were sent for maintenance, thus bringing forward by a few months the regular date of their inspection.

All these operations were carried out according to a rigorous schedule and the cable-car could be put back into service after only a few weeks stop.







Section of the old cable composed of numerous strands.



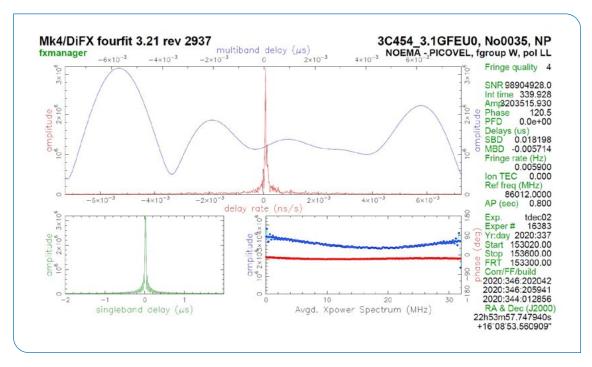


BACKEND GROUP

NOEMA phasing project

Work on the VLBI observing mode has been pursued during 2020. Early January, a new master synthesizer with an improved phase noise has been installed in view of the 2020 EHT dress rehearsal and spring observing session (finally canceled due to the pandemic situation).

Then, by fall 2020, the development of the VDIF formatter board has been resumed to correct for some minor flaws in the hardware. Also, the firmware in charge of encapsulating the beamformer data into VLBI standard digital frames (for transmission to the Mark6 recorders) has undergone extensive testing in collaboration with the MPIfR VLBI technology division.



First NOEMA Fringes in Phased-Array mode on 02 Dec. 2020.

Despite some remaining but non-blocking issues in the firmware, it was decided to run a test in phased-array mode by early December (4GHz 1SB Dual Polarization). With participation of the IRAM 30-meter telescope, fringes at 3 mm at a 32 Gbps recording rate were successfully obtained for the very first time with NOEMA/PolyFiX with 8 active antennas in the phased array. This validated several parts of the VLBI setup, including the real-time control of the beamformer associated with the phase solver software.

Following this milestone, the fabrication of a new batch of VDIF formatter boards has been launched to complete the frequency coverage as required by the EHT frequency plan. Thus, support of the full EHT configuration (4GHz 2SB Dual polarization @64 Gbps) is expected to take place in 2021.

Antenna reference transmission over fiber

Following the lessons learned from the 2019 test campaign, additional hardware and software tools have been developed in order to acquire data with a more convenient and reliable test setup. New on-site measurements of the LOREFoF (LO Reference over Fiber) subsystem were then able to take place during spring 2020.

Loopback tests in a very distant station with the current coaxial cables or with fibers have shown a similar phase drift on short time scales but significant differences over long time scales. This is in coherence with the higher temperature sensitivity of fibers that was characterized last year. Sensitivity of the LOREFoF demonstrator module to EMI has also been highlighted during the tests, especially regarding spurious magnetic fields likely radiated by power-supplies.

With these findings in mind, a first prototype has been developed with the goal of an installation in the limited space available within the receiver cabin to perform in-antenna tests in a configuration as close as possible to the final configuration.



The Test of Reference over Fiber subsystem.

Digital correlators

Additional digitizer cards have been installed in PolyFiX to support the 11th antenna delivered in September 2020. In parallel, work on the second correlator PolyFiX-2 has continued, to prepare for the arrival of the dual-band receivers system. Procurement of some critical components has continued. For instance, a new batch of optical links has been ordered to get rid of the fluctuating manufacturing delays usually encountered on this kind of specific products. At the same time, the correlator boards for PolyFiX-2 have been fabricated. Test of the cards has been started and will extend throughout next year.

Evaluation of very wide bandwidth A/D converters

The next multi-beam receivers currently under development will generate a so far never offered bandwidth. Processing the numerous IF bandwidth implies direct digitization in the receiver cabin. In collaboration with a leading company in A/D converters, some measurements have been performed on a preliminary prototype ADC chip which should be able to digitize at least a 4 to 12GHz instantaneous bandwidth.

Very promising performances have been measured, validating the ADC selected architecture. A design iteration is foreseen to match the high level of integration required from the chip level up to the boards and system levels.



FRONTEND GROUP

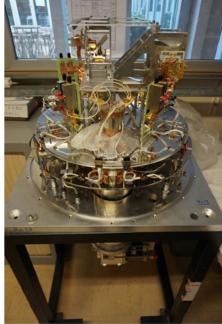
NOEMA operation and maintenance

The NOEMA receivers and cryocooler systems could not be properly serviced during the maintenance of 2020, due to the Covid-19 impact and restrictions on operations. It is planned to recover this delay during the upcoming maintenance of 2021. Similarly, almost all of the planned upgrades could not be performed and should take place normally in 2021.

NOEMA receivers

In 2020, the last receiver for the NOEMA 12^{th} antenna was assembled, together with an additional spare receiver (Rx13) for the Grenoble facilities, allowing an operational receiver to be always available in the IRAM laboratory facilities for testing upgrades.





The two final NOEMA receivers. Left: Receiver for the antenna 12. Right: spare receiver almost completely integrated.

New local oscillators for the NOEMA receivers

Two prototypes for local oscillators for band 1 and bands 2/3 using YIG oscillators were validated. The Band 1 YIG local oscillator is essentially a copy of the existing one used currently in Antennas 7 to 11. A version integrating a power amplifier was tested, allowing to get much more power, as needed for future multibeam receivers. In 2021, the production of all Band 1 YIG oscillators will be completed and we can aim at having only YIG-based oscillators at the observatory by the beginning of 2022.

The Bands 2/3 YIG oscillator was first tested with the Antenna 10 in 2018, but a parasitic signal made part of the LO unusable. It was therefore upgraded and modified in such way as to remove this harmonic (originating from the YIG oscillator fundamental frequency intermodulations at various multipliers). That modification was verified on Antenna 7 in 2020 and after successful validation, the production series for the

12 YIG-based local oscillators can start, to equip all antennas by end of 2022. The goal is therefore that by that time, there will only be YIG-based LOs at NOEMA, which will considerably improve reliability and tuning times, and will allow implementing new observing modes, with fast LO switching capabilities.

Dual-band NOEMA receivers

The first receiver with dual-band capabilities was installed in November 2019 on Antenna 9. In 2020, several commissioning tests were performed, allowing to compare the performance of that receiver compared to optimized single pixel observations. While the results were generally good, a point of concern was the decreased efficiency seen in Band 3 with the new optics of that receiver. In particular the addition of a lens to adjust the focus of Band 3 is likely to be responsible for this 10% degradation. New lens designs were produced in 2020 to correct for this problem.

Two new dual-band systems are under assembly. To this effect, two receiver support frames (so-called "chassis-A") were manufactured in 2020, allowing to accommodate dual-band receivers. An upgrade was also done to replace the translation stage used in the prototype by a similar model but 2.5 times faster. In 2021, if used in combination with the two receivers available in the laboratory, two new full dual-band systems could be installed.

In parallel, since this dual-band mode of observing will deliver twice the IF bandwidth, the Frontend group started in 2020 the design and production of the components needed to double the number of warm IF channels and laser-based transport of the IF signal via optical fibers.





Two new frames called "chassis-A" were manufactured specifically for the dual-band mode of operation. They are partially integrated with the translation stages, optics and moving mechanism used for the calibration.

NOEMA new generation water vapor radiometers

The first prototype of the second-generation water vapor radiometers (WVR-2G) was considerably improved in 2020. A substantial effort was carried out in order to make it more stable by decoupling it from the varying temperature conditions of the receiver cabin. It performed at least as good as the best existing radiometers but with many more channels (14 compared to 3). It also provides added robustness against electro-magnetic perturbations coming from satellites or other microwave links. The series production for 7 new generation radiometers has started in fall 2020. The first two radiometers will be available end of 2021, and all of them will be produced by summer 2022.

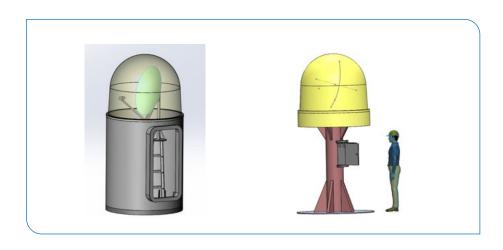


Radiometer prototype in the laboratory after the final upgrades in January 2020, before re-installation on Antenna 8 in February 2020.

Phase monitoring project

This project which started in 2018 in collaboration with the SMA group from the Center for Astrophysics (Harvard, USA), aims at providing a real-time, permanent monitoring system of the observing conditions at the NOEMA site.

In 2020, the final testing of a two-antenna prototype system took place with two parabolic antennas enclosed in protective 1.5 m radomes, separated by 378 m. From September to November, both parabolas were pointed towards the satellite Thor 6 at 0.8° West. The goal was mainly to test new Low Noise Blocks (LNB) that are collecting the satellite incoming signal on the antennas. The system provides real-time atmospheric phase variation to the operators through a web-based page.



Possible solutions for the infrastructure of the phase monitoring project. The parabola has to be enclosed in a protective radome, and installed at sufficient height to prevent snow accumulation.

As the results were very promising, it was agreed that a permanent 3-station system will be installed on the Plateau de Bure in the years to come. After a first infrastructure draft, two simpler versions are under study. Data acquisition system, electronics racks, and software will also have to be improved for better maintainability.

AETHRA Radionet - Multibeams

2020 was the last year for the European project Radionet-AETHRA, in which IRAM has been specifically working on the development of small size array receivers using 1) direct amplifiers (LNAs) in the 70-116 GHz frequency range, and 2) 1x7 pixels at 230 GHz, with mixers, amplifiers, and LO distribution integrated in the same mechanical blocks.

In 2020, the cryostats for both projects were fabricated by two external companies. The impact of the Covid-19 pandemic was high and created substantial delays. Still, the cryostats were delivered and assembled in the IRAM laboratories.



AETHRA 3 mm array receiver. View of the inner contents of the cryostat. Two cold stages are used at 60K and 12K. The cold optics can also be seen. One full line of 3 pixels will populate the receiver.



Test setup, with the cryostat at the middle left. A new spectrum analyzer with noise figure measurement capabilities was purchased in 2020 allowing to noticeably speed up testing.

100GHz 3x3 HEMT receiver

This project is done in close collaboration with the MPIfR in Bonn, the Fraunhofer Institute in Karlsruhe and INAF in Italy. The array includes a 1x3 pixels line for the 75-110 GHz band using active OMT blocks (with HEMT amplifiers integrated in the OMT block). The design is done for 3x3 pixels, but currently, only 3 active OMT blocks are available, sufficient for proper demonstration and testing.

230GHz 1x7 SIS mixer receiver

The 1x7 pixels for the 200-270 GHz band is using 2SB SIS mixers developed during the previous Radionet project AETHER. The cryostat is based on the NOEMA cryostat, scaled down slightly, and uses the same cold head, a 3-stage Sumitomo Industries RDK-3ST allowing to cool down the detectors to about 4K. The cryostat was delivered in November 2020 and integration followed soon afterwards.



AETHRA 1x7 pixels for 230 GHz using 2SB SIS mixers. The cryostat is a scaled version of the NOEMA cryostats and will use the same Sumitomo 3 stage cold heads, allowing to reach temperatures below 4 K.

CMD test cryostat

A new test receiver was designed at IRAM. It will be used in the Grenoble headquarters to measure cold mirror deformations (CMD) at temperatures down to 4K. It is particularly interesting to measure the NIKA-2 dichroic thermal deformations during the cooling process. To do so, the system will use a laser array and measure the reflected signal positions during the cooling phase to estimate deformations. This setup will eventually be available to measure any optical or electrical components at cold conditions.

SUPERCONDUCTING DEVICES GROUP

As in the last few years, the superconducting devices group has been able to focus most of its work to development of new technologies for upcoming instruments. At the same time, the group has continued to work on upgrading the machine park, to be ready for future demanding projects such as the multibeam receivers for the 30-meter telescope. Despite a total stop of the clean room during the 8 weeks of lockdown in spring, noticeable advancements on all fields were possible.

Machine park upgrades

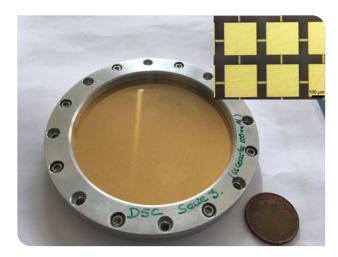
After a significant investment in 2019 in a new e-beam lithography system, attention has been focused on defining and finding replacement for critical but end-of-life equipment in the clean room. A small window of opportunity in early Autumn, between the two pandemic peaks, was used to investigate and select a new surface profiler tool. This critical instrument is used in all IRAM technology steps to measure film heights and film stress. The new surface profiler will arrive early 2021, and will allow these steps to be done in an automated fashion, permitting more reliable quality control and a better device yield.

SIS junctions on 10µm thick silicon

The major research focus for the Superconducting Devices Group in 2020 as in 2019 has been a continuation of the development of SIS junction technology on 10 micrometer thick silicon substrates. This new process, based on silicon-on-insulator (SoI) technology, has several advantages over the classical approach, most importantly the possibility to design junctions with a wider IF bandwidth. Moreover, it allows to use beam-lead technology for advanced packaging, and in the future advanced integration of functionality on chip. In 2020 two major technology issues have been tackled; they were linked to stress in the beam leads, that rendered packaging and electrical contacting of these fragile devices very difficult. Devices can now reliably be packaged, and first confirmation tests on increased IF bandwidth of these junctions are foreseen for early 2021.

Dichroic filters

In-house development of dichroic filters for dual-band operation has continued. These filters consist of an assembly of metallic structures on thin quartz substrates. Tiny air gaps between the different substrates severely hamper their performance. Therefore, we have investigated techniques to glue together these wafers without introducing inhomogeneity due to varying glue thickness, nor introducing supplementary dielectric loss in the glue. These developments have significantly improved devices performances, which is now comparable to or better than that of commercially available filters.



Prototype dichroic filter for NOEMA dual-band operation. The filter consists of 5 layers of 1-3 μ m thick gold plated patterns (two layers are shown in the inset), deposited on 4 fused silica wafers of 163 μ m thickness, that are subsequently glued together.

Other developments

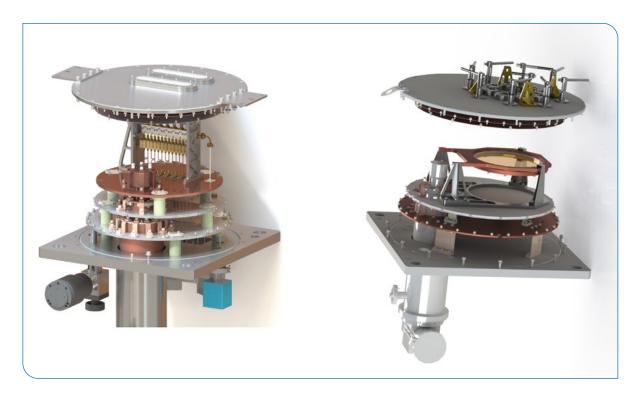
Other development projects from 2020 include production of beam-lead devices for the 100-GHz HEMT prototype receiver array, a thorough investigation of the causes of frequency dispersion in large KID arrays, made possible by the automated measurement routines of our newly acquired e-beam lithography system, and small series productions of devices for various projects.

MECHANICAL GROUP

The workshop received more than 100 requests for the production of mechanical high-precision elements, in particular many items for frontend projects (mixer blocks, horns, LOs, etc). This served several NOEMA projects, as the new generation Water Vapor Radiometers or the dual-band receivers, as well as the RadioNet AETHRA-supported development of multi-beam prototype arrays.

Cryostats

Three cryostats were designed almost simultaneously by the Frontend group: two multibeam receiver prototypes and one lab cryostat to study deformations at low temperatures. IRAM adopted a different internal organization for these projects, as compared to the previous generation (NOEMA receivers) cryostats. The mechanical group engineering office has produced the 3D drawings and was also in charge of the procurement of these cryostats, which were produced by external companies.



3D mechanical drawing of the 7-pixels prototype array and CMD test cryostats.

NOEMA antennas construction

2020 saw the completion of Antenna 11, which was delivered to the commissioning team in August, with a delay of a couple of months only due to the Covid-19 pandemic. The construction of Antenna 12 started during the fall, with the assembling of the large mechanical pieces of the pedestal. IRAM selected a new contractor in charge of the production of the quadruped legs, a critical element of the antenna reflectors.

In parallel, an important development started in 2019 was continued: because the future baseline extension includes a track segment with a slight (<2%) slope, the antennas must be equipped with a security braking system ensuring their velocity can be kept under control in any scenario (e.g. failure of the main drive systems). Each antenna has a weight of approximately 130 tons. Several options were investigated, including passive and active systems, and their pros and cons were carefully assessed. Final choice was to equip antennas with a backstop: in case of emergency, a hydraulic clamp stops the antenna and a disc antiskid system allows it to be moved only in the upward direction, i.e. back to the flat part of the track. In order to precisely assess the performances of such a system and to optimize its control, a test-bed was conceived and constructed. Tests early 2021 should lead to a final design to be implemented on NOEMA antennas.



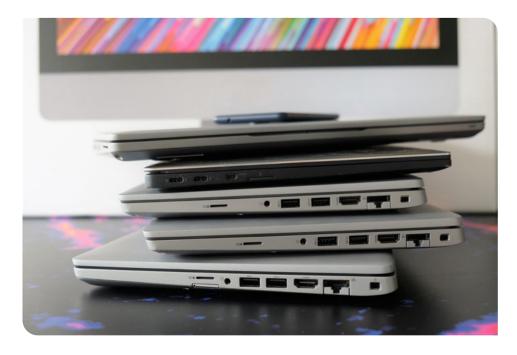
Test-bed of the emergency brake system to be installed on NOEMA antennas.

COMPUTER GROUP

Major upgrades of several software tools used at IRAM were implemented, in particular the tools managing the park of computers running under various operating systems. This was even more important in the context of the Covid-19 pandemic, which triggered a huge increase of remote home office work.

A set of new web tools for home office

To ensure that all IRAM staff working from home can rely on collaboration services, various web-based tools were deployed or upgraded. This included in particular an Intranet wiki (DokuWiki), a video conference (JitsiMeet), an instant messaging (Rocket.Chat), an issue tracker (Zammad) and a file transfer (Lufi) service. All these tools are open-source and operated by the Computer Group on the IRAM computing infrastructure.



Several operating systems are supported.

Managing Linux...

During the last years, the choice was to run one single application per virtual machine, because it allows updates to be performed with a better granularity. The drawback is that the number of virtual machines has increased to around 100. Specific tools to configure such large farms of virtual servers were needed. Among the different existing software solutions, IRAM has chosen Ansible to manage the configuration of its Linux servers because it is an agent-less and easy-to-learn solution.

Windows...

Until last year, the park of Windows computers were managed with a set of individual tools like WSUS, MDT, GPO and PowerShell. Nevertheless, with more than 100 managed Windows devices, a more powerful solution

was clearly needed. Microsoft System Center has been chosen as the new systems management software. System Center offers a single console to manage the life cycle of the devices from the commissioning to the decommissioning. Every step can be easily automated to guarantee fast and repeatable procedures. System Center is also compatible with home office: when the computers are in the IRAM office, they are managed directly, when they connect from outside, they are managed through the IRAM VPN. The commissioning of System Center has been completed at the beginning of the year 2020.

...and Apple computers

The Computer Group has also studied the possibility to manage Apple laptops like Windows ones. Technically, it is possible to install a plugin to manage MacBooks with Microsoft System Center, but it is a complex and not very convenient solution. Therefore, the Computer Group has preferred to opt for JAMF, a mobile device management (MDM) software for Apple devices.

The most spectacular feature is the zero-touch configuration: the serial number of any new MacBook is automatically added to IRAM's Apple Business Account and linked to IRAM's JAMF instance. Should the MacBook be stolen, the IRAM Computer Group can remotely erase and lock the device. Even if it is reset to its factory configuration, it will connect immediately again to the IRAM JAMF instance and it will be locked again.

SCIENCE SOFTWARE ACTIVITIES

Towards new NOEMA observing modes: polarimetry and VLBI

Mid-2020, a first polarimetry measurement was acquired with NOEMA towards a quasar. As the PolyFiX correlator is optimized to cover the widest possible bandwidth, it does not simultaneously measure the cross-polarisation correlations required for full-Stokes polarimetric observations. Instead, the four correlation products are computed in a time sequence. For this purpose, dedicated electronic devices are used to switch the IF signals between the vertical and horizonthal polarization at the entrance of the correlator in a predefined time sequence.

In December 2020, the first Very Long Baseline Interferometry (VLBI) fringes were obtained between NOEMA and several other observatories including the IRAM 30-meter telescope. To do this, the signals from each antenna are phased individually so that they can be added coherently. This means that the 12 NOEMA antennas then behave as a single antenna whose collected area is 12 times the area of a single 15-meter antenna. This is an important step to allow NOEMA to bring its high sensitivity to the benefits of VLBI campaigns, such as the EHT at 1mm or the GMVA at 3mm.

In order to achieve these two milestones, many parts of the software had to be updated. In polarimetry mode, new software had to be written to control the cycle of polarization switches and additional bookkeeping had to be integrated into the raw data format through the correlator. For the VLBI mode, an automated phasing loop was introduced in addition to major changes in the correlator software. The test of both observing modes required changes in the NOEMA raw data format.

New tools for NOEMA: Observation Management System

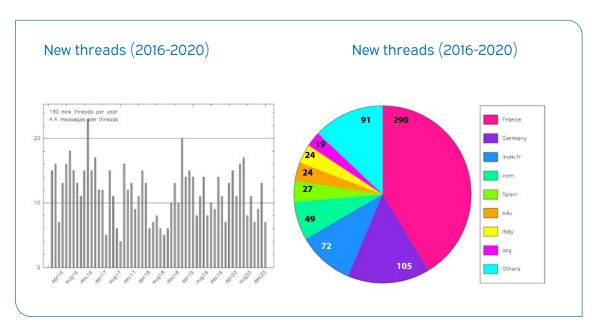
The Observation Management System (OMS) is intended to be a set of independent web-based tools with similar look and feel in order to handle observation projects from proposal submission to distribution of the data to the principal investigators. One of the OMS tools, the Proposal Management System (PMS) has been in operation for many years. In 2020, several major new applications were put in production. First, a web-based sensitivity estimator was introduced to replace the previous tool that required local installation of the GILDAS software. Another major milestone was the release of the Setup Management System (SMS), after several years of development and testing by IRAM astronomers. The SMS allows the principal investigators to easily prepare and validate the details of their NOEMA observing run. In addition to automating many steps such as the selection of calibrators or the setting up of wide-field observations, this tool includes a messaging function and the possibility to check the evolution of the observing setups in order to enable an easy collaboration between the principal investigators and the IRAM staff. Finally, an automated tracking system of the NOEMA status has been implemented. The status of the array (observations, stopped, idle, etc.) is continuously monitored and automatically stored in a database so that unbiased statistics can be calculated.



New functionalities were added to the System Management System (SMS).

Improvements of the data reduction software

In addition to the major projects described above, several improvements of the data reduction software were implemented. The key IRAM 30-meter software packages PAKO, the astronomer interface to the telescope, and PIIC, the calibration software for the NIKA2 continuum camera, were ported to the gfortran compiler in order to facilitate their installation by end users. A novel bandpass calibration scheme was developed and implemented in the NOEMA data calibration pipeline. The routines that handle the calibrator's spectral index for the absolute flux calibration of NOEMA data were refurbished, tested, and implemented in the calibration pipeline. Although workarounds were already available in 2018, shortly after the PolyFiX correlator was put into operation, the new procedures add robustness to the calibration process. The user manual of CLIC, the calibration software for NOEMA, was fully rewritten After a period of performance profiling, various software optimizations increased the speed of the NOEMA pipeline by 75%. Last but not least, user support continued in 2020 with statistics similar to those of the last five years.



140 independent threads (one every 2.6 days), with an average flux of 4.4 emails per threads, were treated in 2020 by the science software team.

IRAM ARC NODE

In the framework of the European ALMA Regional Center (ARC), IRAM is supporting ALMA users, in particular during the data reduction phase. Users can obtain direct help from an IRAM astronomer, in a way similar to the support provided for the NOEMA projects. Travel funding is available for users affiliated to the IRAM funding agencies to visit IRAM and receive face-to-face support. The IRAM ARC node is open to all interested scientists, with an emphasis on the IRAM community (France, Germany, and Spain). One of the goals is to provide to the astronomical community a common support for the IRAM and ALMA facilities, hence maximizing the scientific synergies between the observatories.

In 2020, the Covid-19 pandemic strongly impacted the ALMA operations and led to the cancellation of the Cycle 8 call for proposals. The observatory itself was shut down in March. The "Return to Operations" started on October 1st 2020 with the goal to resume scientific operations in March 2021, after one "blank year". The pandemic also affected the IRAM ARC node, reducing the activity load and preventing any face-to-face support visits. A double remote support solution was implemented, in which both the supported astronomes and the local contact can work from home on the same data reduction session. In 2020, the node supported 80 Cycle 7 projects as contact scientists plus eight Cycle 6 projects carried over. Three face-to-face user support visits were done in Grenoble, and five support sessions were done remotely. An ALMA data reduction training session was organized in the frame of the support of a large program.











Security instructions and evacuation exercise during the Covid pandemic



Administration

The Covid-19 pandemic had an obvious important impact on IRAM operations, on all sites. The Grenoble headquarters were totally closed during three weeks in March-April 2020. Since then, a significant part of the work is done remotely from home, whenever possible. A preliminary analysis of the financial impact due to the Covid-19 pandemic showed a balance between additional costs and savings.

The year 2020 highlighted the importance of the dematerialization process in all administrative activities and allowed the Administration group to reach more flexibility and efficiency. As outlook for 2021, the Administration group will work amongst other things to implement a new purchasing software.

Staff

IRAM employed 121.5 FTEs in 2020 (92.6 in France and 28.9 in Spain). Statistics show a stable proportion of 28% women at IRAM. IRAM continues to have good results in 2020 in terms of the equal pay index between men and women.

IRAM continues to invest in training for the staff, which has a high priority. Unfortunately, due to the Covid-19 pandemic most of the training sessions have been canceled. A side effect was the reduction of the training expenditures in 2020, down -50% in comparison to 2019.

Financial situation of IRAM

Operating

| Income in k€ | Actual 2020 |
|--------------------------|-------------|
| Associates' contribution | 12 859 |
| Other income | 906 |
| Total income | 13 765 |

| Expenses in k€ | Actual 2020 |
|--|-------------|
| Operation (staff costs and operating expenses) | 13 801 |
| Total expenditure | 13 801 |

Investment

| Expenses in k€ | Actual 2020 |
|-------------------|-------------|
| Base Investment | 1 188 |
| Specific projects | 4 075 |
| Total | 5 263 |

IRAM staff list

IRAM Headquarters, Grenoble, France

Direction

SCHUSTER Karl-Friedrich **GUETH Frédéric** DELLA BOSCA Paolo MOREAU Sonja SERLET Murielle

Director Deputy Director

Administration

DELAUNAY Isabelle FERREIRA Dina MARCOUX Stéphane MESSMER Romain PALARIC Laurent SIMON Lauriane SIMONE Jeannine

ZACHER Karin

Head of Administration

Astronomy & Science Support Group

NERI Roberto ANTONELLINI Stefano ARUMUGAM Vinodiran BARDEAU Sébastien BERJAUD Catherine BERTA Stefano **BREMER Michael BROGUIERE** Dominique CASTRO CARRIZO Arancha CHAHINE Layal CHAPILLON Edwige CONTURSI Alessandra CORTZEN Isabella CUNNINGHAM Nichol DE SOUZA MAGALHAES Victor

FEHER Orsolya

HERRERA CONTRERAS Cinthya

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Head of Backend Group

Superconducting Devices Group

DRIESSEN Eduard BARBIER Arnaud BILLON-PIERRON Dominique Head of Superconducting Devices Group

Computer Group

BLANCHET Sébastien CHALAIN Julien DUMONTROTY Patrick MICHAUD Jean-Yves REYGAZA Mickaël ROCHE Jean-Christophe Head of Computer Group

Mechanical Group

LEFRANC Bastien COUTANSON Laurent DANNEEL Jean-Marc JUBARD Vincent LAZARO Gaëtan PASCAUD Victor Head of Mechanical Group

NOEMA, Plateau de Bure, France

GAUTIER Bertrand BOISSIER Jérémie

AZPEITIA Jean-Jacques CASALI Julien CAYOL Alain CHAUDET Patrick COMBE Kevin CONSEIL Yann CONVERS Bruno DAN Michel DI LEONE Cécile LAPEYRE Laurent LE CORRE François-Gildas LEONARDON Sophie MASNADA Lilian MOURIER Yvan OLIVIA Stéphanie RAMBAUD André SALGADO Emmanuel SARRAZIN Jérémy ZANINELLO Pierre

Station Manager Deputy Station Manager

IRAM 30-meter telescope, Granada, Spain

SÁNCHEZ PORTAL Miguel PEÑALVER Juan

AMAYA Sergio BONGIOVANNI Angel Manuel CÓRDOBA Antonió DAMOUR Fréderic EŞPAÑA Gloria GÁLVEZ Gregorio GARCÍA José GARCÍA Verónica GARCÍA RODRIGUEZ Juan Pedro JOHN David KIM Wonju LADJELATE Bilal Smain LARA María LOBATO Enrique MELLADO Pablo MORENO GARCÍA Laura MORENO María MUÑOZ GONZÁLEZ Miguel MYSERLIS Ioannis NAVARRO Santiago PAUBERT Gabriel PEULA Víctor PIERFEDERICI Francesco ROBERTSON II William RODRÍGUEZ MARTÍNEZ Mónica RUIZ Carmen RUIZ Ignacio RŲIZ Manuel SÁNCHEZ Antonio SÁNCHEZ Rosa María SÁNCHEZ Salvador SANTARÉN Juan Luis SANTIAGO Joaquín SERRANO David SIEVERS Albrecht

TORNE Pablo

Station Manager Deputy Station Manager

Telescope schedules

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| 192-16 | The NIKA2 Cosmological Legacy Survey (NIKA2 GT-LP Set1 -N2CLS) | Guilaine Lagache, Alexandre Beelen, Nicolas Ponthieu, Remi Adam, H. Aussel, Matthieu Bethermin, Veronique Buat, Frederic Boone, Emanuele Daddi, David Elbaz, Daizhong Liu, Morgane Cousin, Francois-Xavier Desert, Juan Macias-Perez, Denis Burgarella, Herve Dole, Peter Ade, Philippe Andre, Alain Benoit, Aurelien Bideaud, Nicolas Billot, O. Bourrion, M. Calvo, A. Catalano, Gregoire Coiffard, Barbara Comis, S. Doyle, Carsten Kramer, Samuel Leclercq, Frederic Mayet, A. Monfardini, Francois Pajot, Enzo Pascale, Laurence Perotto, Giampaolo Pisano, Vincent Reveret, Alessia Ritacco, Louis Rodriguez, Charles Romero, Florian Ruppin, Karl-Friedrich Schuster, Albrecht Sievers, Robert Zylka |
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| D04-20 | NIKA2 follow-up of the Swift GRB201024.12 afterglow | Angel Bongiovanni, Bilal Ladjelate, Miguel Sanchez Portal |
| D05-20 | Is phosphine present in Venus' atmosphere? | Raphael Moreno, Carsten Kramer, Emmanuel Lellouch, Roberto Neri, Gabriel Paubert, Miguel Sanchez Portal |
| GMVA-20B-178- MH5 | Ultra-high-resolution Imaging of the Nearest Gamma- ray Emitting NLSy1 1H0323+342 | Kazuhiro Hada, Filippo D'Ammando, Akihiro Doi, Monica Orienti, Kiyoaki Wajima, Gabriele Giovannini, Marcello Giroletti, Bong Won Sohn, Mahito Sasada, Lang Cui, Keiichi Asada, Masanori Nakamura, Valentina Vacca |
| GMVA-20B-179-MI9 | Ultra-deep Imaging the Acceleration and Collimation Zone in the jet of 4C+73.18 | Kunwoo Lee, Jongho Park, Masanori Nakamura, Sascha Trippe |
| GMVA-20B-243- MK16 | Extreme blazars under the GMVA microscope | Matthias Kadler, Eduardo Ros, Marcello Giroletti, Daniela Dorner, Shoko Koyama, Thomas Krichbaum, Rocco Lico, Karl Mannheim, Roopesh Ojha, Julian Sitarek, Luis Wachter |
| GMVA-20B-339- MK18 | The nature of molecular gas in the jet-launching region of young radio galaxy. | Dongjin Kim, Thomas Krichbaum, Biagina Boccardi, Alan L. Roy, Eduardo Ros, Anton Zensus, Anne-Kathrin Baczko, Jae-Young Kim, Violette Impellizzeri, Uwe Bach, Matthias Kadler |

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| D20AA | A novel method to identify the most massive, dusty starburst galaxies and AGN/starburst composites at z>4 with IRAM and LOFAR? | Nicole Nesvadba, Rachel Cochrane, Philip Best, Rohit Kondapally, Kenneth Duncan |
| D20AB | Completing z-GAL: the Comprehensive Redshift Survey of Bright Herschel Galaxies | Pierre Cox, Tom Bakx, Helmut Dannerbauer, Roberto Neri |
| D20AC | The nucleus and inner coma of the great comet C/2020 F3 (NEOWISE) | Jeremie Boissier, Dominique Bockelee-Morvan, Nicolas Biver, Raphael Moreno, Katia Hadraoui, Jacques Crovisier, Martin Cordiner, Stefanie Milam |
| D20AD | Monitoring the Abnormal Behaviour of Betelgeuse | Thavisha Dharmawardena, Karl M. Menten, Ka Tat Wong, Peter Scicluna, Steve Mairs, Albert Zijlstra, Iain McDonald |
| D20AE | The Evolving Pulsar Wind Nebula in SN 2012au | Joseph Bright, Raffaella Margutti, Deanne Coppejans, Giacomo Terreran, Lindsay DeMarchi |
| D20AF | Search for Millimeter Emission From AT2020xnd, an AT2018cow Analog | Anna Ho, Daniel Perley, Yuhan Yao |
| D20AG | Request for Continued Observations of AT2020xnd: Establishing A New Class of Millimeter-Bright Transients | Anna Ho, Yuhan Yao, Daniel Perley, David Kaplan, Dougal Dobie, Andrew O'Brien |
| E19AA | Confirming the detection of the [CII] line in the $z=7.12$ galaxy GN-108036 | Rodrigo Herrera-Camus, Linda Tacconi, Natascha Förster Schreiber, Reinhard Genzel, Dieter Lutz, Alessandra Contursi |
| E19AB | Water deuteration in the Class 0 protostar CepE-mm | Bertrand Lefloch, Juan Ospina-Zamudio, Cecile Favre, Jaime Pineda, Dominique Segura-Cox, Ana Lopez-Sepulcre, Claudio Codella, Paola Caselli, Cecilia Ceccarelli |
| E19AD | Gas and Dust Properties in a Red Quasar Firstly Discovered at z>7 | Seiji Fujimoto, Gabriel Brammer, Johan Fynbo, Georgios Magdis, Sune Toft, Charles Steinhardt, Francesco Valentino, Isabella Cortzen |
| L19MA | Galactic star formation MIOP: From clouds to cores | Karl M. Menten, Friedrich Wyrowski, Henrik Beuther, Gisela Ortiz Leon, Thanh Dat Hoang, Antonio Hernandez-Gomez, Sumeyye Suri, Caroline Gieser, Nicola Schneider, Timea Csengeri, Sylvain Bontemps, Frederique Motte, Nichol Cunningham, Jan Martin Winters, Ka Tat Wong, Wonju Kim |
| L19MB | Redshift determination of high-redshift Herschel lensed galaxies | Alain Omont, Roberto Neri, Alexandre Beelen, Steve Eales, Tom Bakx, R. Gavazzi, Simon Dye, Rob Ivison, Ismael Perez-Fournon, Ivan Oteo, Helmut Dannerbauer, Frank Bertoldi, Matthew Lehnert, Chentao Yang, Zhiyu Zhang, Gianfranco De Zotti, Dominik A. Riechers, David Clements, Lucia Marchetti, Joshua Greenslade, Melanie Krips, Catherine Vlahakis, Michal J. Michalowski, Andrew Baker, Paul van der Werf, Pierre Cox |
| L19ME | From Protostars to Planet-Forming Disks | Paola Caselli, Thomas Henning, Jaime Pineda, Dominique Segura-Cox, Dmitry Semenov, Bo Zhao, Mario Tafalla, Grigorii Smirnov-Pinchukov, Cecilia Ceccarelli, Edwige Chapillon, Nichol Cunningham, Anne Dutrey, Stephane Guilloteau, Izaskun Jimenez-Serra, Ana Lopez-Sepulcre, Sebastian Marino, Maria Maureira, Roberto Neri, Vincent Pietu, Asuncion Fuente |
| L19MC | Search for molecular absorption in AGNs. | Dongjin Kim, Rainer Mauersberger, Thomas Krichbaum, Biagina Boccardi, Anton Zensus, Michael Bremer, Christian Henkel, Francoise Combes |
| L19MD | NOEMA3D: a Comprehensive Census of the Molecular Gas Distribution and Kinematics of Massive Main- Sequence Star Forming Galaxies at the Peak and Winding Down of Galaxy Formation Activity | Reinhard Genzel, Roberto Neri, Linda Tacconi, Natascha Förster Schreiber, Dieter Lutz, Karl-Friedrich Schuster, Alessandra Contursi, Melanie Krips, Stefano Berta, Fabian Walter, Axel Weiss, T. Naab, Richard Davies, Minju Lee, Sedona Price, Thomas Taro Shimizu, Eckhard Sturm, Hannah Uebler, Francoise Combes, David Elbaz, Pierre Cox, Santiago Garcia-Burillo, Antonio Usero, Alberto D. Bolatto, Monica Rubio, Tadayuki Kodama, Rodrigo Herrera-Camus, Ken-Ichi Tadaki, Alvio Renzini, Amiel Sternberg, Andreas Burkert, Avishai Dekel, S. Wuyts, Cinthya Herrera Contreras |
| M18AB | A Comprehensive NOEMA Redshift Survey of the Brightest Herschel Galaxies | Pierre Cox, Tom Bakx, Helmut Dannerbauer, Roberto Neri, Alain Omont, Steve Eales, Rob Ivison, Matthew Lehnert, R. Gavazzi, Stephen Serjeant, Lucia Marchetti, Mattia Negrello, Simon Dye, Dominik A. Riechers, Melanie Krips, Asantha Cooray, Guilaine Lagache, Ismael Perez-Fournon, Ivan Oteo, David Hughes, Hugo Messias, Veronique Buat, Andrew Baker, Catherine Vlahakis, Paul van der Werf, Loretta Dunne, Chentao Yang, Stefano Berta, Alexandre Beelen, Axel Weiss, Cinthya Herrera Contreras |
| M19AA | The First Cloud-by-Cloud Dense Gas Map of an External Galaxy | Eva Schinnerer, Frank Bigiel, Jérôme Pety, Antonio Usero, Adam Leroy, Annie Hughes, Miguel Querejeta, Cinthya Herrera Contreras, Ashley Thomas Barnes, Ilin Lazar, Cosima Eibensteiner, Jakob den Brok, Daizhong Liu, Johannes Puschnig, Toshiki Saito, Axel García-Rodríguez, Sharon Meidt, Eric Emsellem, Jonathan Henshaw, Melanie Chevance, Diederik Kruijssen, Alex Hygate, Andreas Schruba, Maria Jesus Jimenez Donaire |
| S19AW | Ionization Structure of Protoplanetary Disks: Is there a connection with turbulence? | Grigorii Smirnov-Pinchukov, Richard Teague, Dmitry Semenov, Stephane Guilloteau, Thomas Henning, Anne Dutrey, Mario Flock |
| S19BB | Testing time-dependent chemistry in the expanding CSE IRC+10216 | Michel Guélin, Jose Cernicharo, Marcelino Agundez, Jan Martin Winters, Carl Gottlieb, Guillermo Quintana-Lacaci |
| S19BM | CO(1-0) in a Candidate Accretion-Induced Starburst in a Local Dwarf Galaxy | Monica Rubio, Bruce Elmegreen, Cinthya Herrera Contreras, Debra Elmegreen, Jorge Sanchez Almeida, Casiana Munoz-Tunon, Elias Brinks, Deidre Hunter, Leslie Hunt, Juan Cortes |

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| S19BP | Mapping molecular gas in one of the most extended Halpha nebulae in X-ray cool cores | Ming Sun, Alastair Edge, Francoise Combes, William R. Forman, Hao Chen |
| S19BV | Probing the molecular gas content of galaxies in an over-dense group at z~0.7: a test case for environmental quenching | Thierry Contini, Jonathan Freundlich, Benoit Epinat, Philippe Salomé, Avishai Dekel, G. Soucail, Léo Michel-Dansac, Sandro Tacchella, Leindert Boogaard |
| S19BY | CO survey of the most strongly lensed galaxies | Johan Richard, Miroslava Dessauges-Zavadsky, Francoise Combes, Frederic Boone, Daniel Schaerer, Vera Patricio, Wiphu Rujopakarn, John Chisholm |
| S19CC | At the end of the Cosmic Noon: Molecular gas census in intermediate-redshift lensed quasars | Matus Rybak, Jacqueline Hodge, Paul van der Werf, Gabriela Calistro Rivera, Hannah Stacey, John Mckean |
| S19DM | Unraveling the powering mechanism of the cold ISM in z>6 quasars | Roberto Decarli, Bram Venemans, Fabian Walter, Mladen Novak, Eduardo Banados, Dominik A. Riechers, Axel Weiss, Feige Wang, Jinyi Yang, Xiaohui Fan, Jianan Li, Ran Wang, Antonio Pensabene |
| S20AA | Gas and grain interactions: unraveling the methanol story with JWST and NOEMA | Valentine Wakelam, Pierre Gratier, Daniel Harsono, Jacqueline Keane, Adwin Boogert, Jennifer Noble, E. Dartois, Marin Chabot |
| S20AB | NOEMA-SPONGE: the role of neutral hydrogen in diffuse interstellar chemistry | Daniel Rybarczyk, Snezana Stanimirovic, Claire Murray, Ellen Zweibel, Shuichiro Inutsuka, John Dickey, Antoine Gusdorf, Anne Klitsch, Celine Peroux, John Wu, Martin Zwaan |
| S20AD | What is the true evolutionary state of the First Hydrostatic Core candidate CB 17 MMS? | Maria Maureira, Paola Caselli, Hector Arce, Jaime Pineda, Stephanie Spear, Dominique Segura-Cox |
| S20AE | Chemical and dynamical evolution of starless cores | David Navarro Almaida, Asuncion Fuente, Evelyne Roueff, Paola Caselli, Carsten Kramer, Mario Tafalla, Santiago Garcia-Burillo, Rafael Bachiller, Maryvonne Gerin, Stephanie Cazaux, Valentine Wakelam, Nuria Marcelino, Roberto Neri, Marcelino Agundez, Izaskun Jimenez-Serra, Tomas Alonso-Albi, Jaime Pineda, Silvia Spezzano, Sandra Trevino-Morales, Valerio Lattanzi, Jean-Christophe Loison, Ana Chacón-Tanarro, Pablo Riviere-Marichalar, Marina Rodriguez Baras, Romane Le Gal, Gisela Esplugues |
| S20AH | Distinguishing between envelope and embedded disk chemistry of Class I YSOs | Romane Le Gal, Karin Öberg, Ana Lopez-Sepulcre, Charlotte Vastel, Jane Huang, Francois Menard, Charles Law, Bertrand Lefloch, Cecilia Ceccarelli, Cecile Favre, Eleonora Bianchi, Clement Baruteau, Asuncion Fuente, Pablo Riviere-Marichalar, Anaelle Maury, David Navarro Almaida, Edwin A. Bergin, Emmanuel Caux |
| S20AI | Evolution of a massive clump at the intersection of colliding filaments | Julien Montillaud, Cecile Favre, Charlotte Vastel, Mika Juvela, Isabelle Ristorcelli, Tie Liu, David Cornu, Rebeka Bögner |
| S20AJ | Hunting for the coldest cores in Cygnus-X | Keping Qiu, Yue Cao, Bo Hu, Junhao Liu |
| S20AL | Disruption or Assembly? The Origin of the Widespread SiO Emission in a Massive IRDC | Wonju Kim, Gary Fuller, Nichol Cunningham, Nicolas Peretto, Fabien Louvet, Di Li, Jinjin Xie, Bethany Jones, Ken'ichi Tatematsu, Adam Avison, James Urquhart, Ka Tat Wong |
| S20AM | Marking the End of Evolution: The association of class II methanol masers with high-mass star formation | Nichol Cunningham, Gary Fuller, Adam Avison, Alessio Traficante, Bethany Jones, Shari Breen, Wonju Kim |
| S20AR | CO as a Tracer of Pebble Growth and Drift in Protoplantary disks | Ke Zhang, Edwin A. Bergin, Arthur Bosman |
| S20AT | A tale of two masses: HD and CO | Ewine F. Van Dishoeck, Alice Booth, Margot Leemker, Melissa McClure, Leon Trapman |
| S20AX | Dust properties, temperatures, and the HCN X-factor of individual Giant Molecular Clouds in the Andromeda Galaxy, using unique NOEMA/SMA synergies | Jan Forbrich, Charles Lada, Christopher Faesi, Glen Petitpas, Sebastien Viaene, Jérôme Pety, Jonathan Toomey |
| S20AZ | First cloud-scale measurement of $N_{\rm g} H^{\dagger}$ in a normal starforming galaxy | Axel García-Rodríguez, Antonio Usero, Maria Jesus Jimenez Donaire, Frank Bigiel, Ivana Beslic, Johannes Puschnig, Melanie Chevance, Diederik Kruijssen, Ashley Thomas Barnes, Eva Schinnerer, Toshiki Saito, Christopher Faesi, Daizhong Liu, Andreas Schruba, Miguel Querejeta, Adam Leroy, Cinthya Herrera Contreras, Jérôme Pety, Simon Glover |
| S20BB | Searching for molecular gas in HI-rich red spirals | Jianhang Chen, Yong Shi, Zhiyu Zhang |
| S20BD | Mapping two Relic Compact Elliptical Galaxies | Francoise Combes, Philippe Salomé, Valeria Olivares |
| S20BH | Feedback from low luminosity radio AGN: Strong jet-ISM interaction in B2 0258+35 | Suma Murthy, Pierre Guillard, Raffaella Morganti, Tom Oosterloo, Dipanjan Mukherjee |
| S20BJ | Probing Merger-driven Gas Inflows in Dual-AGNs | Meicun Hou, Zhiyuan Li, Zongnan Li, Xin Liu |
| S20BQ | A systematic search for ultra-bright strongly lensed galaxies in the Planck all-sky survey | Mattia Negrello, Gianfranco De Zotti, Matteo Bonato, Andrea Enia, Pierre Cox, Stefano Berta, Tiziana Trombetti, Carlo Burigana, Lucia Marchetti, Roberto Neri, Amvrosiadis Aristeidis, Alain Omont, Matthew Smith |
| S20BT | Tracing starburst HyLIRGs in host galaxies of high-z QSOs | Y. Sophia Dai, Alain Omont, Pierre Cox, Xue-Bing Wu, Rob Ivison, Chentao Yang, Roberto Neri, Belinda Wilkes, Jacqueline Bergeron, Melanie Krips |
| S20BV | How thermalized is the molecular ISM at z~2? | Melanie Kaasinen, Daizhong Liu, Fabian Walter, Leindert Boogaard, Manuel Aravena, Ian Smail, Chelsea Sharon, Dominik A. Riechers, Gergo Popping |
| S20BY | Gas Properties in the Early Universe: collisional excitation or IR pumping in the Cloverleaf Quasar? | Michel Guélin, Carsten Kramer |
| S20BZ | Looking down the barrel through a multi-phase quasar outflow at high redshift with VLT and NOEMA | Pasquier Noterdaeme, Raghunathan Srianand, Francoise Combes, Sergei Balashev, Jens-Kristian Krogager, Peter Laursen, Neeraj Gupta |
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| S20CA | Redshift Scans of Submillimeter Galaxies in the Hubble Frontier Fields: A Pilot Program | Logan Jones, Amy Barger, Lennox Cowie |
| S20CD | Unveiling the host galaxies of the most luminous quasars at $z\sim3.5$ | Jan-Torge Schindler, Bram Venemans, Fabian Walter, Eduardo Banados, Marcel Neeleman |
| S20CE | Coming out from the curtain: Molecular gas condition in an optical/near-IR dark sub-mm bright AGN host | Zhiyu Zhang, Chentao Yang, Chao-Wei Tsai, Natsuki H. Hayatsu, Ping Zhou, Rob Ivison |
| S20CK | Exploiting a snapshot survey of the 3,083 reddest Herschel sources to reveal distant protoclusters | Vinodiran Arumugam, Melanie Krips (NOEMA coordinator), Rob Ivison, Helmut Dannerbauer, Loretta Dunne, Malcolm Bremer, Tom Bakx, Seiji Fujimoto, Geraint Lewis, Jianhang Chen |
| S20CL | A Band 2 Follow-up to Candidate Protoclusters in NIKA2 Science Verification Field | Longji Bing, Alexandre Beelen, Guilaine Lagache, Matthieu Bethermin, Rémi Adam, Nicolas Ponthieu, Juan Macias-Perez, Alain Omont, Roberto Neri, Stefano Berta |
| S20CP | How fast could the Universe form a massive galaxy? | Pablo Pérez-González, Jesus Martin-Pintado, Helmut Dannerbauer, Giovanni Fazio, Guillermo Barro, Miguel Pereira Santaella, Belén Alcalde Pampliega, Ángela García Argumánez |
| S20CS | NOEMA line scans toward z>5 candidate AzTEC-ALMA SMGs | Soh Ikarashi, Rob Ivison, Karina Caputi, Kotaro Kohno |
| S20CV | First insights into the ISM of the most distant radio galaxy at z=5.72 | Bram Venemans, Fabian Walter, Marcel Neeleman, Aayush Saxena, Eduardo Banados Chris L. Carilli |
| S20CW | Probing the CO spectral line energy distributions (SLEDs) in two luminous quasars at z~6 | Jianan Li, Karl M. Menten, Frank Bertoldi, Fabian Walter, Roberto Neri, Alain Omont, Pierre Cox, Xiaohui Fan, Michael A. Strauss, Dominik A. Riechers, Ran Wang, Jeff Wagg Desika Narayanan, Roberto Decarli, Eduardo Banados, Yali Shao, Yuanqi Liu, Qiong Li, Bram Venemans, Yu Gao, Chris L. Carilli |
| S20CY | The host galaxies of the most distant radio-loud quasars at z>6 | Eduardo Banados, Yana Khusanova, Sofia Rojas, Chiara Mazzucchelli, Bram Venemans, Roberto Decarli, Fabian Walter |
| S20DA | Dense Gas Content of Star-Forming Cores in the Best- Studied AGN Host and Starburst at z~6.5 | Dominik A. Riechers, Fabian Walter, Roberto Decarli, Axel Weiss, Pierre Cox, Roberto Neri |
| S20DB | Probing "Cosmic Dawn" : The physical properties of a z=8.78 galaxy | Nicolas Laporte, Frederic Boone, Richard Ellis, Guido Roberts-Borsani, Romain Meyer, Roberto Maiolino, Gareth Jones |
| W18BY | 50pc imaging of the molecular outflow in M82 | Nico Krieger, Fabian Walter, Axel Weiss, Adam Leroy, Alberto D. Bolatto, Laura Zschaechner, Sylvain Veilleux |
| W18CS | Probing the molecular gas content in galaxies with gas flows: a test case for self-regulated star formation models | Nicolas Bouche, Thierry Contini, Jonathan Freundlich, Léo Michel-Dansac, llane Schroetter, Johan Richard, Johannes Zabl |
| W19AA | Protostellar Chemistry and the Carbon Content of Terrestrial Worlds | Edwin A. Bergin, Jes Jorgensen, Ke Zhang, Thomas Rice, Merel van 't Hoff |
| W19AG | Envelope to Disk: The Composition of Accretion | Dominique Segura-Cox, Jaime Pineda, Paola Caselli, Anika Schmiedeke, Nichol Cunningham, Maria Maureira |
| W19AH | lonization: explaining the split between big and small Protostellar Disks | Maria Maureira, Jaime Pineda, Dominique Segura-Cox, Paola Caselli, Bo Zhao |
| W19AI | Are disk asymmetries driven by large scale accretion flows? | Jaime Pineda, Anika Schmiedeke, Paola Caselli, Dominique Segura-Cox, Bo Zhao, Nichol Cunningham, Roberto Neri, Maria Maureira |
| W19AL | 2MASS J0233+6156: a newly discovered eruptive star | Orsolya Feher, Agnes Kospal, Carlos Contreras Pena, Fernando Cruz-Saenz de Miera |
| W19AM | Caught in the act: Varying infall towards a periodically brightening protostar | Hendrik Linz, Bringfried Stecklum, Henrik Beuther, Alessio Caratti O Garatti, Marian Szymczak |
| W19AN | Catching the missing: NOEMA survey of massive cold cores in Cygnus-X | Keping Qiu, Yue Cao, Yuwei Wang |
| W19AP | Investigation for mass growth of massive young stellar objects by the feedback of companion gas clump | Chen Xi |
| W19AQ | AU-scale structure in the diffuse ISM: shocks and multiphase nature | Daniel Rybarczyk, Snezana Stanimirovic, Ellen Zweibel, John Dickey, Shuichiro Inutsuka, Antoine Gusdorf |
| W19AR | The protostellar and interstellar content of the IC443 SNR | Antoine Gusdorf, Pierre Dell'Ova, Maryvonne Gerin, Fabien Louvet, Denise Riquelme, Marco Padovani, Martin Houde, Bilal Ladjelate |
| W19AU | The Molecular Content of Small Protoplanetary Disks | Vincent Pietu, Jean-Paul Melisse, Anne Dutrey, Stephane Guilloteau, Edwige Chapillon |
| WI9AV | GG Tau A: a 3mm Large spectral Survey in the densest binary TTauri disk | Anne Dutrey, Edwige Chapillon, Stephane Guilloteau, Vincent Pietu, Ya-Wen Tang, Jeffrey Bary, Audrey Coutens, Diep Pham Ngoc, Liton Majumdar, Emmanuel Di Folco, Otoniel Denis-Alpizar, Jean-Paul Melisse, Tracy Beck, Thi Phuong Nguyen |
| W19AW | Searching for Cold Circumbinary Disks Around High Mass X-Ray Binaries | Kevin Fogarty, Dimitri Mawet, Julie Hlavacek-Larrondo, Jack Steiner, Andy C. Fabian, Dom Walton, Lauren Weiss, David Lafreniere, Myriam Prasow-Vâmond |
| W19AX | Streams, disk and chemistry in the S-type AGB star RS Cnc | Thibaut Le Bertre, Jan Martin Winters, Ka Tat Wong, Do Thi Hoai, Pham Thi Tuyet Nhung, Pham Tuan-Anh, Pierre Lesaffre, Diep Pham Ngoc, Pierre Darriulat |
| W19BJ | CO observation of a star-forming SO galaxy PGC 34107 | Xue Ge, Qiusheng Gu, Zhiyu Zhang, Rubén García-Benito, Mengyuan Xiao, Zhengyi Chen |
| W19BK | Mapping Molecular Gas in Shell Galaxies | Brisa Mancillas, Francoise Combes, Pierre-Alain Duc |

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| W19BL | SFE distributions in interacting galaxies Difference between S+E and S+S pairs | Cong Kevin Xu, Ute Lisenfeld, Yu Gao, Min Yun |
| W19BN | Search for a cool accretion disk around M31 black hole | Anne-Laure Melchior, Francoise Combes, Igor Chilingarian, Ivan Katkov, Daniel Maschmann |
| W19BR | A CO survey and molecular gas scaling relations of the SUBWAYS AGN sample | Chiara Feruglio, Marcella Brusa, Stefano Bianchi, Francesco Tombesi, Smita Mathur, Fabio La Franca, Cristian Vignali, Yair Krongold, Raffaella Morganti, Enrico Piconcelli, Anna Lia Longinotti, Fabrizio Fiore, Michele Perna, Gerard Kriss, Ehud Behar, Massimo Gaspari, Alessandro Marconi |
| W19BS | Molecular gas content in post starburst E+A galaxies with massive AGN-driven winds | Richard Davies, Dieter Lutz, H. Netzer, Dalya Baron |
| W19BV | Unveiling the MS to SB transition with a sample of intermediate redshift ULIRGs | Santiago Garcia-Burillo, Dimitra Rigopoulou, Georgios Magdis, Miguel Pereira Santaella, Francoise Combes, Ismael García-Bernete, Almudena Alonso-Herrero, Axel Weiss |
| W19BZ | Short depletion scales in cluster LIRGs: the LoCuSS sample | Pascale Jablonka, Francoise Combes, Gianluca Castignani, Chris Haines, Melanie Krips, Monique Arnaud |
| W19CF | Investigating the molecular component of fast multiphase outflows into the CGM | Christy Tremonti, James Geach, Amanda Kepley, Gregory Rudnick, John Moustakas, Alison Coil, Ryan Hickox, Aleks Diamond-Stanic, David Rupke, Paul Sell |
| W19CJ | Characterizing Molecular Gas in Quenching Galaxies at z>1 | Sirio Belli, Alessandra Contursi, Natascha Förster Schreiber, Dieter Lutz, Linda Tacconi Rodrigo Herrera-Camus, Hannah Uebler, Thomas Taro Shimizu, Sedona Price, Rebecca Davies, Erica Nelson, Minju Lee, Jinyi Shangguan |
| W19CL | A titanic wind destroying the interstellar medium of a massive starburst galaxy at z=1.4? | Annagrazia Puglisi, Emanuele Daddi, Marcella Brusa, Chiara Mancini, Giulia Rodighiero, Ivan Delvecchio, Andrea Enia |
| WI9CR | Completion: A survey of molecular gas in the Hubble Deep Field North (HDF-N) | Fabian Walter, Roberto Decarli, Mladen Novak, Dominik A. Riechers, Manuel Aravena, Axel Weiss, Gergo Popping, Emanuele Daddi, Daizhong Liu, Mark T. Sargent, Frank Bertoldi, Benjamin Magnelli, Roberto Neri, Jérôme Pety, Pierre Cox, Elisabete Da Cunha, Pascal Oesch, Rychard Bouwens, Richard Ellis, Dan Stark |
| W19CS | Resolving the [CI] and Mid-J CO Line Emission from Giant Molecular Clouds in Strongly Lensed Starbursts at $z=2-3$ | Kevin Harrington, Axel Weiss, Benjamin Magnelli, Eric Faustino Jiménez-Andrade, Min Yun, Frank Bertoldi, David Frayer, Patrick Kamieneski, Q. Daniel Wang, Amit Vishwas, T.K. Daisy Leung, Nichol Cunningham |
| W19CW | Probing the dense gas star formation law in intense dusty starbursts at z~2-4 | Raoul Canameras, Nicole Nesvadba, Chentao Yang, Sabine König, Alexandre Beelen, Ryley Hill, Ruediger Kneissl, Emeric Le Floc'h, Sangeeta Malhotra, Alain Omont, Douglas Scott |
| W19DA | Gas Properties in the Early Universe from the study of molecular lines: 1.5 mm observations of high excitation lines in the Cloverleaf Quasar | Michel Guélin, Chentao Yang, Carsten Kramer, Robert Zylka, Alain Omont |
| W19DB | Probing the excess gas origin in QSO2343+12 | Scott Chapman, Axel Weiss, Axel Weiss |
| W19DD | Rise of the Titans: Molecular Gas Excitation in Hyper- Luminous Starbursts at z=3-6 | Dominik A. Riechers, Ismael Perez-Fournon, Rob Ivison, Roberto Neri, Alain Omont, Pierre Cox, David Clements |
| W19DE | Exploring the Nature of the Broad Line Bright Herschel Galaxy HerBS-89a at z=2.95 | Stefano Berta, Pierre Cox, Roberto Neri, Alain Omont, Alexandre Beelen, Tom Bakx, Matthew Lehnert, Andrew Baker, Veronique Buat, Helmut Dannerbauer, Loretta Dunne, Simon Dye, Steve Eales, R. Gavazzi, Andrew Harris, Cinthya Herrera Contreras, Rob Ivison, Shuowen Jin, Melanie Krips, Guilaine Lagache, Lucia Marchetti, Hugo Messias, Mattia Negrello, Ismael Perez-Fournon, Stephen Serjeant, Sheona Urquhart, Catherine Vlahakis, Axel Weiss, Paul van der Werf, Chentao Yang, Anthony Young, Dominik A. Riechers |
| W19DF | $\mbox{CH}^{+}\mbox{(I-O)}$ in high-redshift starburst galaxies : probes of massive turbulent reservoirs | Alba Vidal Garcia, Edith Falgarone, Benjamin Godard, Roberto Neri, Paola Andreani, Rob Ivison, Chentao Yang, Martin Zwaan, Alain Omont, Fabian Walter, Edwin A. Bergin, Nick Indriolo |
| W19DG | Redshift Scans of Bright Submm Galaxies in a Candidate Proto-cluster in the GOODS-N Field | Logan Jones, Amy Barger, Lennox Cowie |
| W19DI | A critical test of the nature of cold dust in submillimeter bright galaxies at $z{\sim}4$ | Emanuele Daddi, Shuowen Jin, Georgios Magdis, Daizhong Liu, Eva Schinnerer, Francesco Valentino, Yu Gao, Antonello Calabro', Padelis Papadopoulos, Isabella Cortzen |
| W19DN | Measuring for the first time the massive-end of the dust mass function of galaxies at $z{\sim}4$ | Benjamin Magnelli, Eva Schinnerer, Mark T. Sargent, Eric Faustino Jiménez-Andrade, Caitlin Casey, Jorge Zavala, Manuel Aravena, Dominik A. Riechers, Daizhong Liu, Frank Bertoldi, Alexander Karim, Tsan-Ming Wang |
| W19DO | Capitalising on an ALMA snapshot survey of the 3,083 reddest Herschel sources | Vinodiran Arumugam, Rob Ivison, Melanie Krips, Loretta Dunne, Helmut Dannerbauer |
| W19DQ | Final Piece to The Most Puzzling z>4 SMG in GOODS- North, New Chapter for A Blind Line Emitter Field? | Daizhong Liu, Emanuele Daddi, Eva Schinnerer |
| W19DS | Exploration of a Small Angular Separation Quasar/SMG Pair | Marcel Neeleman, Nissim Kanekar, Fabian Walter |
| W19DU | The Formation and Growth of Supermassive Black Holes in the Early Universe | Anna-Christina Eilers, Bram Venemans, Fabian Walter, Joseph Hennawi, Roberto Decarli, Eduardo Banados, Feige Wang, Chiara Mazzucchelli, Frederick Davies |

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| W19EB | Witnessing the formation of the first large-scale structures in the universe | Roberto Decarli, R. Gilli, M. Mignoli, Cristian Vignali, Eros Vanzella, Marcella Brusa, Andrea Comastri, Antonio Pensabene, Riccardo Nanni, Alessandro Peca, Nico Cappellutti, Barbara Balmaverde |
| W20AA | The intriguing mm-spectrum of lapetus | Emmanuel Lellouch, Alice Le Gall, Lea E Bonnefoy, Cedric Leyrat, Raphael Moreno, Jeremie Boissier |
| W20AF | Connecting scales in protostellar disk formation | Edwin A. Bergin, Merel van 't Hoff, Felipe Alarcon, Arthur Bosman, Jane Huang, Kamber Schwarz, Jenny Calahan, John Tobin |
| W20AQ | The OMC-2 FIR6c outflow: A new astrochemical laboratory | Mathilde Bouvier, Ana Lopez-Sepulcre, Cecilia Ceccarelli, Nadia Balucani, Claudio Codella, Marta De Simone, Joan Enrique-Romero, Bertrand Lefloch, Pedro Ruben Rivera-Ortiz, Nami Sakai, Andre Schutzer, Fanny Vazart, Satoshi Yamamoto |
| W20BB | Investigating the deuterium fraction in the diffuse and translucent clouds through sensitive absorption observations | Luo Gan, Zhiyu Zhang, Siyi Feng, Thomas Bisbas |
| W20BC | The Molecular Content of Small Protoplanetary Disks | Vincent Pietu, Jean-Paul Melisse, Anne Dutrey, Stephane Guilloteau, Edwige Chapillon, Emmanuel Di Folco |
| W20BK | Determining the cause of the extremely unusual behaviour of Betelgeuse with NOEMA | Thavisha Dharmawardena, Karl M. Menten, Ka Tat Wong, lain McDonald, Albert Zijlstra, Steve Mairs, Peter Scicluna, Miguel Montargès, Anita M. S. Richards, Fabrice Herpin |
| W20CM | At the end of the Cosmic Noon: Completing the molecular gas census in lensed quasars at z=1.0-1.5 | Matus Rybak, Paul van der Werf, Jacqueline Hodge, Hannah Stacey, Gabriela Calistro Rivera, Marta Frias Castillo |
| W20CW | The molecular gas properties of high-z galaxies as traced by [Ci] and CO | Isabella Cortzen, Roberto Neri, Pierre Cox, Stefano Berta, Anthony Young, Alain Omont, Flora Stanley, Tom Bakx, Helmut Dannerbauer, Alexandre Beelen, Dominik A. Riechers, Axel Weiss, Andrew Baker, Steve Eales, Rob Ivison, Shuowen Jin, R. Gavazzi, Stephen Serjeant, Lucia Marchetti, Mattia Negrello, Matthew Lehnert, Melanie Krips, Ismael Perez-Fournon, Asantha Cooray, Chentao Yang, Sheona Urquhart, Hugo Messias, Veronique Buat, Catherine Vlahakis, Paul van der Werf, Loretta Dunne, Andrew Harris |
| W20DB | Molecular gas across the circum-galactic medium of Enormous Lyα Nebulae | Qiong Li, Jiangtao Li, Ran Wang, Jianan Li |
| W20DC | A multi-phase, multi-scale, multi-wavelength investigation of quasar feedback at high redshift. NOEMA: The dense molecular component. | Pasquier Noterdaeme, Sergei Balashev, Jens-Kristian Krogager, Francoise Combes, Neeraj Gupta, Raghunathan Srianand |
| W20DD | A Survey of CO(3-2) in the 20Mpc environment of HS1549+19, the most overdense protocluster known at z>2 | Scott Chapman, Frank Bertoldi, lan Smail, Chuck Steidel, Yuichi Matsuda, Manuel Aravena, Nikolaus Sulzenauer |
| W20DX | NOEMA line scans toward z>5 candidate ALMA-AzTEC SMGs (Continue) | Soh Ikarashi, Rob Ivison, Karina Caputi, Kotaro Kohno |
| W20EO | A Vigorously Star-forming Red Quasar Firstly Discovered at z>7 | Seiji Fujimoto, Gabriel Brammer, Sune Toft, Georgios Magdis, Francesco Valentino, Charles Steinhardt, Isabella Cortzen, Luis Colina, Rui Marques-Chaves, Thomas R. Greve, M. Vestergaard, Peter Jakobsen, Vasily Kokorev |

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| 2584 | Molecular Gas in Distant Brightest Cluster Galaxies | Castignani, G; Combes, F.; Salomé, P.; Freundlich, J. | 2020, A&A, 635, A32 |
| 2585 | Molecular Gas in CLASH Brightest Cluster Galaxies at z \sim 0.2 - 0.9 | Castignani, G.; Pandey-Pommier, M.; Hamer, S. L.; Combes, F.; Salomé, P.; Freundlich, J.; Jablonka, P. | 2020, A&A, 640, A65 |
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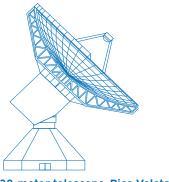
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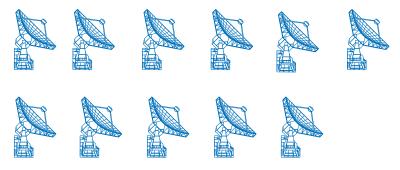
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30-meter telescope, Pico Veleta



11 x 15-meter Interferometer, NOEMA

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