Counting quasiparticles: generation-recombination noise in microwave resonators

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Detection by pair breaking

- 1. Most electrons are paired in equilibrium
- 2. A photon breaks a pair
- 3. Resistance increases
- 4. Detection with microwave resonator





Quasiparticle number fluctuations

In equilibrium: generation and recombination balanced





Number of excitations fluctuates: fundamental noise!



Quasiparticle number fluctuations in equilibrium



$$N_{qp}\tau = 2\tau_0 N_0 V \frac{(kT_c)^3}{(2\Delta)^2}$$

Noise level constant as a function of temperature!!

Wilson and Prober, PRB 69, 094524 (2004)

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Measurement: alumumium resonator

- 50 nm Al film, sputtered on sapphire
- T_c=1.1 K
- Half wave CPW resonators
- Resonant at 6.6 GHz, Q~40k, Qi~160k



Calibration: subtract system noise

- System noise measured off-resonance
- System noise subtraction





Measurement: Noise at different temperatures



Clear sign of quasiparticle number fluctuations!! Fundamental noise reached!



Quasiparticle lifetime



 $4N\tau$ S

2 methods: consistent lifetime But it saturates at low Temperature



Noise is a signal!

Counting quasiparticles with microwave resonators



Noise spectrum -> number of quasiparticles

Lifetime limited by excess quasiparticles



Fundamental sensitivity for KIDs



$$NEP_{A}(\omega) = \sqrt{S_{A}} \left(\frac{\eta\tau}{\Delta} \frac{dA}{dN_{qp}}\right)^{-1} \sqrt{1 - \omega^{2}\tau^{2}}$$



Fundamental sensitivity for KIDs



NEP_{q-r} behaves as theoretically expected

Low temperature NEP limited by excess quasiparticles SRON Microresonator Workshop 2011

Sources of quasiparticles

- Stray light creates quasiparticles -> Light-tight setup No photon noise observed
- (Cosmic ray) hits -> filter peaks out of noise analysis



Sources of quasiparticles

- The number of quasiparticles and lifetime are microwave readout power dependent / limited
- Means that we are sure other effects do not limit Nqp!
- Small range: one side bifurcation, other side amplifier noise



Resonator noise limited by intrinsic generation-recombination noise

Noise is a signal: powerful tool to count quasiparticles and determine their lifetime

Both saturate consistently at low temperature - due to the readout power



Phys. Rev. Lett. 106, 167004, 2011

This opens the door to fundamentally limited KID - detectors

Microwave absorption changes the distribution function



- Quasiparticle distribution function can change drastically
- Resulting distribution depends on ratios of scattering/recombination rates
- Number of quasiparticles and their lifetime energy dependent
 - Microwave absorption can also result in effective cooling

Full treatment will require solving kinetic equations

- B.I. Ivlev, S.G. Lisitsyn and G.M. Eliashberg, 'Nonequilibrium excitations in superconductors in high-frequency fields', J. Low Temp. Phys. 10, p449 (1973)
- J.-J. Chang and D.J. Scalapino, 'Kinetic-equation approach to nonequilibrium superconductivity', Phys. Rev. B 15, p2651 (1977)





Amplitude vs phase noise

- Phase noise is dominated by two level system noise in the dielectric (Rami's thesis)
- Amplitude noise dominated only by amplifier

