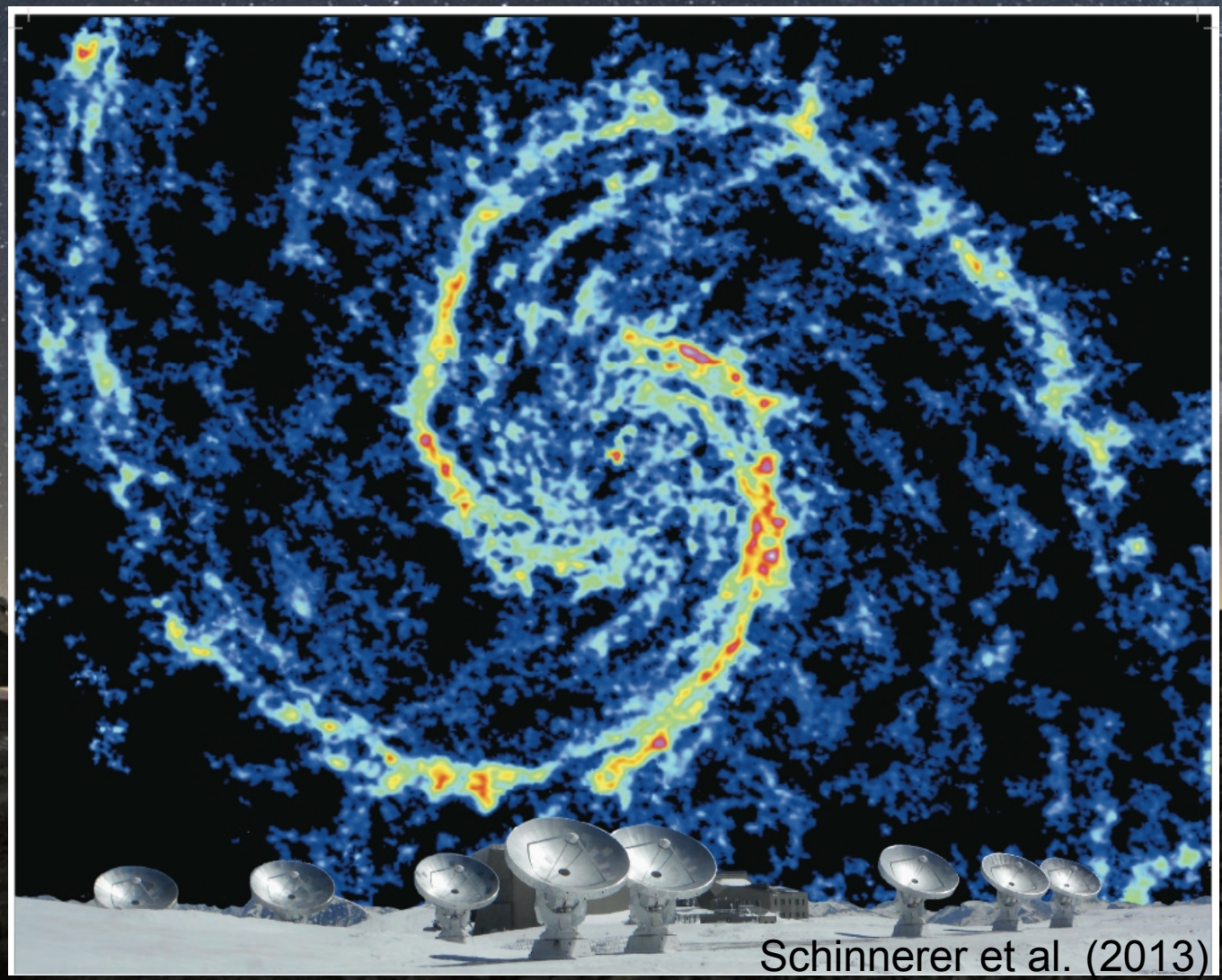
How to use NOEMA

Melanie Krips
IRAM, Grenoble

How to use NOEMA

Motivation

* LP from 2009!
* 60 pointings
* Only 1 line ^{12}CO
* 5-6 antennas in array at the time
* 200h of t_{obs} !



Schinnerer et al. (2013)

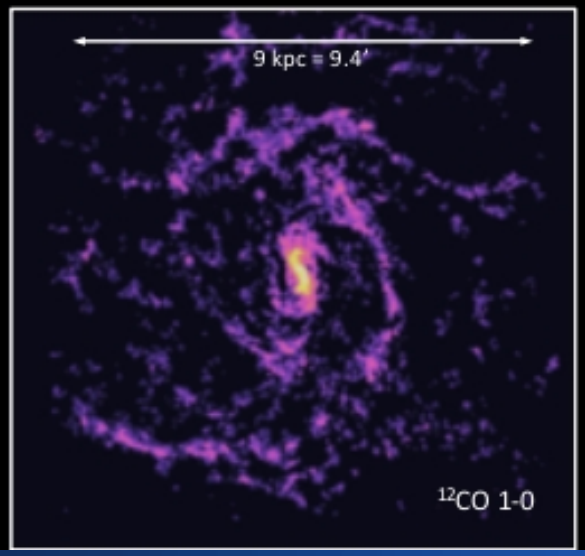
How to use NOEMA

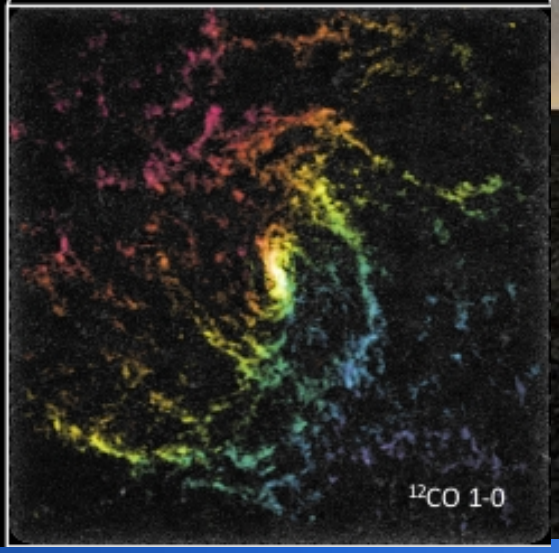
Motivation

* 2016!
* >900 pointings
* only ^{12}CO
* 6 antennas in array
* ~50h of t_{obs} !
  (not as deep as PAWs!)

Molecular clouds in IC342
PI A.Schruba (MPE)

- D = 3.3 Mpc, M(gas) = 10^{10} $M_\odot$, SFR = 1.9 $M_\odot$/yr
- NOEMA + IRAM 30m cover 70% of the SF disk
- NOEMA = 1250-field mosaic, 60 pc resolution = 3.8"
- 1500 molecular clouds with S/N > 5

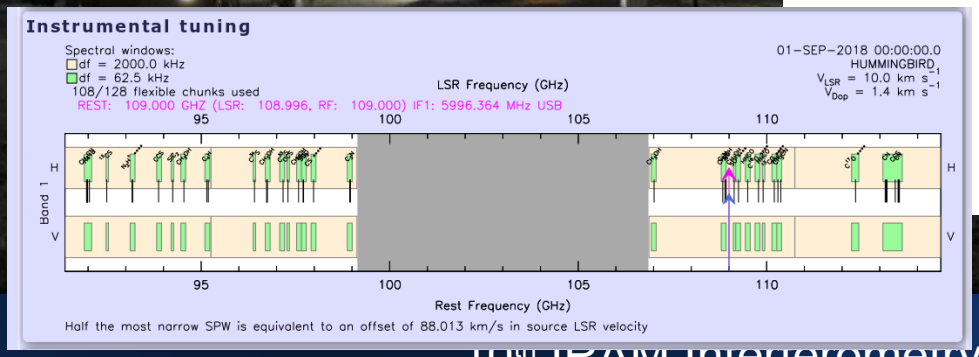


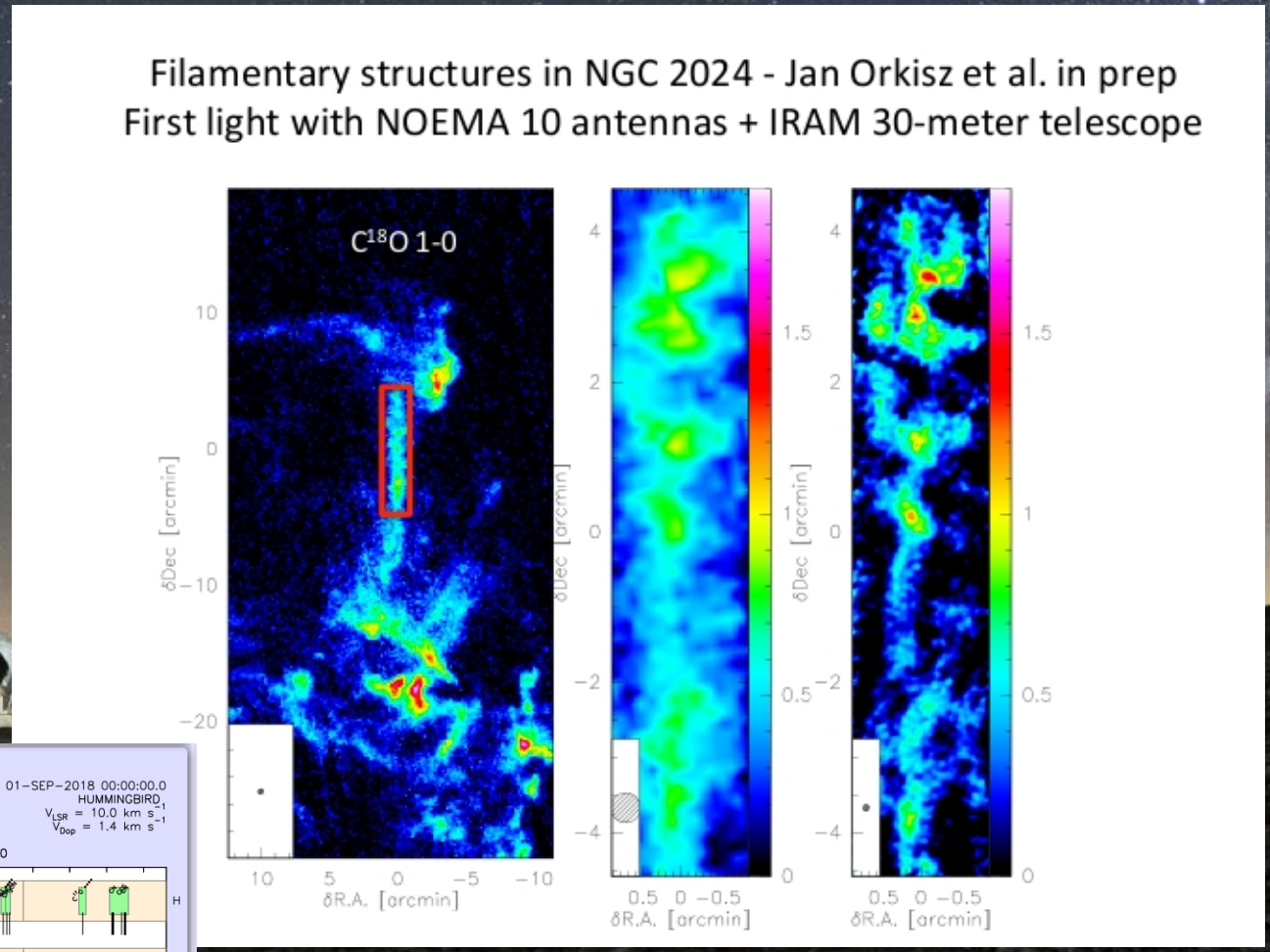


How to use NOEMA

Motivation

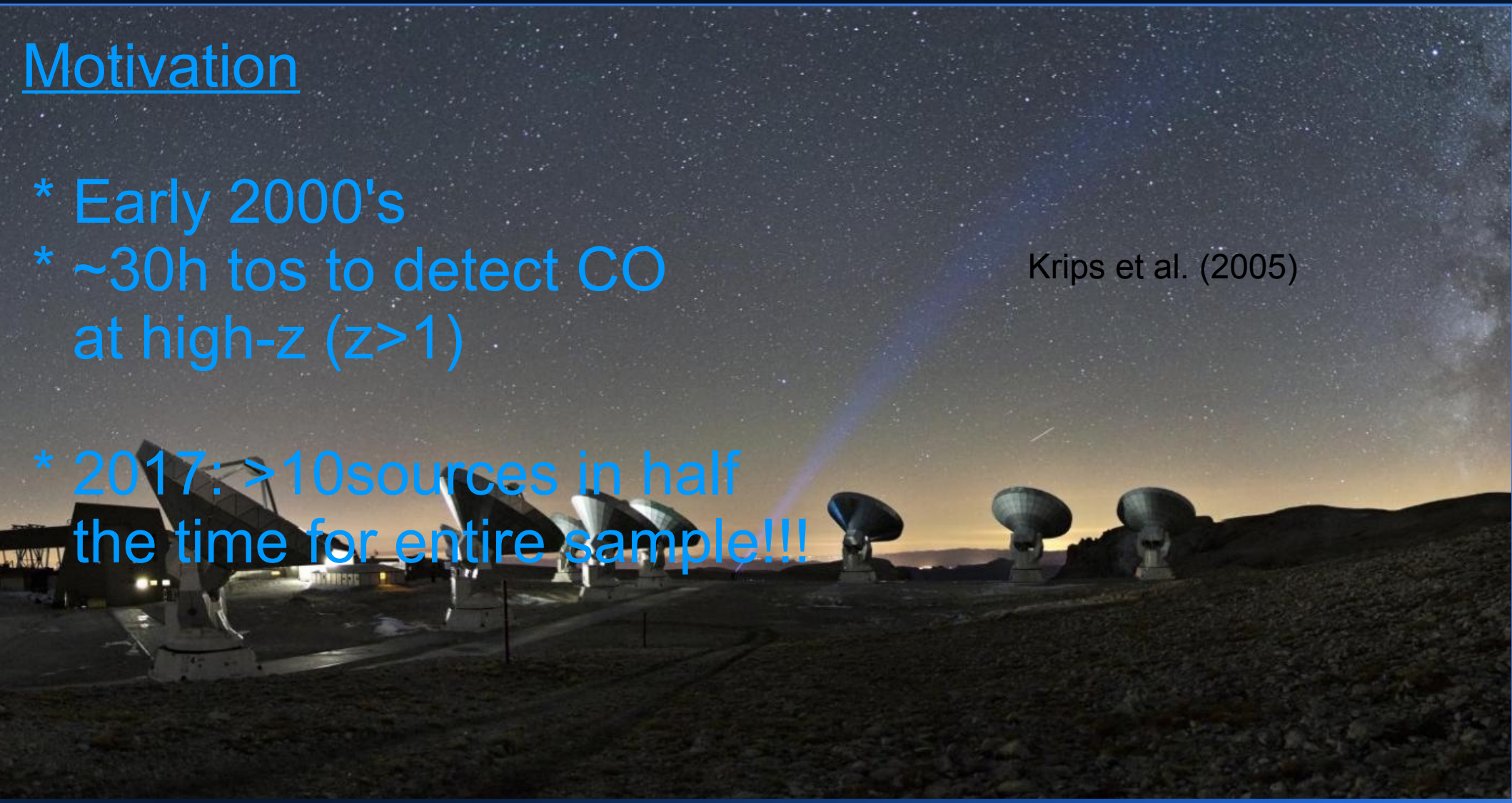
- * Aug/Sep2018!
- * >100 pointings
- * Many lines!
- * 8-10 antennas in array
- * ~16h of t_{obs} !!!!!!





Orkisz et al., in prep.

How to use NOEMA



Motivation

* Early 2000's
* ~30h tos to detect CO at high-z (z>1)

Krips et al. (2005)

* 2017: >10sources in half the time for entire sample!!!

How to use NOEMA

How to use NOEMA

This gives you (scientific) ideas?
(If it did not already .... ;) ....)

GREAT!

How to use NOEMA

Proposal Preparation

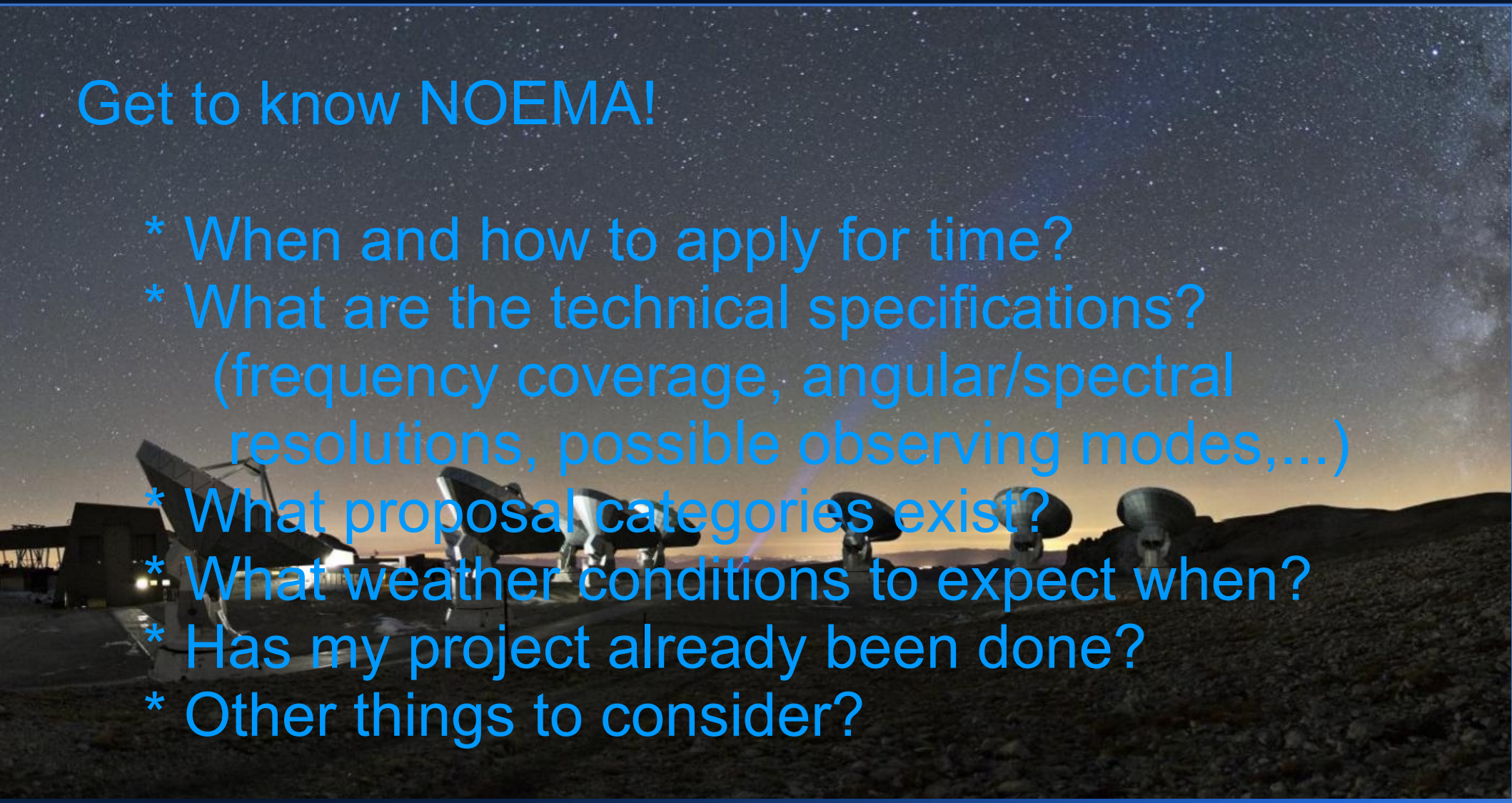
Reviewing Process

Observations

Data Reduction

How to use NOEMA

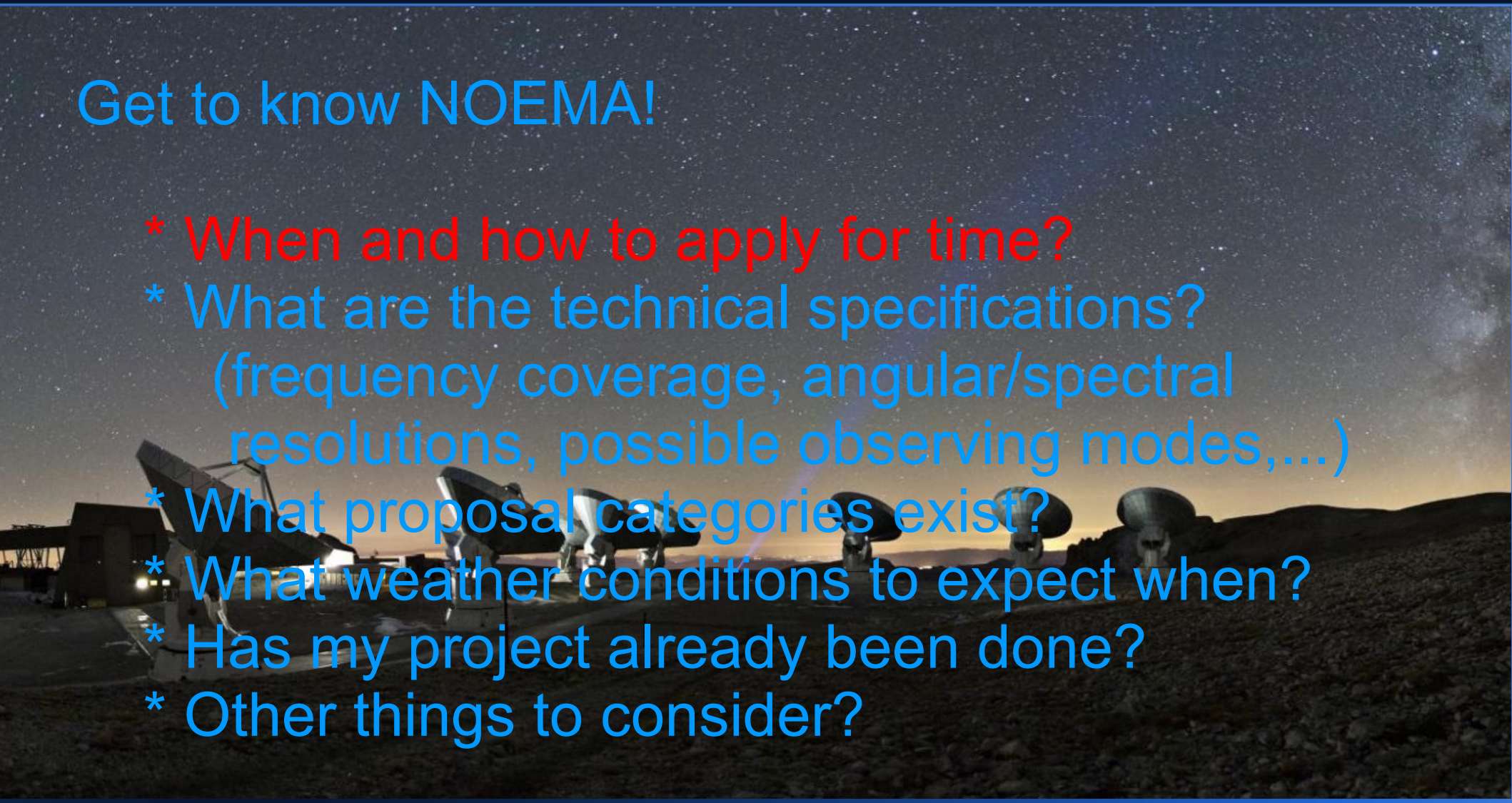
How to use NOEMA



Get to know NOEMA!

* When and how to apply for time?
* What are the technical specifications? (frequency coverage, angular/spectral resolutions, possible observing modes,...)
* What proposal categories exist?
* What weather conditions to expect when?
* Has my project already been done?
* Other things to consider?

How to use NOEMA



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How to use NOEMA

Can everyone apply?

YES!

3. Up to 15% of the available observing time may be invested into projects submitted by PIs affiliated with institutes in non-IRAM partner countries.

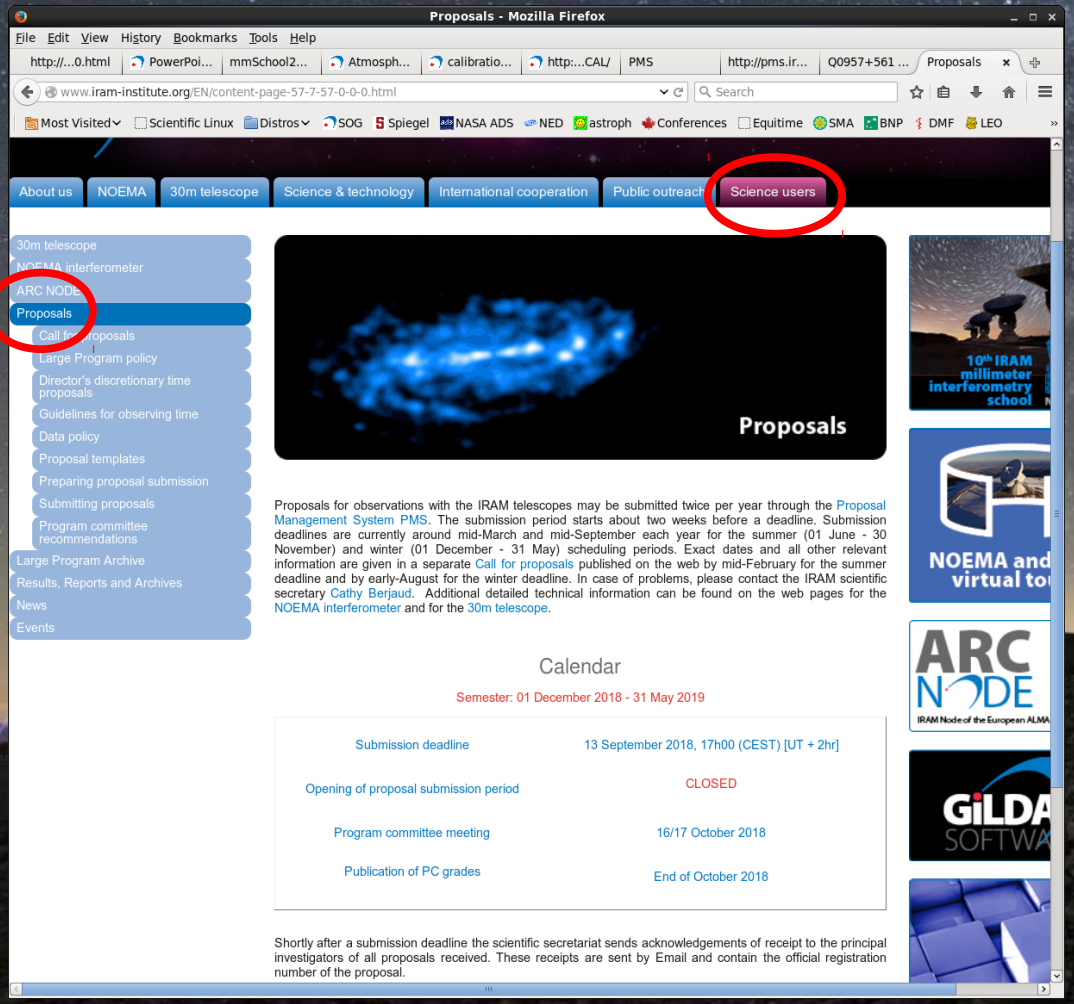
Geographic distribution of NOEMA users across the world between 2010 and 2017. During this period, 1790 individual investigators from 43 countries submitted proposals for the NOEMA Interferometer.

How to use NOEMA

When to apply for time?

2 deadlines per year:

* 2nd week of March
  Summer
  01 June – 30 Nov
* 2nd week of Sep
  Winter
  01 Dec – 31 May



How to use NOEMA

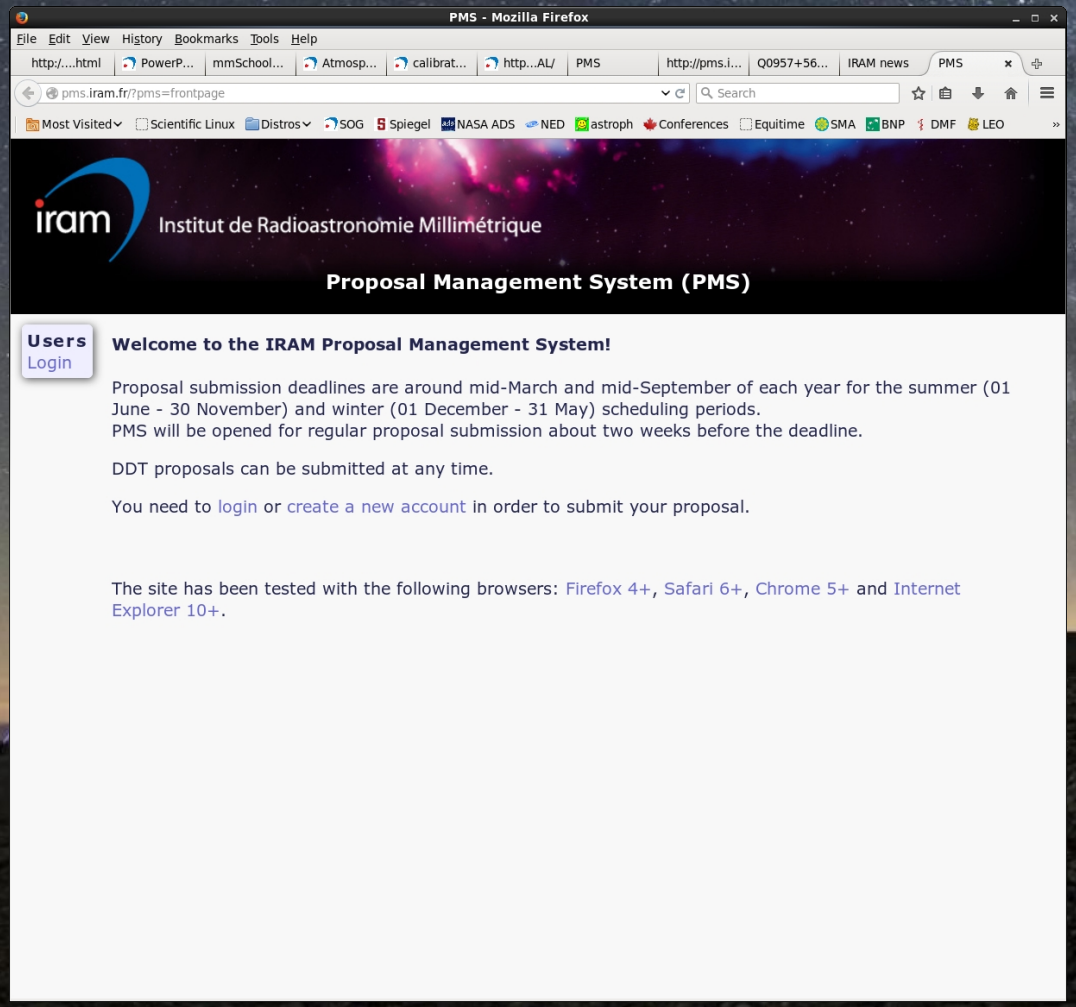
How to apply for time?

Use online proposal system to apply:

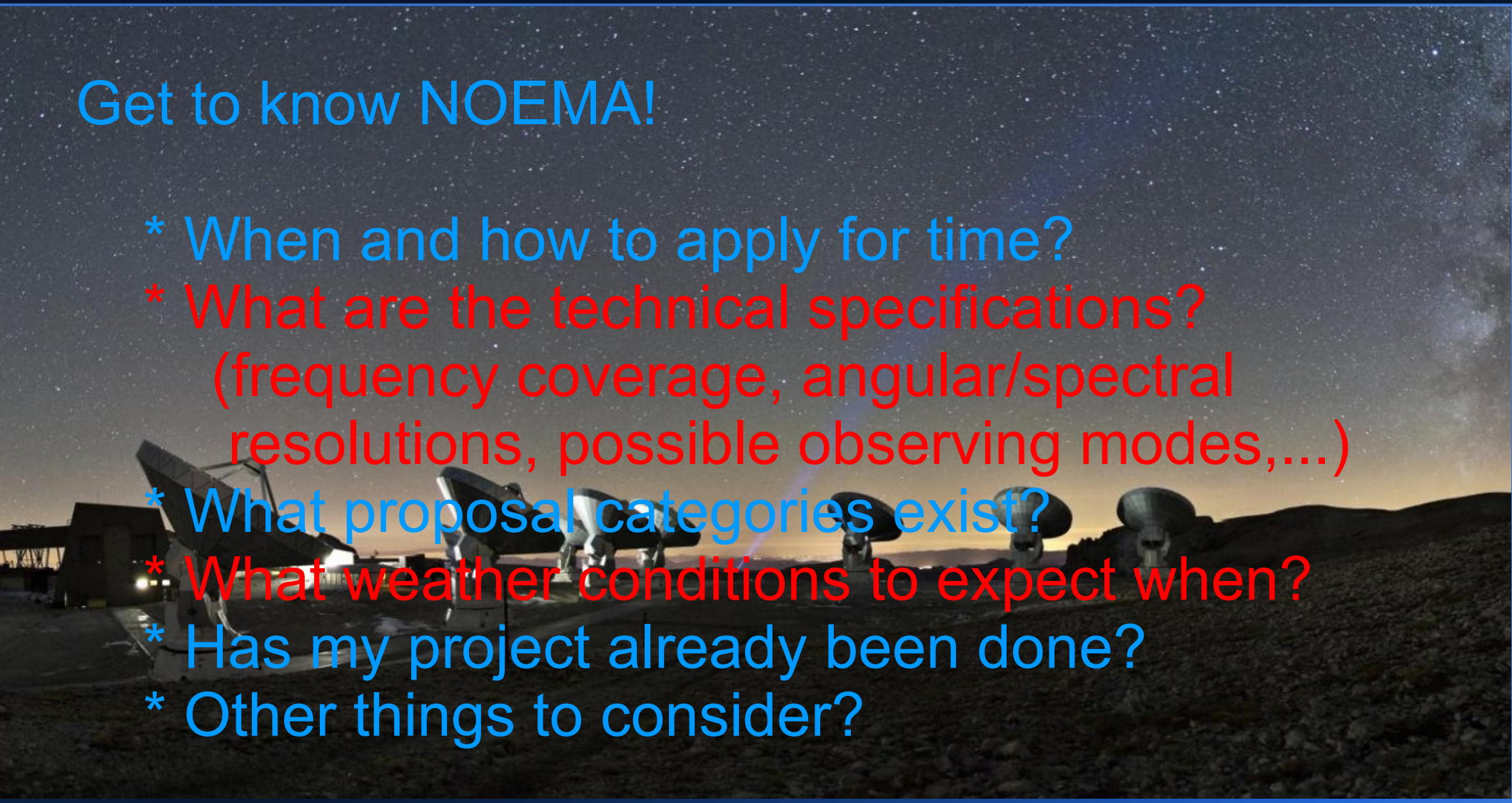
* PMS:

http://pms.iram.fr

-> see talk by Ch. Lefevre



How to use NOEMA



Get to know NOEMA!

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* What proposal categories exist?
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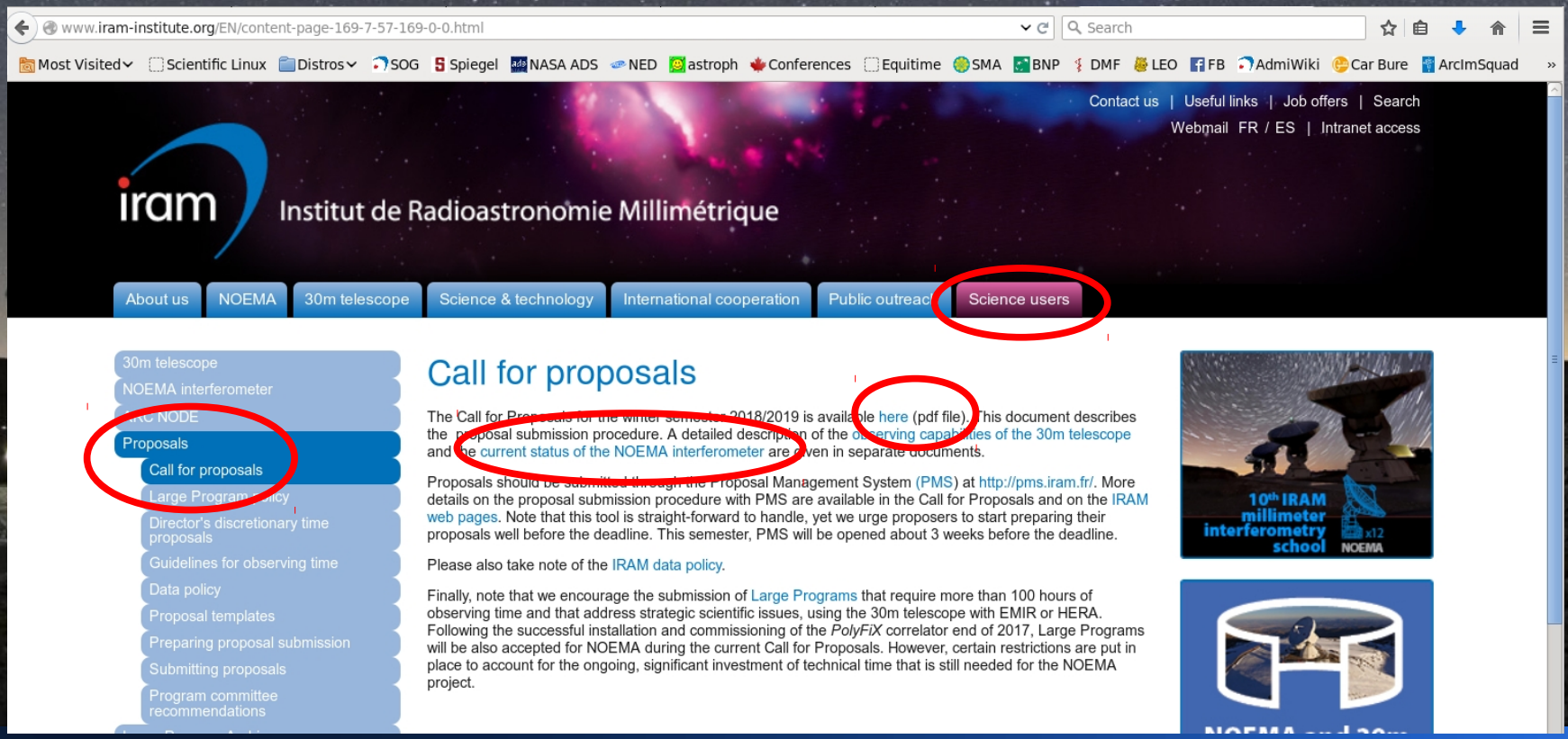
How to use NOEMA

Technical Considerations:

* Read the Call for Proposals and the "Technical Handbook" for NOEMA:



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www.iram.fr/GENERAL/calls/w18/w18.pdf

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Page: 1 of 5 — Automatic Zoom

Call for Proposals on IRAM Telescopes

The deadline for submission of observing proposals on IRAM telescopes, both the NOEMA interferometer and the 30-meter telescope, covering the scheduling period 1 December 2018 to 31 May 2019, is

13 September 2018, 17:00 CEST (UT + 2 hours)

IRAM proposals should be submitted through the *Proposal Management System* (**PMS**) at URL:

http://pms.iram.fr/

PMS provides on-screen instructions to guide the proposal editor through the submission process. The procedure consists in filling in an on-line form with the details of the requested observations (source coordinates, receiver setups, array configuration, etc.), and to upload a single file in pdf format containing the scientific and technical justification. A LaTeX template is provided from the PMS submission page for your convenience. This file may be customized, or the pdf file can be generated with another software, but in any case **proposers should respect the following requirements**: (1) A normal proposal may contain up to two pages of text describing the scientific aims and the technical justification (4 pages for a Large Program, see below) (2) up to two pages of figures, tables, and references may be added (but the text should not be mixed with figures, tables, and references!), and (3) the font size must be at least 11pt.

For a proposal to be complete, PMS requires that all authors validate their identity (e–mail and affiliation) and their participation to the proposal before the deadline. The editor of the proposal will have to send invitations to all authors through PMS by clicking an *invitation* button. We urge proposal editors to invite the authors through PMS well before the deadline to give them enough time to validate their identity before the deadline. Authors that fail to validate their participation will automatically be dropped from the proposal.

How to use NOEMA

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www.iram.fr/GENERAL/calls/w18/NOEMACapabilities.pdf

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Page: 1 of 8 — 70%

IRAM NOEMA interferometer

Observing Capabilities and Current Status

Melanie Krips

July 30, 2018

This document is updated twice a year to reflect the capabilities of the interferometer at the time of the *Call for Proposals* publication. Non-trivial changes with respect to the previous version are marked in red. Note that this document contains active links marked with a `different font` for an easy access to documentation, e.g. on the IRAM web pages. The full links are also given on the last page of this document.

Table 1: Configuration Schedule

Conf	Scheduling Priority Winter 2018/2019
C	November – December
D	December – January
A	January – February
C	February – March
D	March – May

1 Progress of NOEMA

1.1 NOEMA's 10th antenna

The commissioning of antenna 10 is scheduled for the end of summer 2018. We plan to operate with all 10 antennas during the winter period 2018/2019.

2 Conditions for the next winter session

During the course of the winter semester, we plan to schedule three configurations. The number of antennas is dependent on the availability of antenna 10. A preliminary configuration schedule for the winter period is outlined in Table . Adjustments to this provisional configuration planning will be made according to commissioning requirements in the frame of NOEMA, proposal pressure, weather conditions, and other contingencies. The configuration schedule given in Table should be taken as a rough guideline, in particular for astronomical targets that cannot be observed during parts of the winter period because of sun avoidance constraints.

The winter semester is preferred for high frequency (1 mm) and high angular resolution observations. Nevertheless, we encourage the submission of proposals that ask for observations at lower frequencies (2 mm & 3 mm) and lower angular resolution for which a significant amount of time can be invested. Observations in band 4 will not be offered this semester.

Unfinished A-rated programs from the current summer semester 2018 will be carried over into the upcoming winter semester. However, B-rated projects (or their sub-parts) from the summer semester 2018 that have not been started by the end of November 2018 will not be carried over and should be re-submitted. Investigators who wish to check the status of their project may consult the `interferometer schedule` on the IRAM website. This page is updated daily.

2.1 General Proposal Considerations

Please give high importance to the quality of your proposal. The NOEMA interferometer is a powerful, but complex instrument, and proposal preparation requires special care, especially in light of its new capabilities. In particular, your proposal should not only justify the scientific interest, but also the need for NOEMA. Proposers should note in their application whether the same or a similar proposal was or is intended to be submitted to ALMA, in which case a special justification is required to explain why IRAM interferometer time is needed.

Don't hesitate to contact the NOEMA Science Operations Group (sog@iram.fr) in case of doubts and for

How to use NOEMA

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Page: 1 of 8 — 70%

All that is new is marked in red

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How to use NOEMA

Technical Considerations:

* What deadline to choose?
  - What frequency (band) do you need?
  - Which configuration do you need?
  - Are the constraints wrt to source visibility?

How to use NOEMA

Frequency Coverage/bands offered:

* 3 bands offered: 3mm, 2mm & 1mm (band 4 currently not available)

Table 3: Receiver characteristics

	Band 1	Band 2	Band 3
F_{LO1} range/[GHz]*	82.0–108.256	138.616–171.256	207.744–264.384
F_{sky} range/[GHz]*	70.4–119.872	127.000–182.872	196.128–276.000
T_{rec}/[K]**	25–45	35–55	40–70
G_{im}/[dB]	-15-...-10	-15...-10	-15 ... -10

* Guaranteed LO1 frequency ranges per offered band. The LO1 frequency is the center frequency between the USB and LSB that can both be simultaneously observed in one tuning (see Fig 1). The center frequency of the USB (LSB) is separated by +(−)7.744 GHz from the LO1 frequency. With an effective width of 7.744 GHz per sideband the lowest and highest sky frequencies that can be covered per tuning are hence $F_{sky}=F_{LO1}\pm11.616$ GHz. The lowest and highest LO1 frequencies per band define the F_{sky} ranges that are guaranteed for this call.

** for LSB and USB.

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How to use NOEMA

Frequency Coverage/bands offered:

* 2 sidebands covering each 7.744GHz in 2 polars

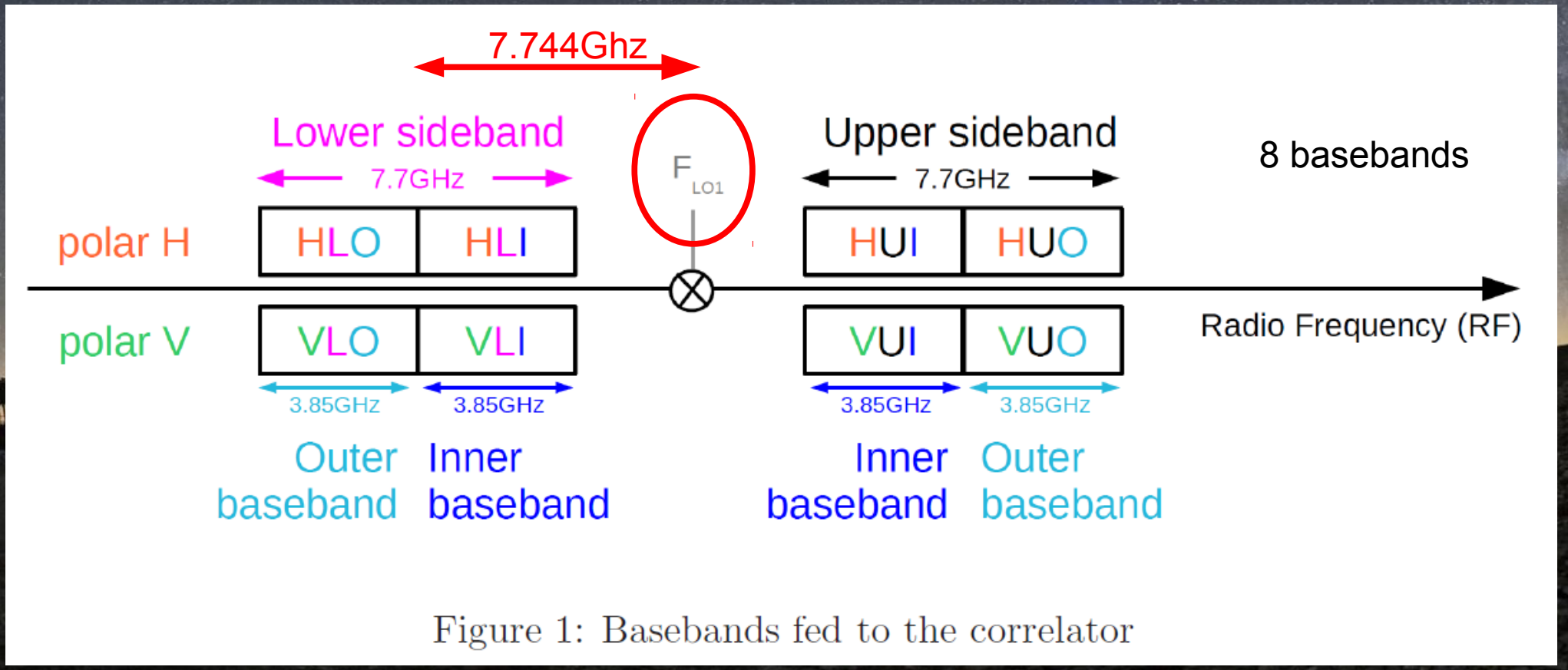

Figure 1: Basebands fed to the correlator

How to use NOEMA

Frequency Coverage/bands offered:

* 2 sidebands covering each 7[illegible] in 2 polars

F_{sky} coverage

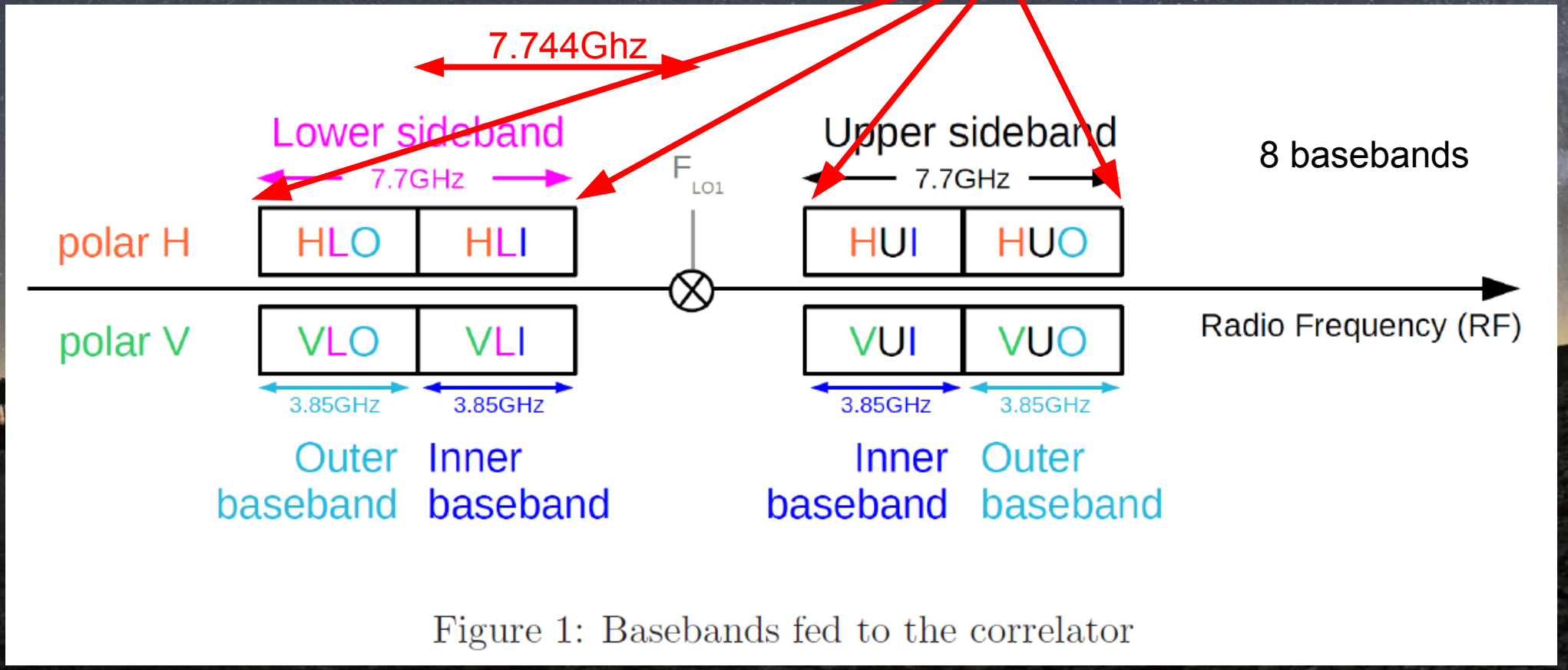

Figure 1: Basebands fed to the correlator

How to use NOEMA

Frequency Coverage/bands offered:

* 3 bands offered: 3mm, 2mm & 1mm (band 4 currently not available)

Table 3: Receiver characteristics

	3mm Band 1	2mm Band 2	1mm Band 3
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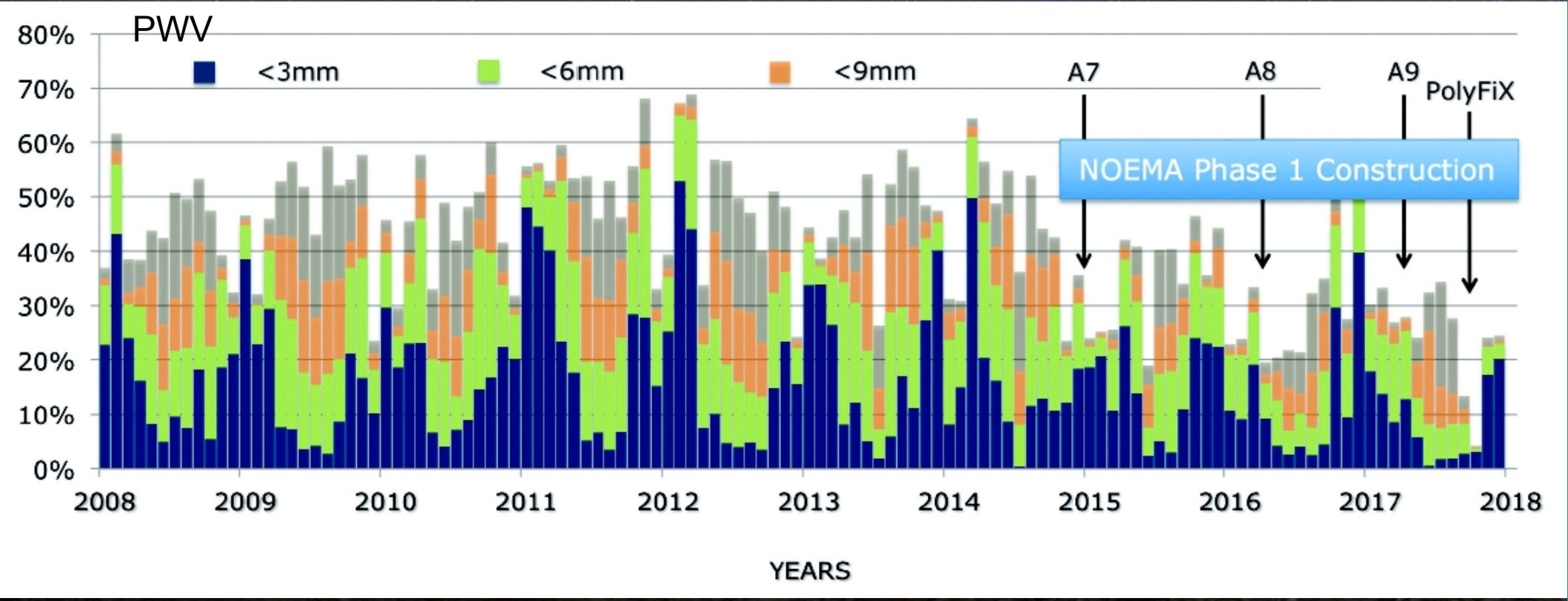
How to use NOEMA

When to observe what band:

NOEMA correlation time and atmospheric water vapor statistics during on sky operation over the last ten years.

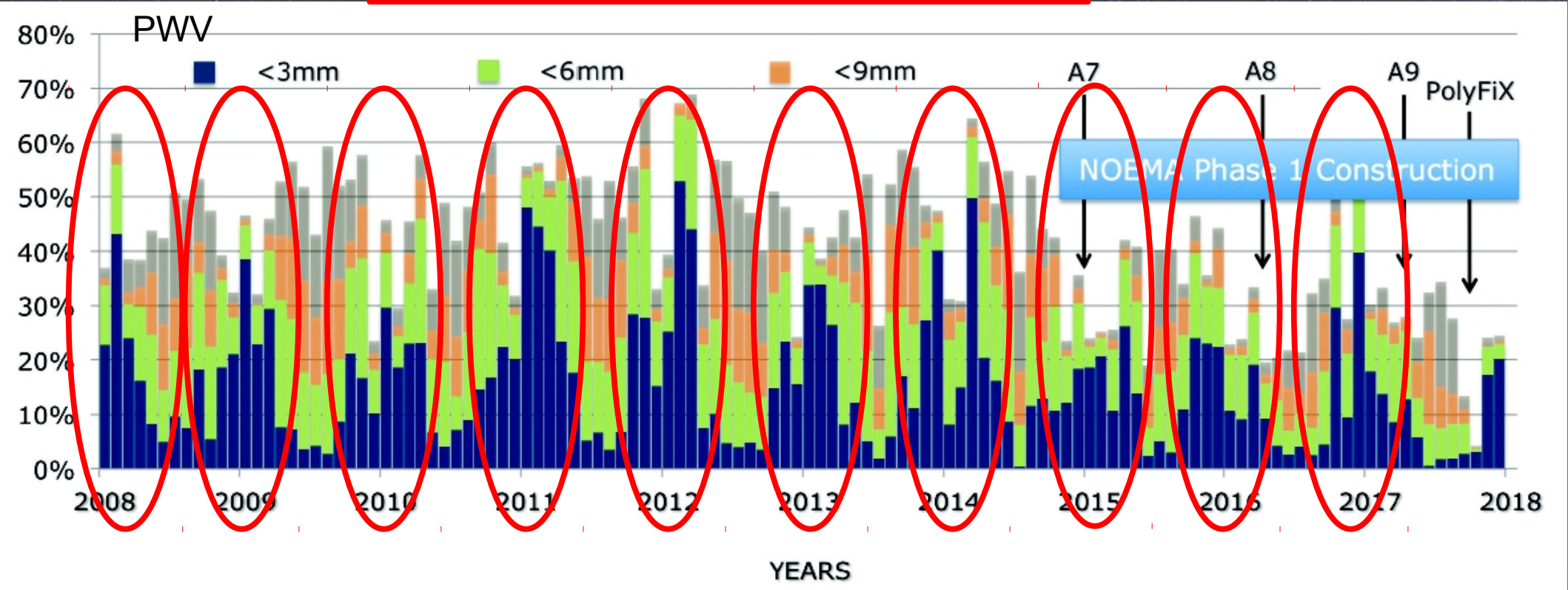


How to use NOEMA

When to observe what band:

Winter semester has the best weather conditions (lowest pwv)!

NOEMA correlation time and atmospheric water vapor statistics during on sky operation over the last ten years.



How to use NOEMA

When to observe what band:

Band3 (1mm): pwv<5mm

Band2 (2mm): pwv<7mm

Band1 (3mm): pwv<15mm

How to use NOEMA

When to observe what band:

Band3 (1mm): pwv<5mm

Band2 (2mm): pwv<7mm

Band1 (3mm): pwv<15mm

WINTER!
(Sep - Apr)

SUMMER!
(May - Aug)

How to use NOEMA

Offered Configurations:

* In total 3 configurations available: D, C and A

WINTER

10D	W12	W08	W05	E10	E04	N17	N13	N09	N05	N02
10C	W23	W20	W09	E23	E18	E10	E03	N20	N17	N11
10A	W27	W23	W08	E68	E24	E16	E03	N46	N29	N20

Conf	Scheduling Priority Winter 2018/2019
C	November – December
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A	January – February
C	February – March
D	March – May

SUMMER

Name	Stations								
8D	W12	W08	W05	E10	E04	N13	N09	N02	–
9C	W20	W12	W09	E16	E10	E03	N29	N20	N11

Conf	Scheduling Priority Summer 2018
D	June – August
C	September – November

How to use NOEMA

Offered Configurations:

* In total 3 configurations available: D, C and A

WINTER

10D	W12	W08	W05	E10	E04	N17	N13	N09	N05	N02
10C	W23	W20	W09	E23	E18	E10	E03	N20	N17	N11
10A	W27	W23	W08	E68	E24	E16	E03	N46	N29	N20

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SUMMER

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9C	W20	W12	W09	E16	E10	E03	N29	N20	N11

Conf	Scheduling Priority Summer 2018
D	June – August
C	September – November

How to use NOEMA

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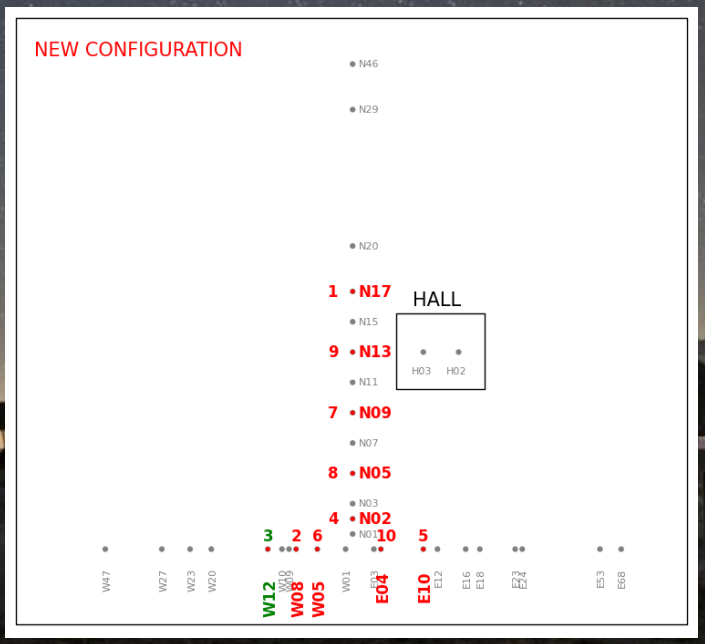
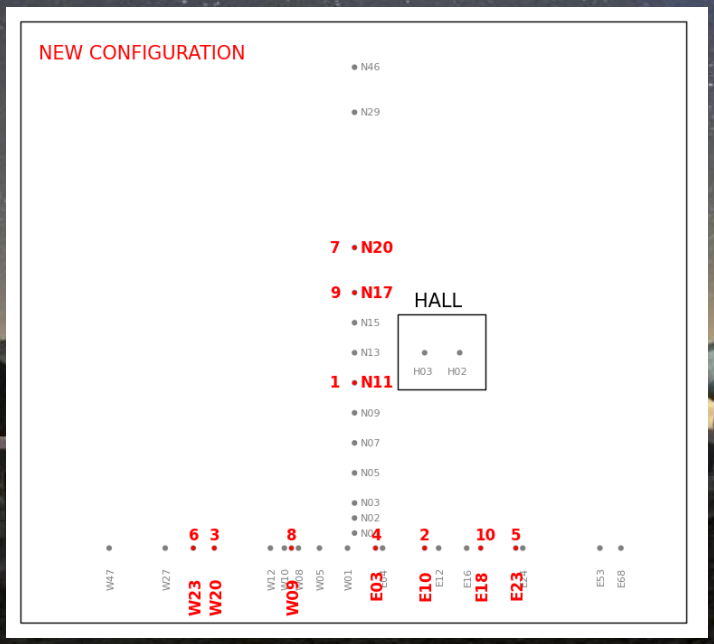
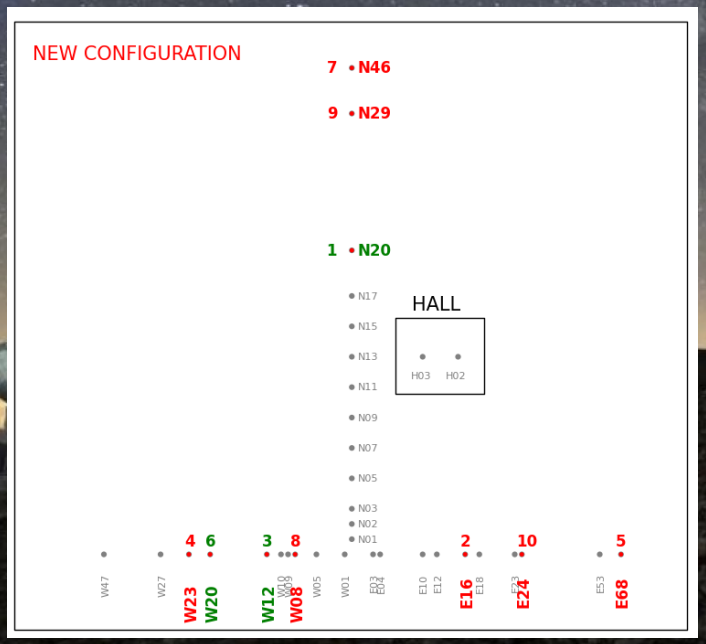
Configuration D



Configuration C



Configuration A



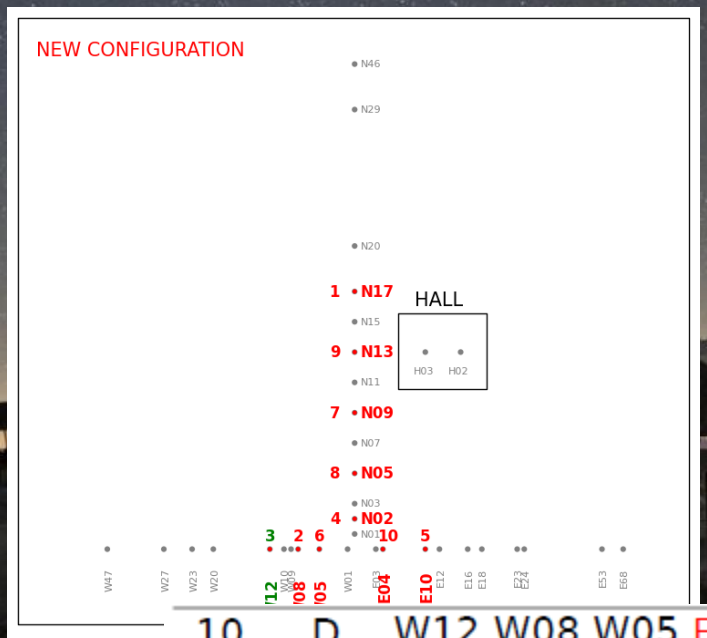
Increasing baseline length (b_{max} from ~80m to ~800m)

How to use NOEMA

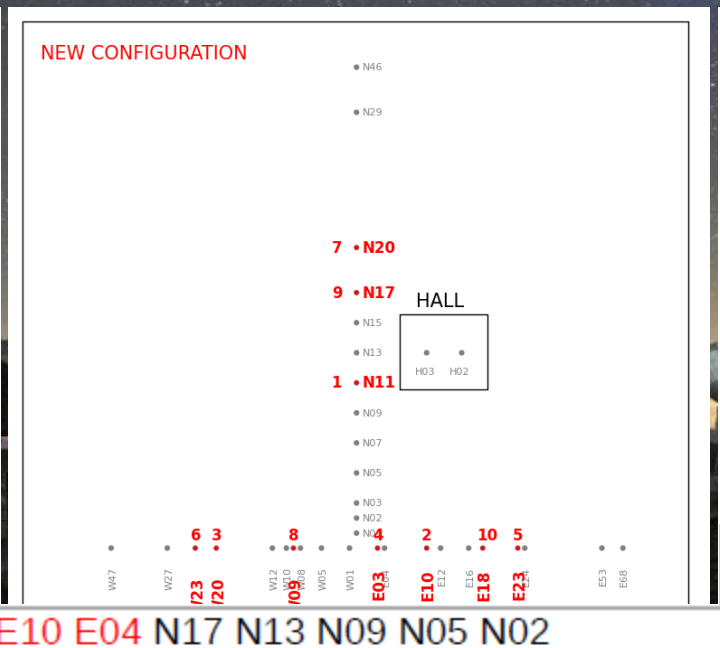
Offered Configurations:

* In total 3 configurations available: D, C and A

Configuration D



Configuration C



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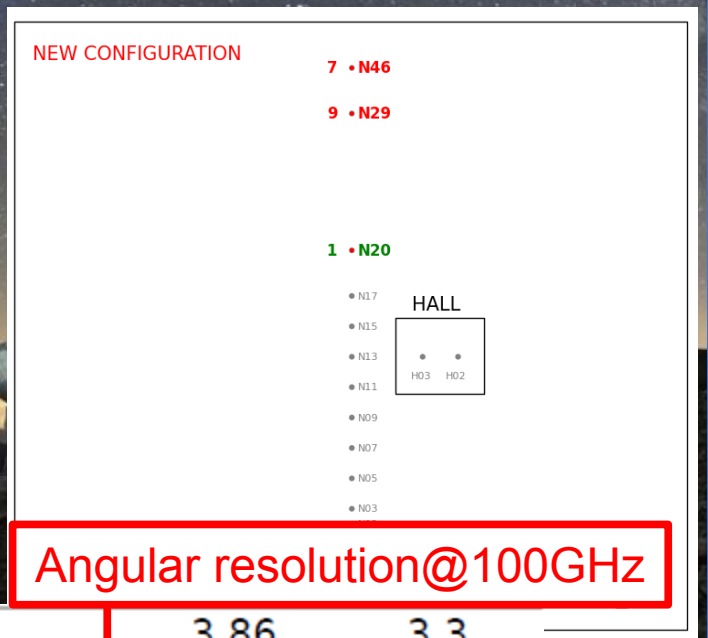


Angular resolution@100GHz

10	D	W12 W08 W05 E10 E04 N17 N13 N09 N05 N02	3.86	3.3
	C	W23 W20 W09 E23 E18 E10 E03 N20 N17 N11	2.01	1.75
	A	W27 W23 W08 E68 E24 E16 E03 N46 N29 N20	1.14	1.00

How to use NOEMA

Offered Configurations:

* In total 3 configurations available: D, C and A

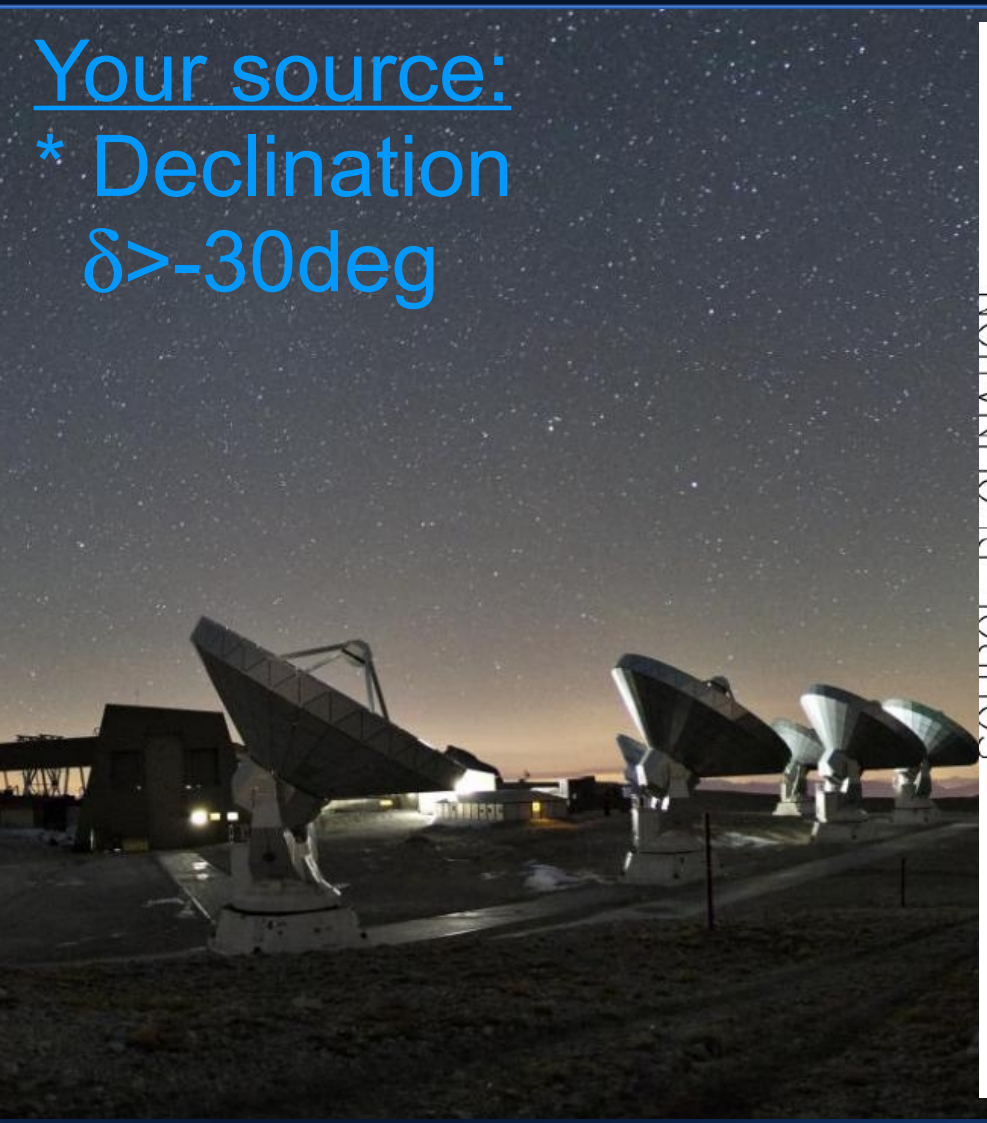
* Short Spacings needed from the 30m?

How to use NOEMA

Your source:

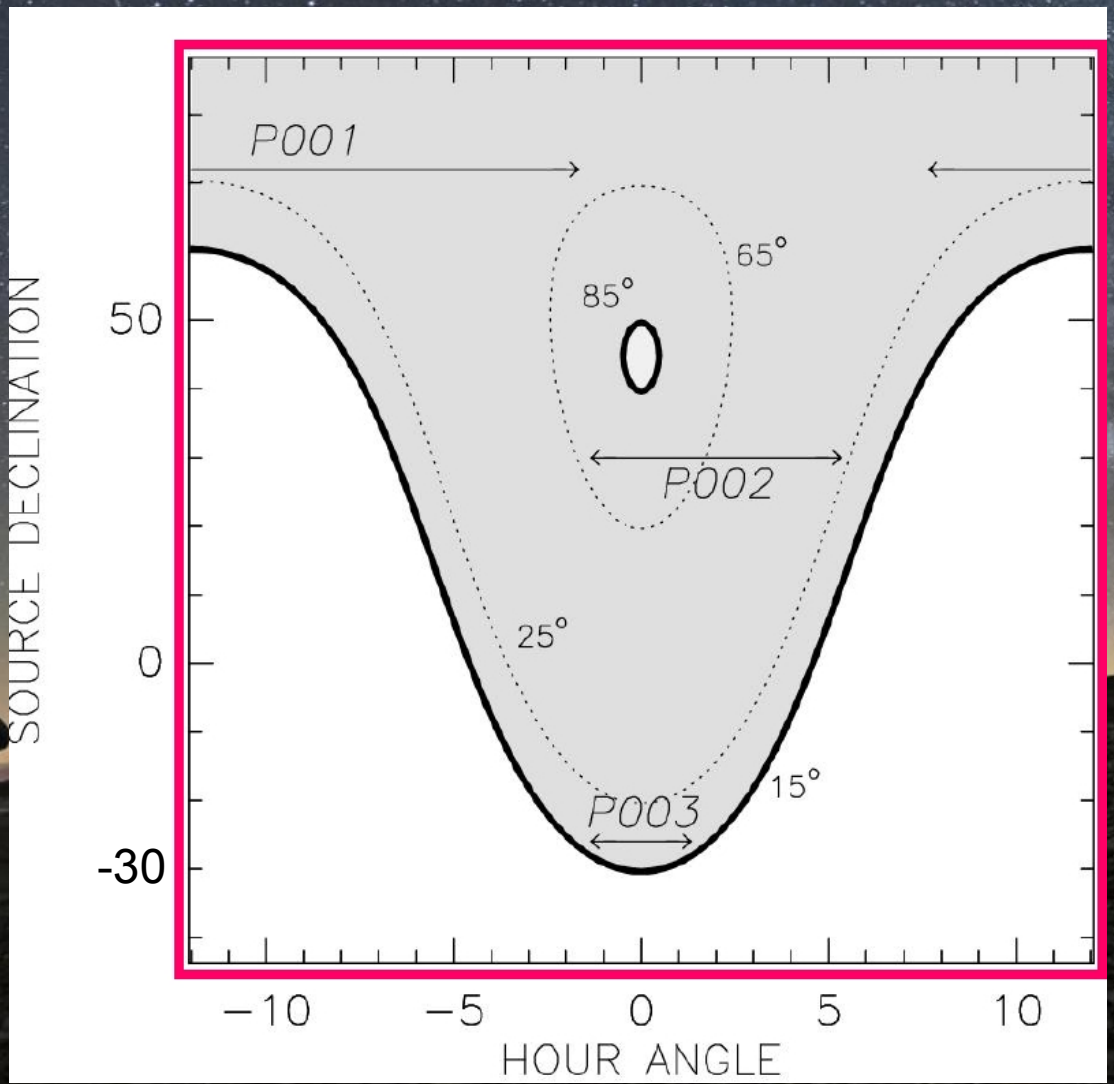
* Source position in the sky (RA, Declination)
  LST range
  Sun avoidance

How to use NOEMA



Your source:

* Declination
  δ>-30deg



How to use NOEMA

Your source:

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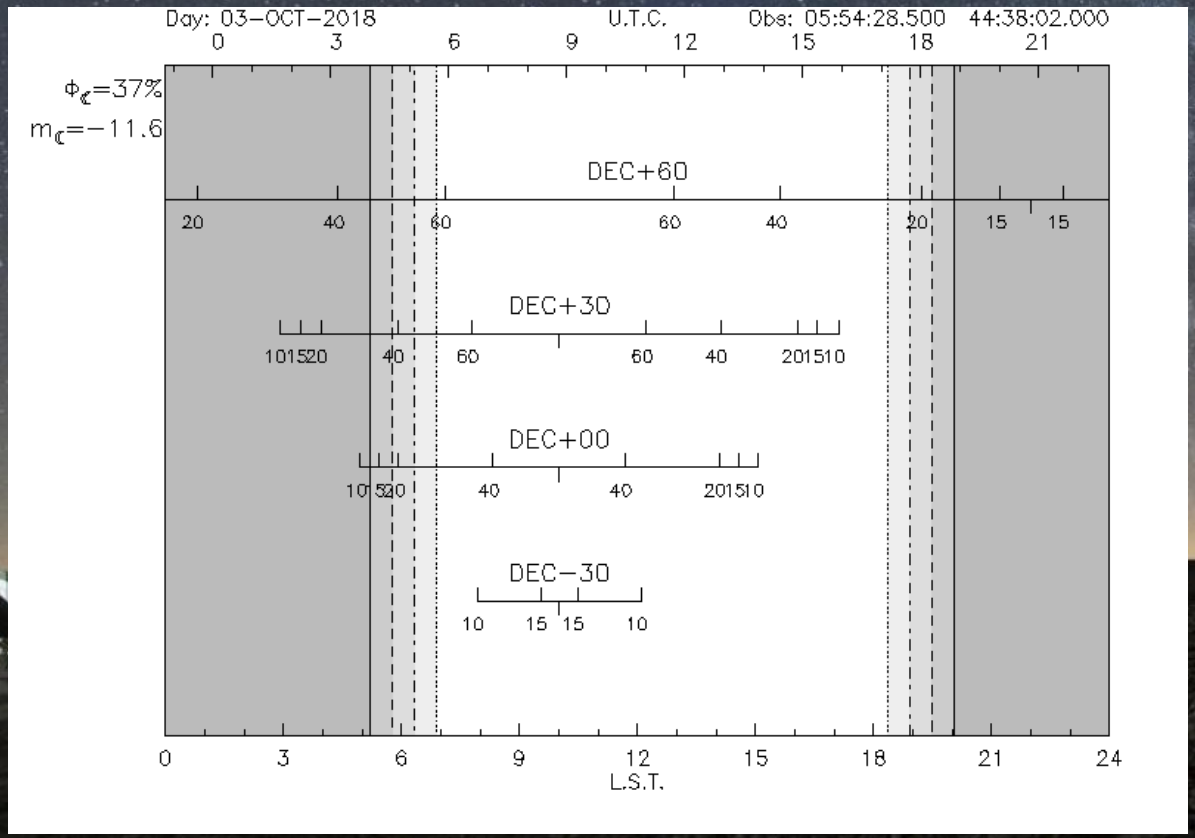
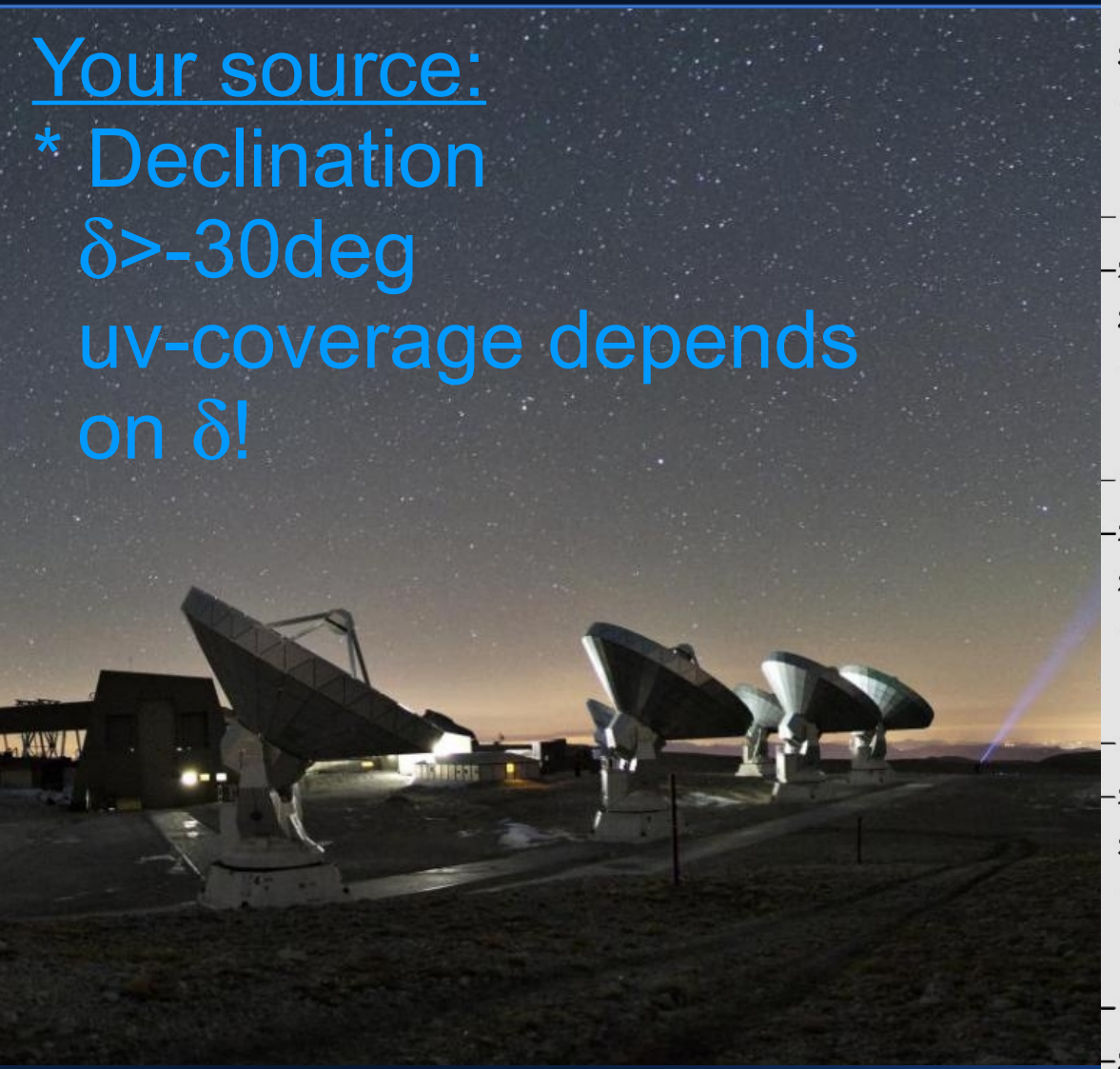
  δ>-30deg

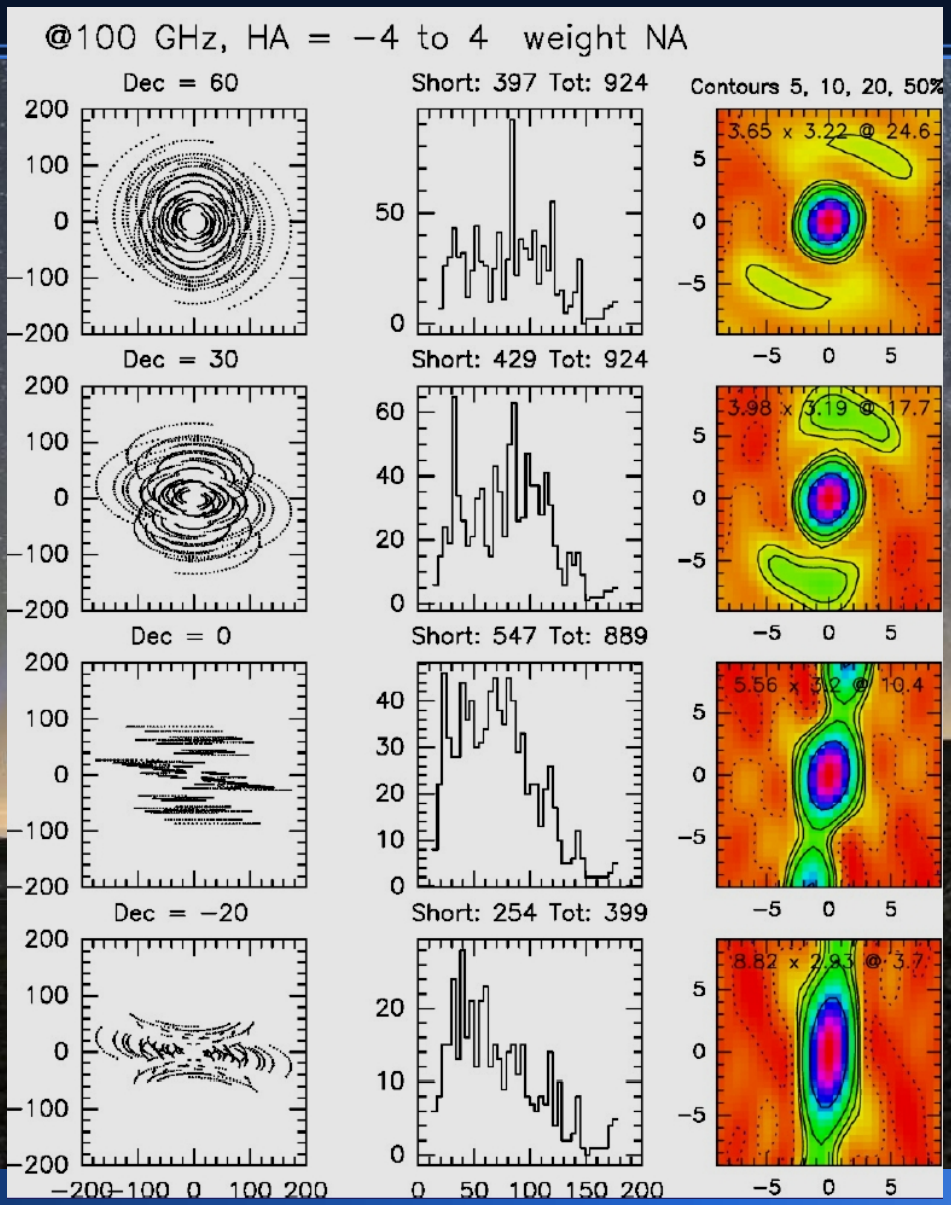


How to use NOEMA



Your source:

* Declination δ>-30deg uv-coverage depends on δ!

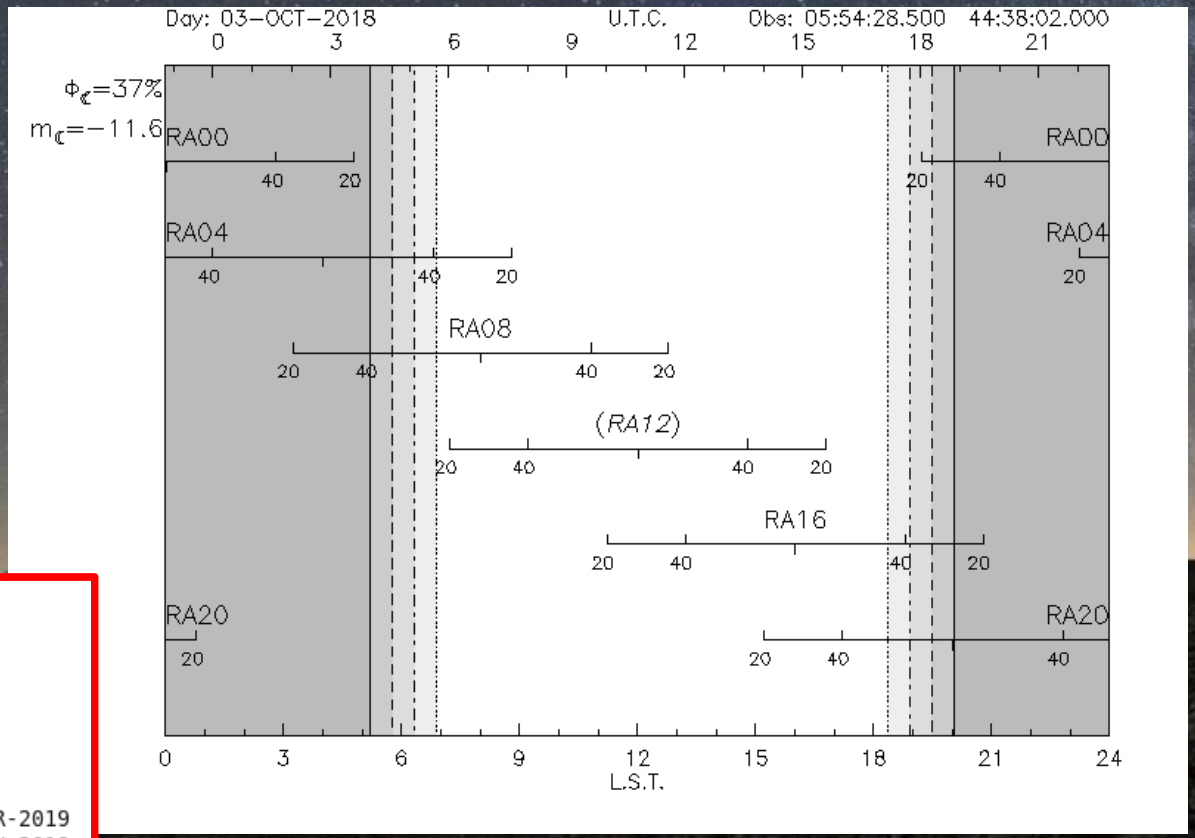


How to use NOEMA

Your source:

* Declination δ>-30deg uv-coverage depends on δ!
* Sun Avoidance/Az: Is source visible?

```
ASTRO> cata az.sou
I-CATALOG,  New source catalog is az.sou
ASTRO> time
ASTRO> obs noema
I-OBSERVATORY,  Selected NOEMA observatory
I-OBSERVATORY,  NOEMA status is set to year 2017
   >> NOEMA receivers with Polyfix backend
ASTRO> hori /sou /night
RA00            Sun distance  169.1   Avoidance 21-FEB-2019 to 26-APR-2019
RA04            Sun distance  129.2   Avoidance 19-APR-2019 to 23-JUN-2019
RA08            Sun distance   70.3   Avoidance 20-JUN-2019 to 25-AUG-2019
RA12            Sun distance   16.7   Avoidance 17-AUG-2018 to 21-OCT-2018
RA16            Sun distance   52.5   Avoidance 05-NOV-2018 to 01-DEC-2018
RA20            Sun distance  111.2   Avoidance 12-JAN-2019 to 07-FEB-2019
ASTRO>
```



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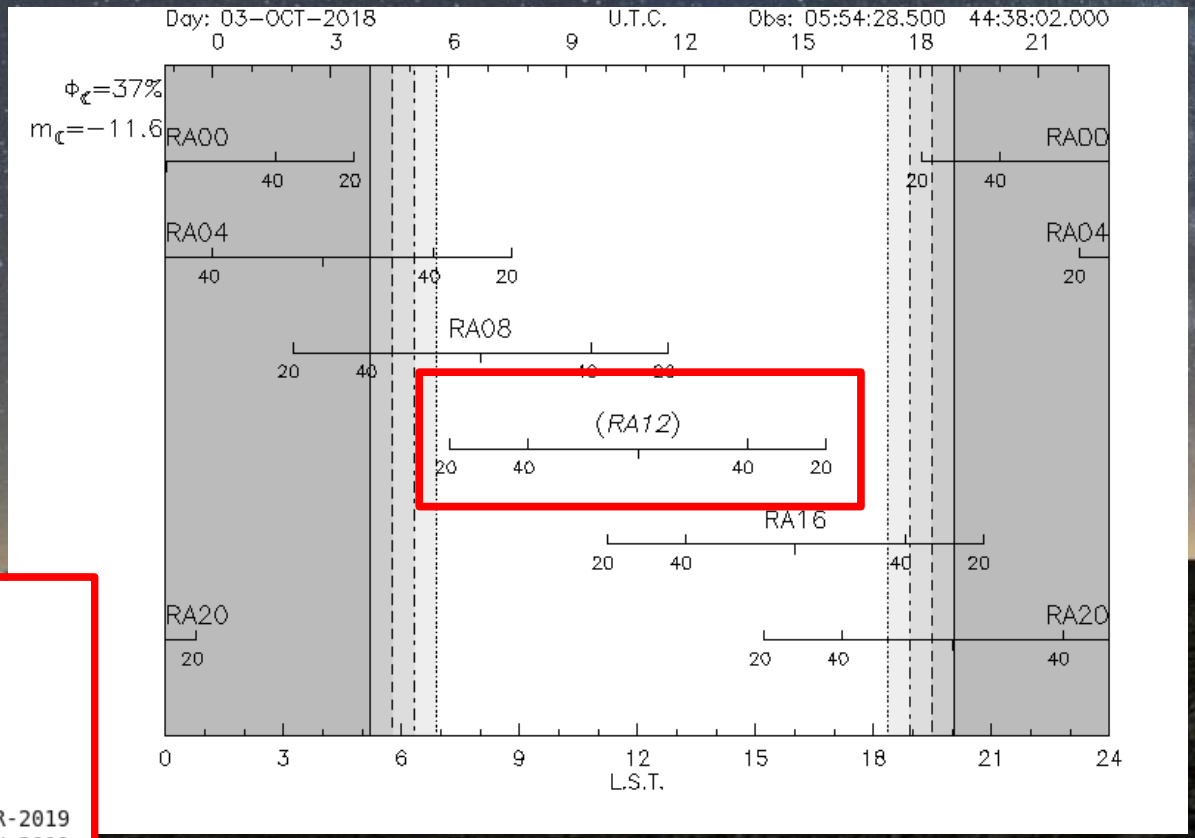


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ASTRO> hori /sou /night
RA00            Sun distance  169.1   Avoidance 21-FEB-2019 to 26-APR-2019
RA04            Sun distance  129.2   Avoidance 19-APR-2019 to 23-JUN-2019
RA08            Sun distance   76.3   Avoidance 20-JUN-2019 to 25-AUG-2019
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ASTRO>
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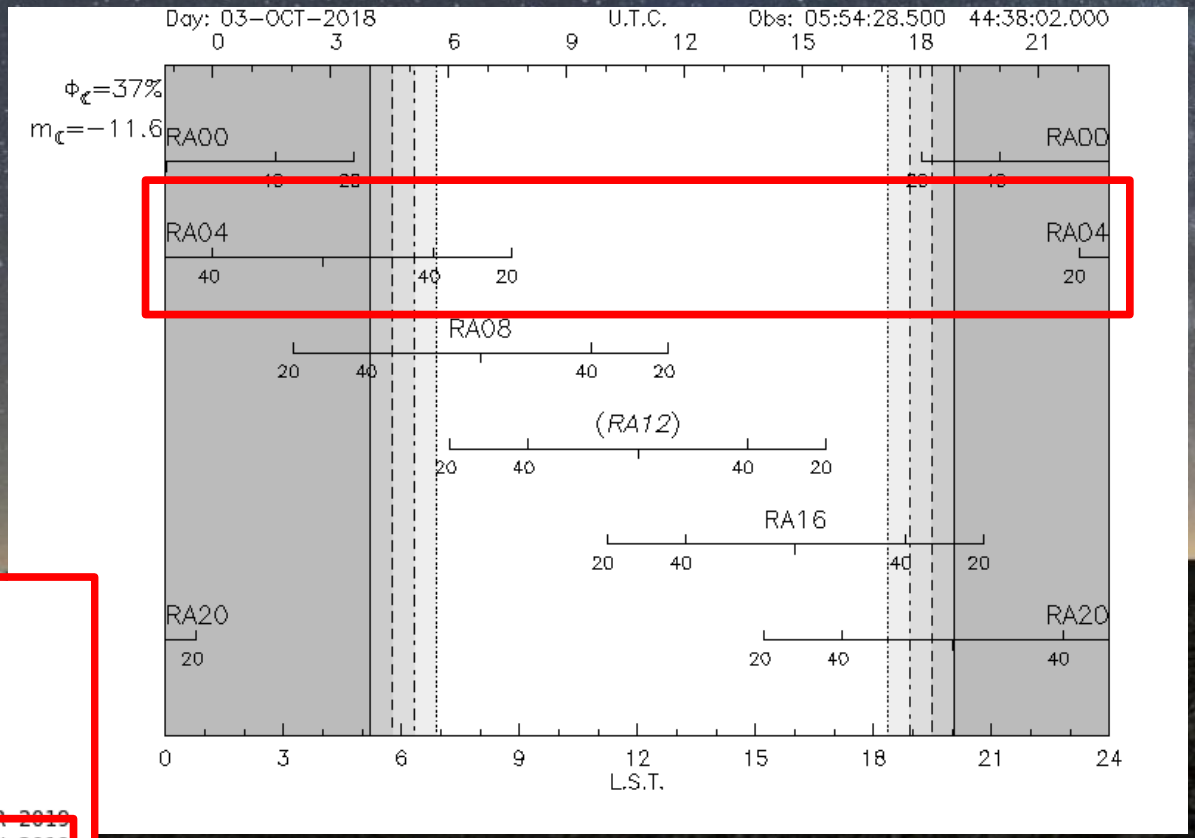


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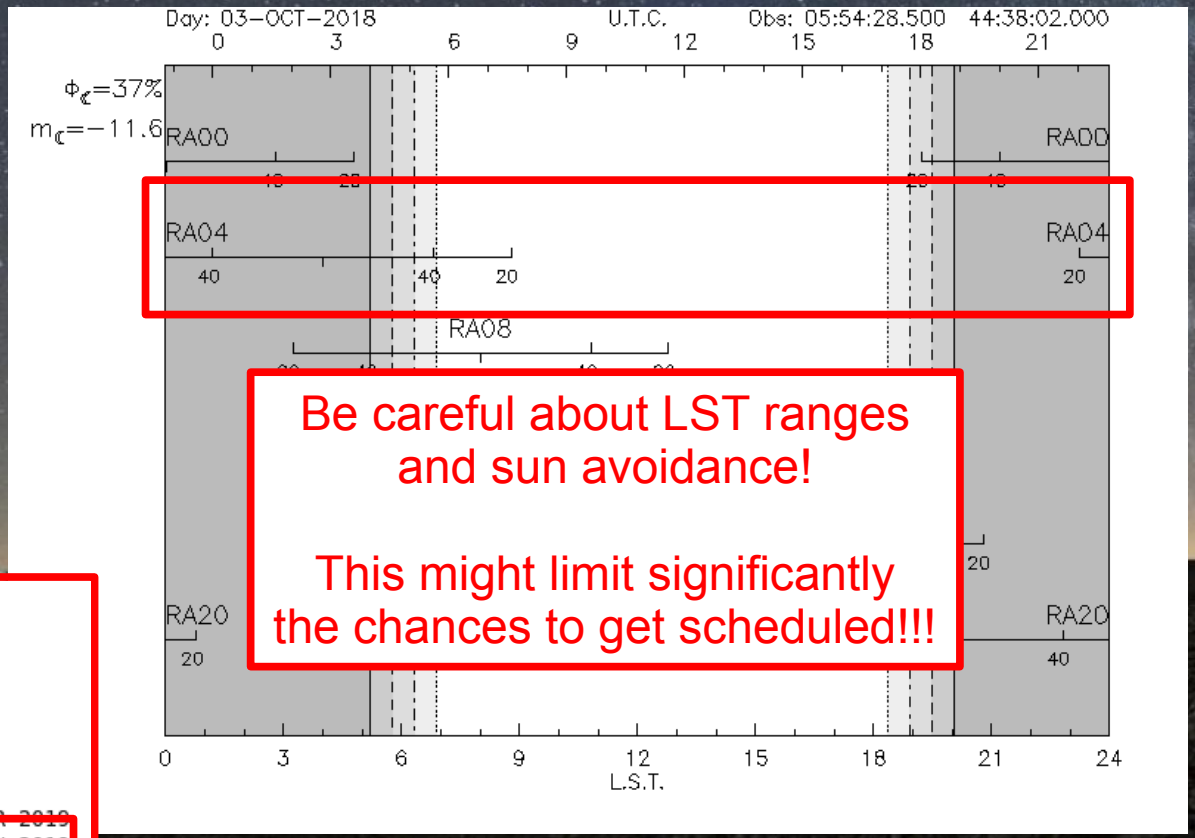


Be careful about LST ranges and sun avoidance!

This might limit significantly the chances to get scheduled!!!

How to use NOEMA

Sensitivity requirements:

* How to calculate sensitivities and corresponding telescope time?

$$\delta S = \frac{2k}{\eta_a A \cdot \eta_j \eta_C} \cdot \frac{\langle T_{sys} \rangle}{\eta_p \sqrt{N(N-1)} \sqrt{\delta\nu \, t_{\rm on}}} \cdot \frac{1}{\sqrt{N_{\rm pol}}}$$

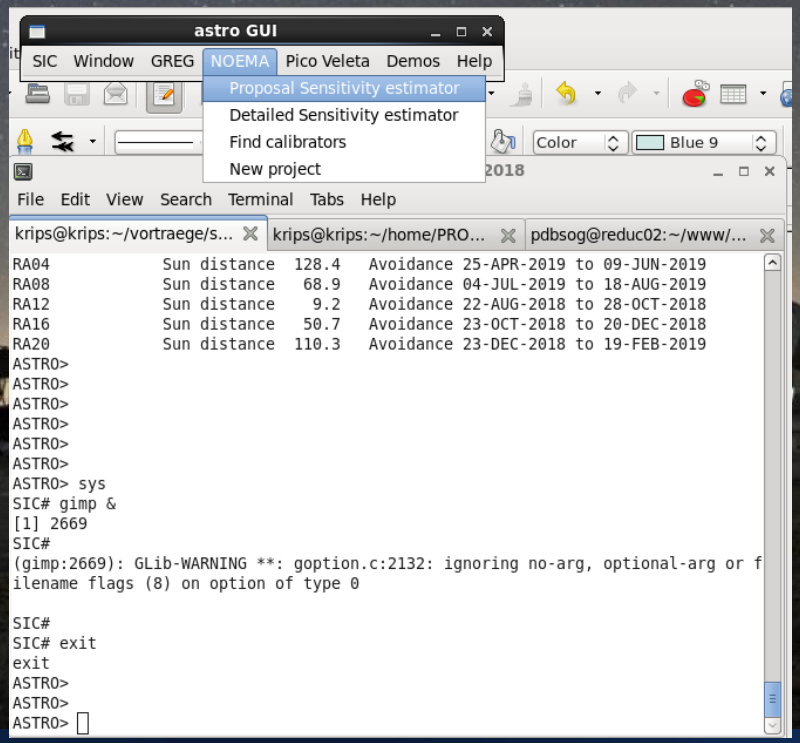
A	collecting area of a single antenna ($176.7\,\mathrm{m}^2$)
η_a	aperture efficiency (0.80 @ 3mm, 0.75 @ 2mm, 0.65 @ 1mm)
η_j	instrumental decorrelation $\eta_j = e^{-\sigma_j^2/2}$ (0.90 to 0.98)
η_C	correlator efficiency ($\eta_C = 0.88$)
k	Boltzmann constant
$\langle T_{sys} \rangle$	average system temperature [K]
η_p	atmospheric decorrelation $\eta_p = e^{-\sigma_p^2/2}$ (0.6 to 0.98)
N	Number of antennas (currently 10)
$\delta\nu$	Spectral Bandwidth [Hz] (62.5 kHz to 15.6 GHz)
$t_{\rm on}$	On-source integration time [s], $t_{\rm obs} = 1.6\, t_{\rm on}$
$N_{\rm pol}$	Number of polarizations (1 or 2)
$\frac{2k}{\eta_a A \cdot \eta_j \eta_C}$	$= J_{pk}$: Conversion factor Kelvin to Jansky 22 Jy/K @ 3mm, 26 Jy/K @ 2mm, 35 Jy/K @ 1mm

How to use NOEMA

Sensitivity requirements:

* How to calculate sensitivities and corresponding telescope time?

  Use either time estimator in PMS or astro:



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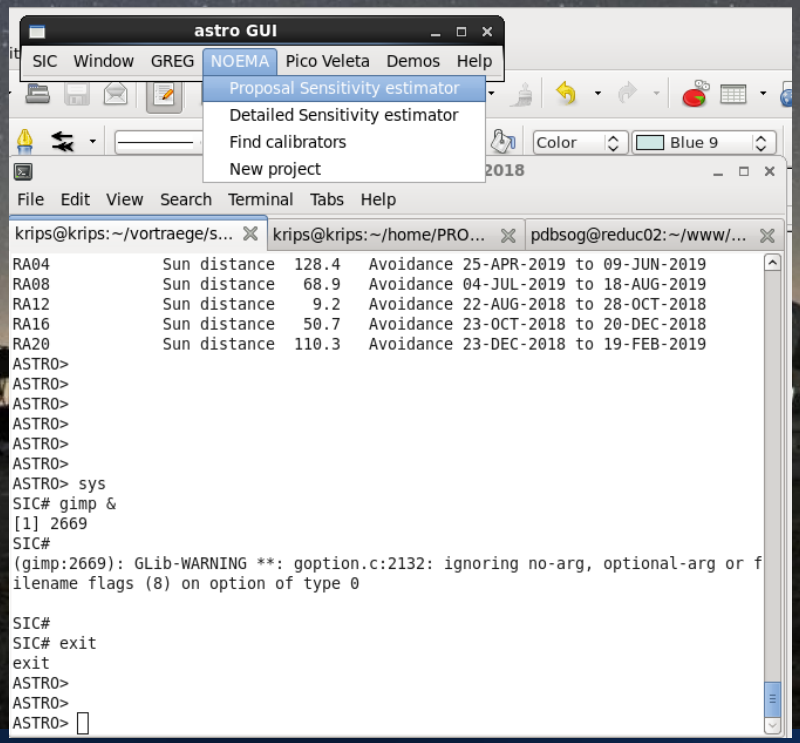
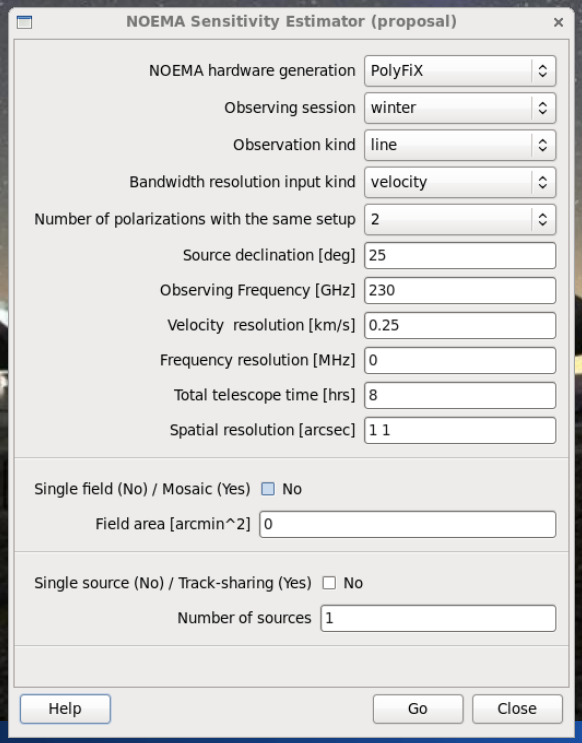
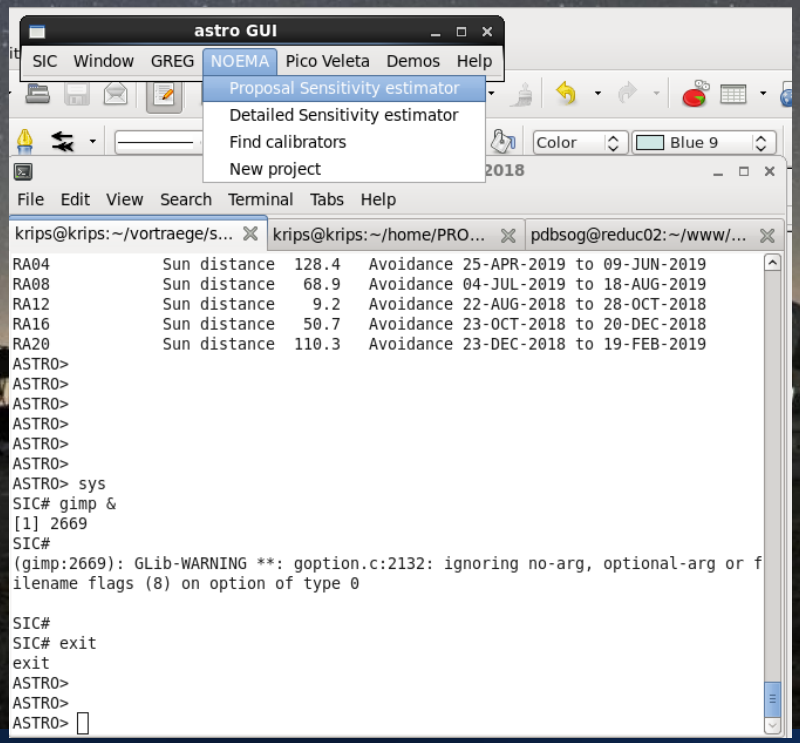


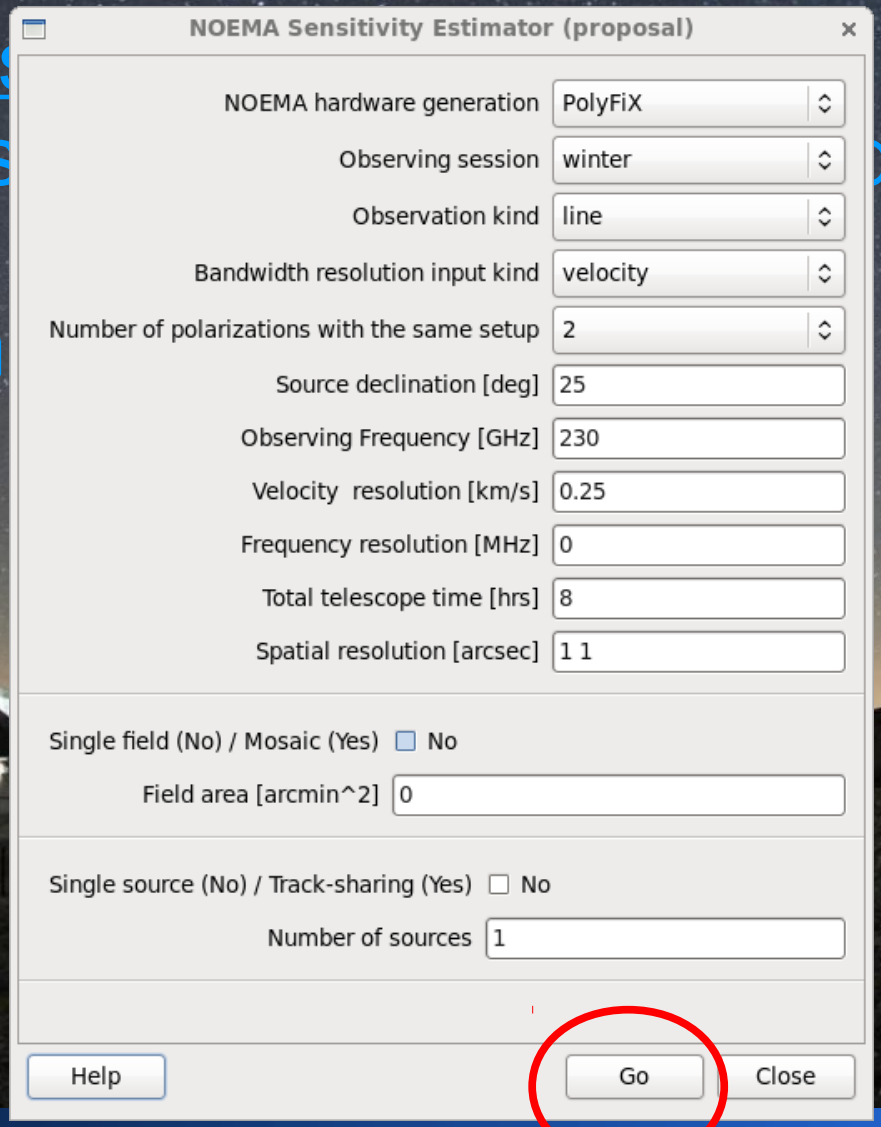


How to use NOEMA

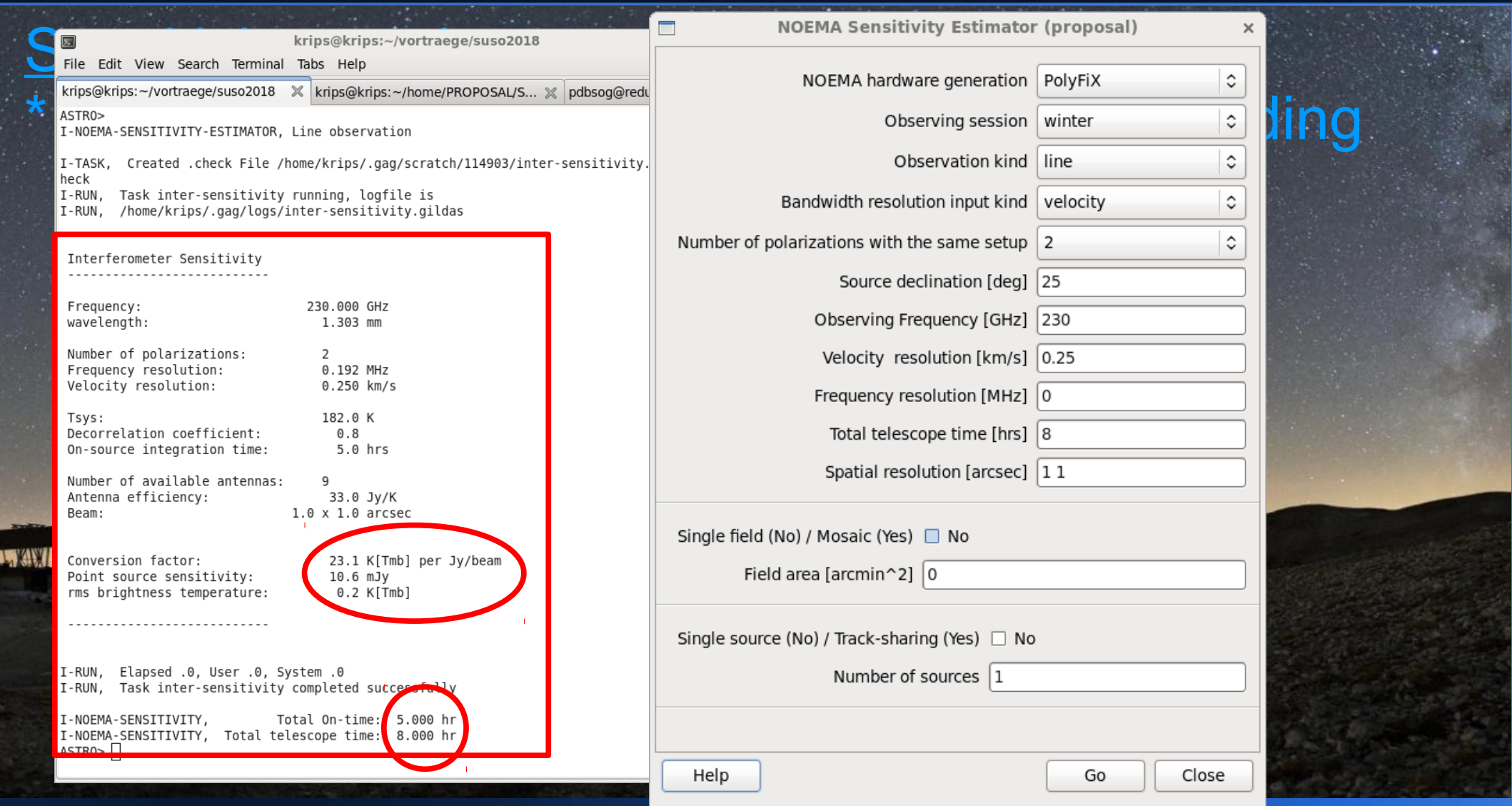
Sensitivity requirements

* How to calculate sens ding telescope time?
  Use either time estim





How to use NOEMA



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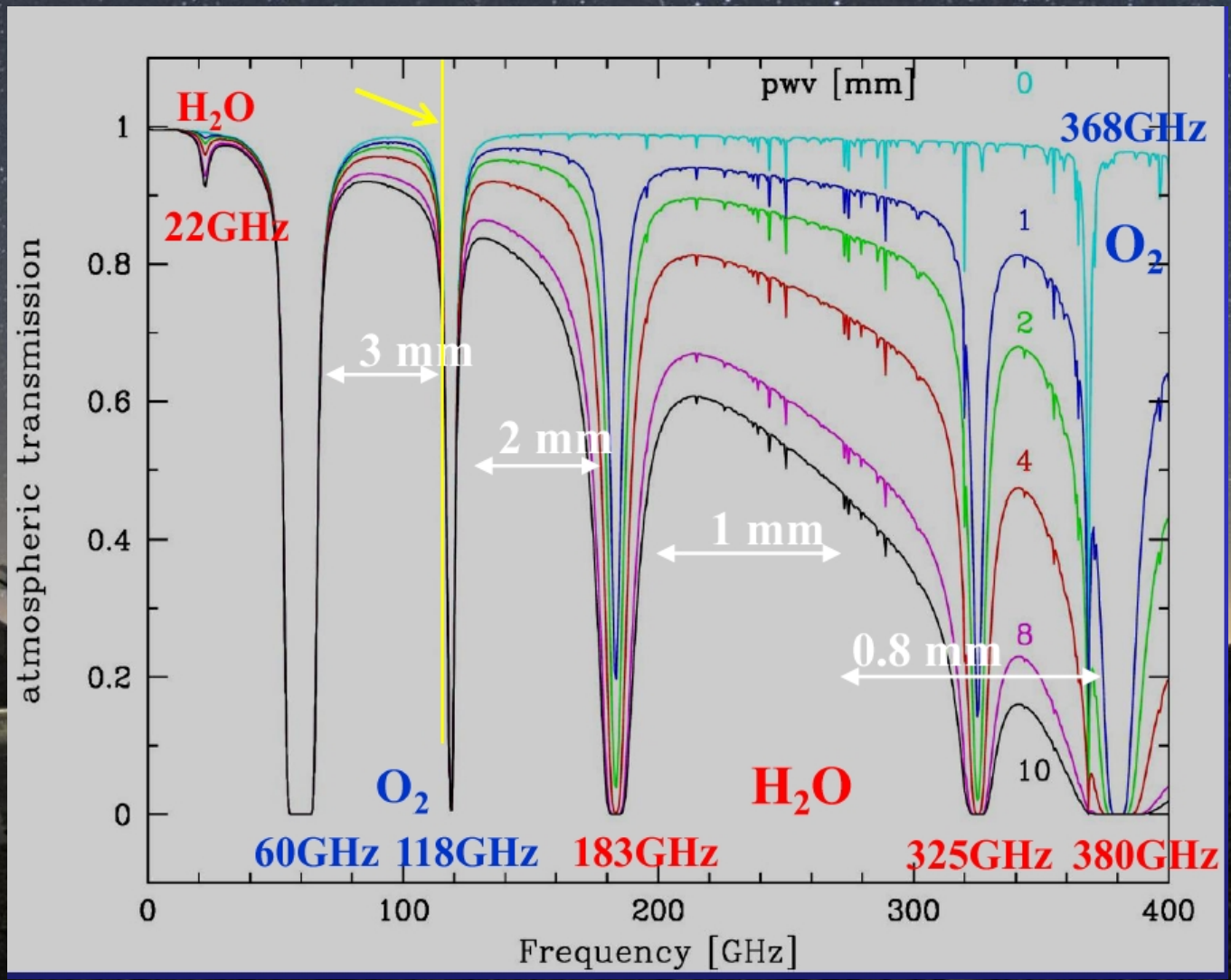


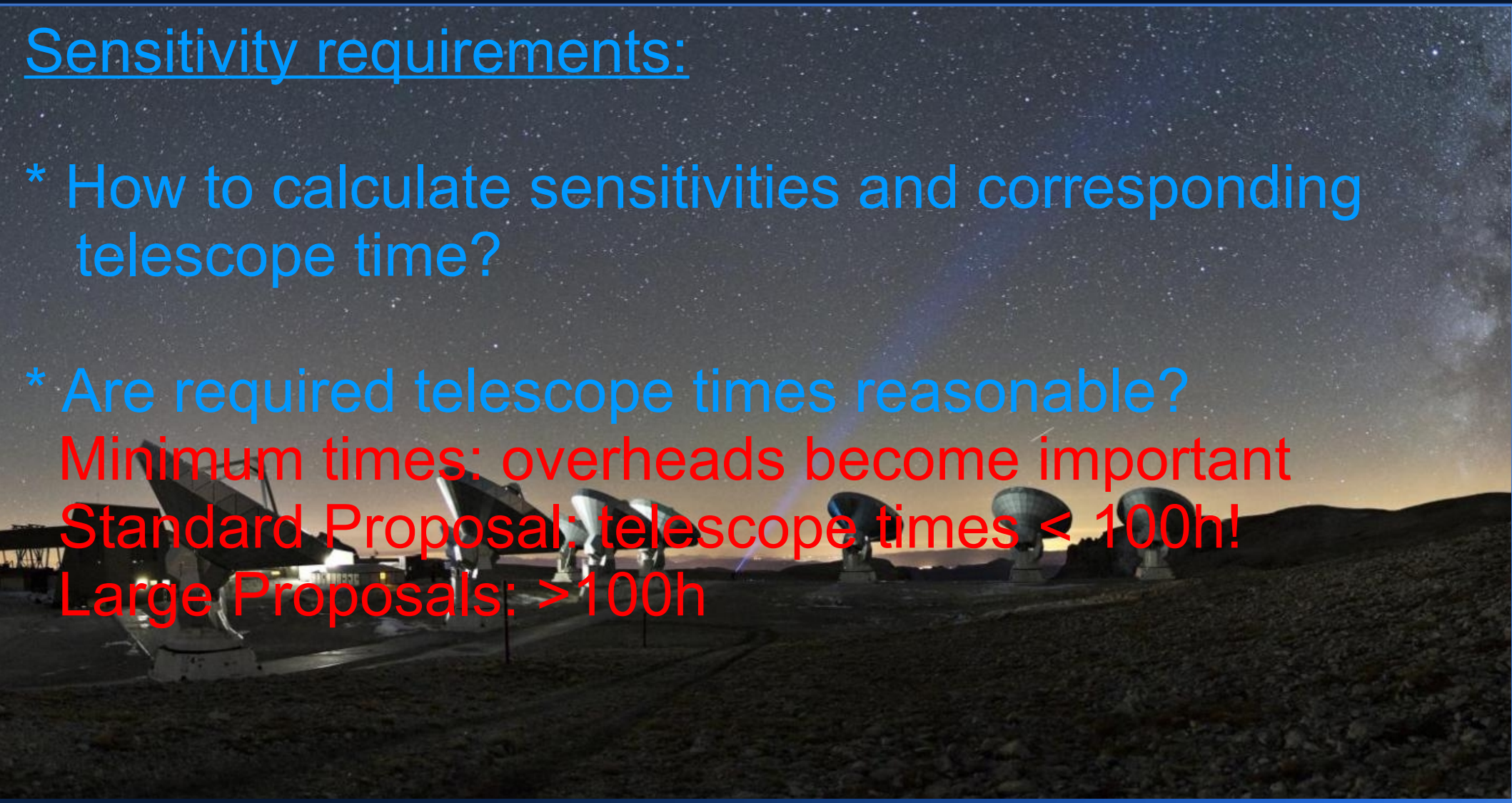
How to use NOEMA

Sensitivity requirements:

* How to calculate sensitivities and corresponding telescope time?

* Are required telescope times reasonable?

How to use NOEMA

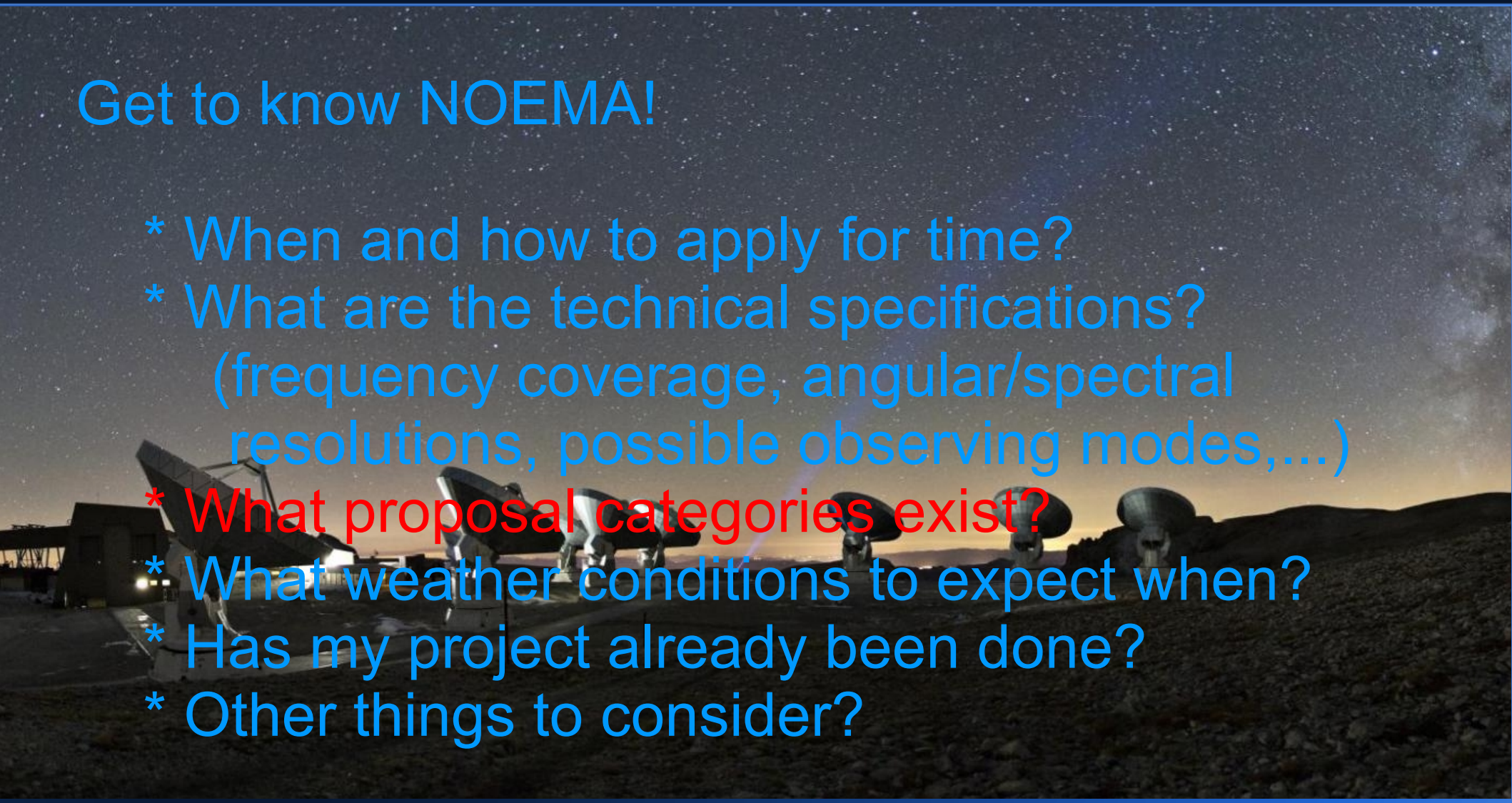


Sensitivity requirements:

* How to calculate sensitivities and corresponding telescope time?

* Are required telescope times reasonable?
  Minimum times: overheads become important
  Standard Proposal: telescope times < 100h!
  Large Proposals: >100h

How to use NOEMA



Get to know NOEMA!

* When and how to apply for time?
* What are the technical specifications? (frequency coverage, angular/spectral resolutions, possible observing modes,...)
* What proposal categories exist?
* What weather conditions to expect when?
* Has my project already been done?
* Other things to consider?

How to use NOEMA

Proposal Categories:

* Standard Proposals (<100h):
  Detection (track-sharing)
  Mapping/Mosaic

* Large Proposals (>100h)

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  Mapping/Mosaic

* Large Proposals (>100h)

Detection:
Only rms noise level deciding factor

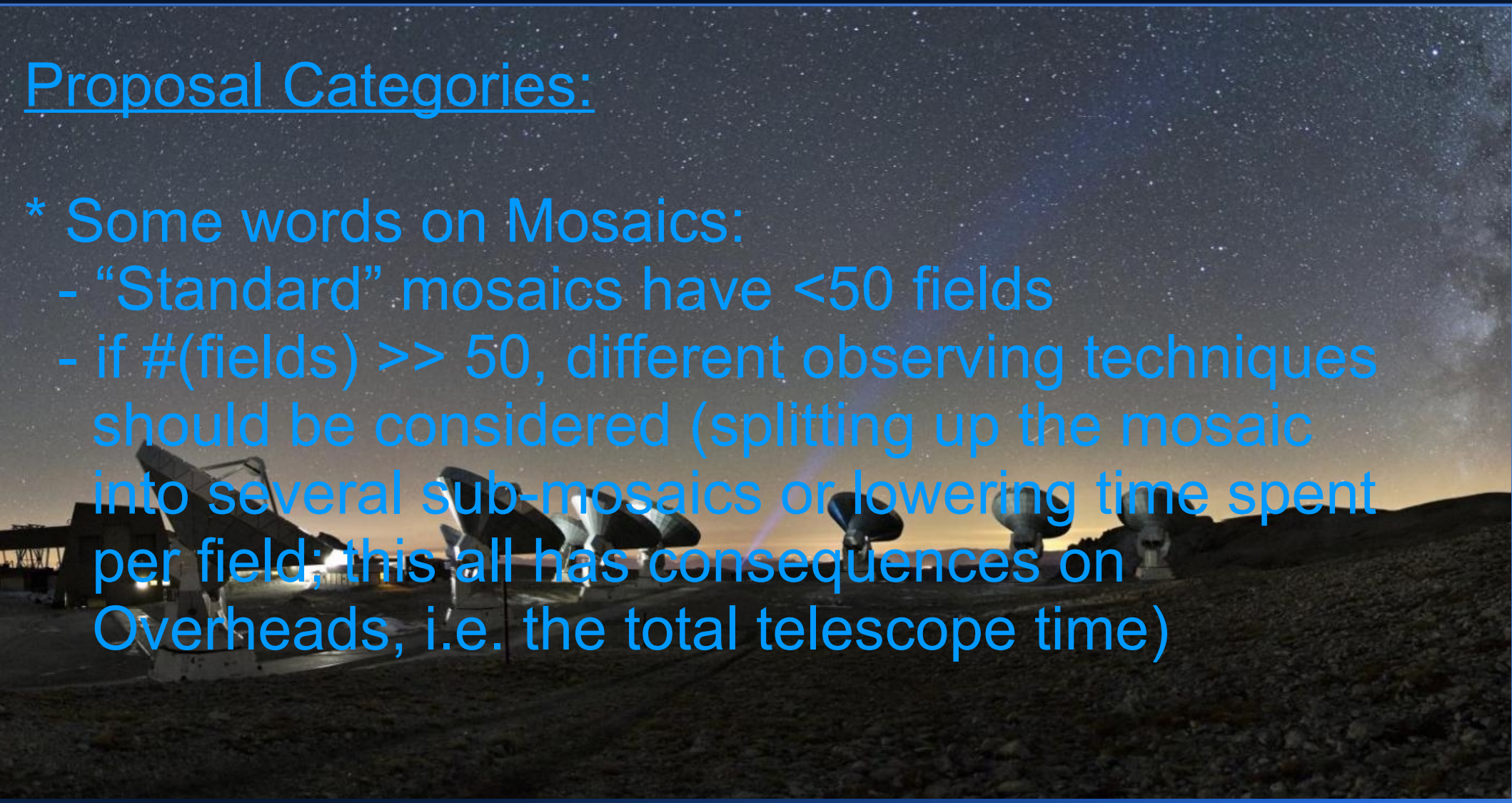
Mapping:
RMS noise level AND uv-coverage deciding factor

How to use NOEMA

Proposal Categories:

* Some words on track-sharing:
  If sources are close enough (<15deg)!
  AND
  You need rather short integration times on them!
  (#(sou)*tobs(sou) ≤ 1-2 tracks (à ~8h)

How to use NOEMA



Proposal Categories:

* Some words on Mosaics:
  - “Standard” mosaics have <50 fields
  - if #(fields) >> 50, different observing techniques should be considered (splitting up the mosaic into several sub-mosaics or lowering time spent per field; this all has consequences on Overheads, i.e. the total telescope time)

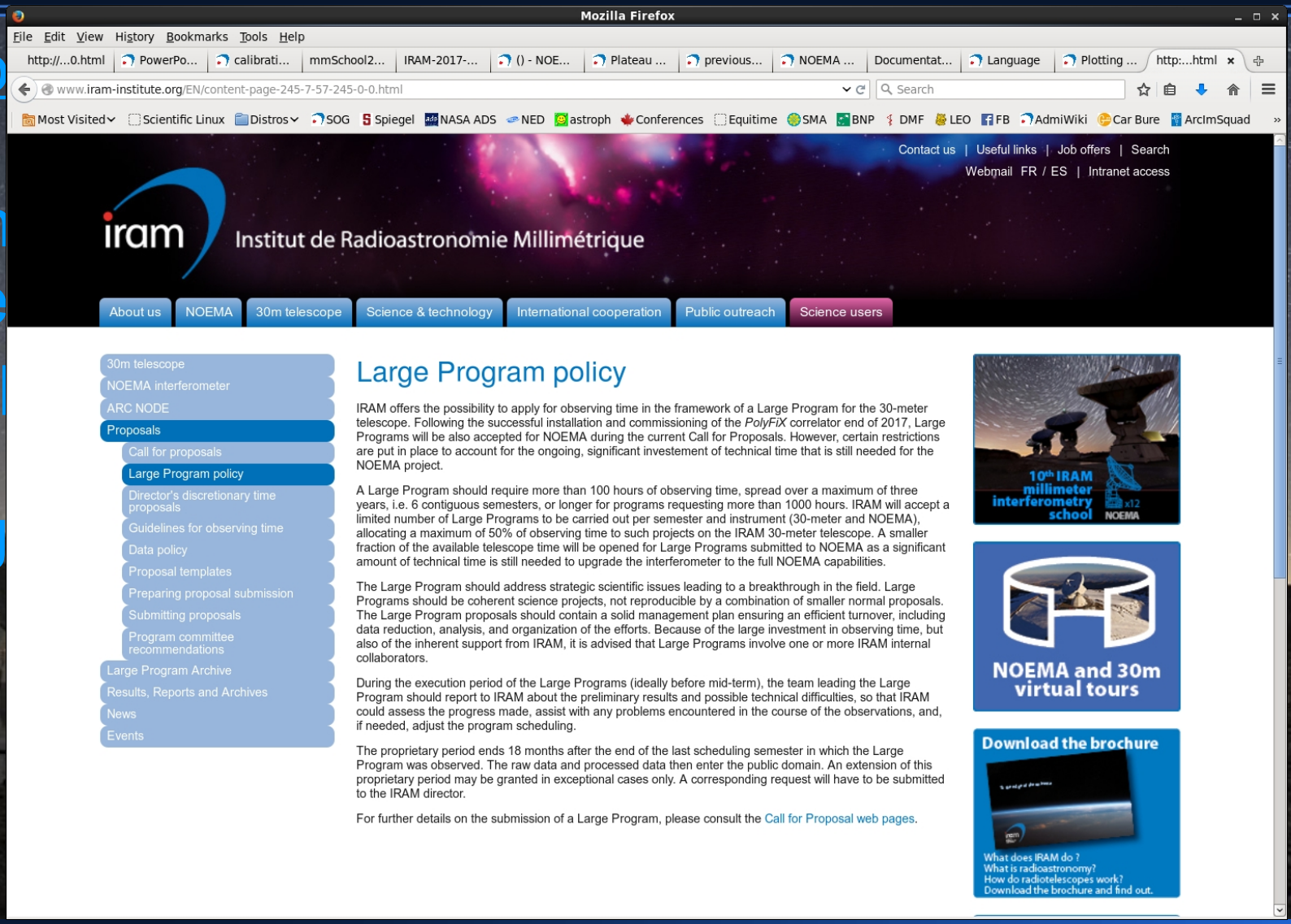
How to use NOEMA

Proposal Categories:

* Standard Proposals (<100h):
  Detection (track-sharing)
  Mapping/Mosaic

* Large Proposals (>100h):

How to use NOEMA



How to use NOEMA

Proposal Categories:

* Standard Proposals (<100h):
  Detection (track-sharing)
  Mapping/Mosaic

* Large Proposals (>100h):

* DDT/ToO projects:
  If urgent in time!
  Decision will be taken by the director

How to use NOEMA

Proposa

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* DDT/To
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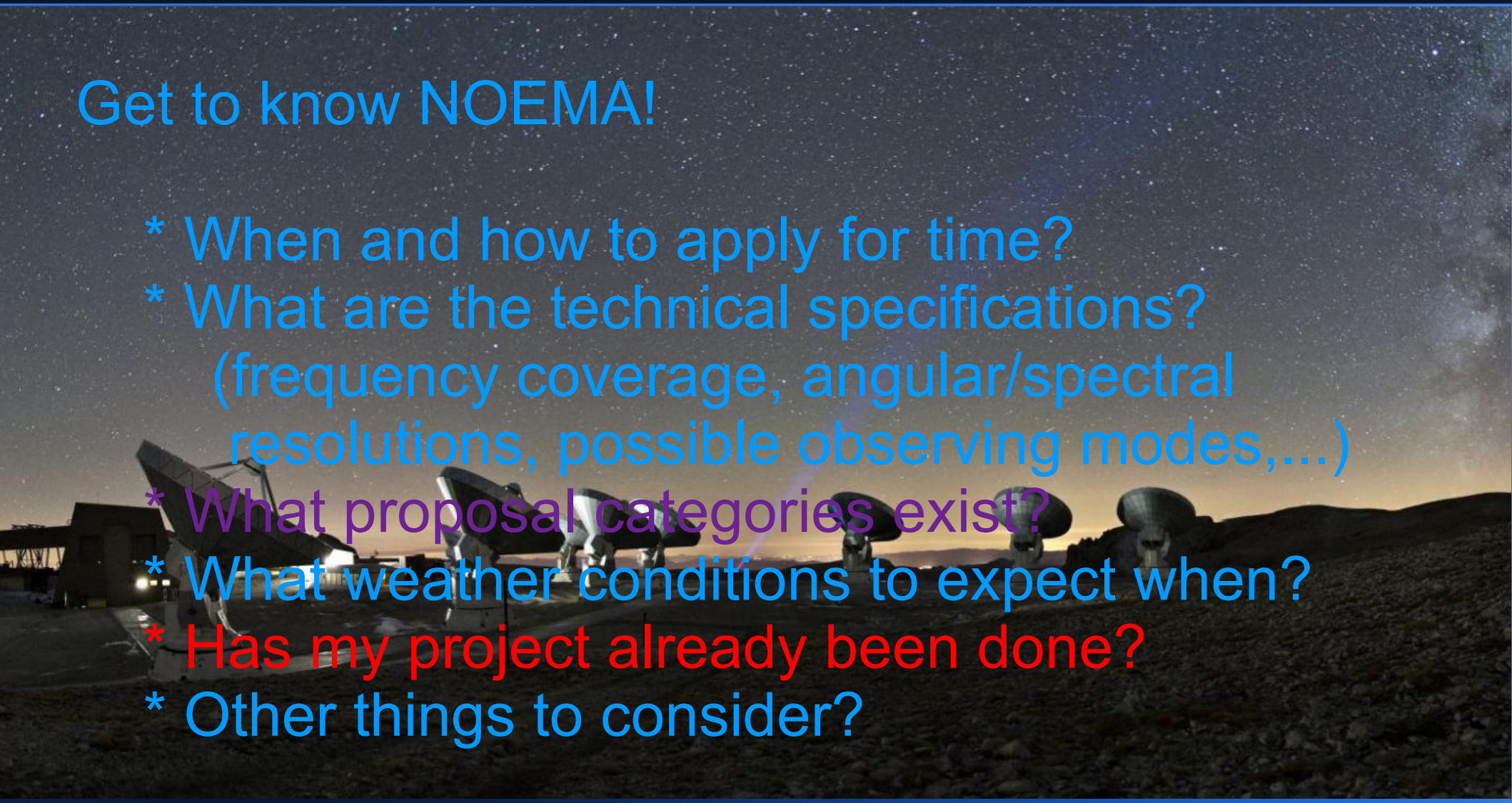
How to use NOEMA



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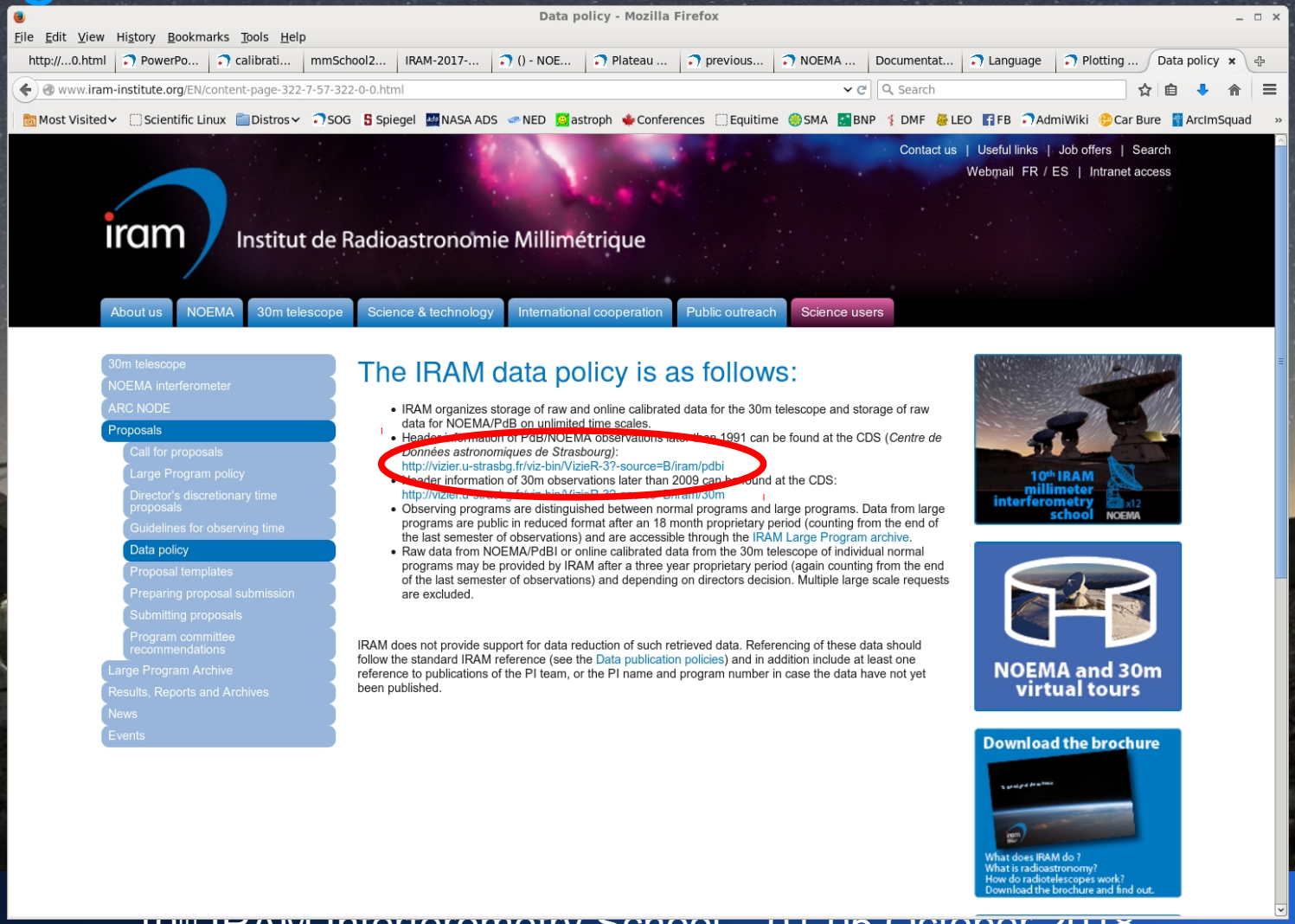
How to use NOEMA

Checking Archival Data:



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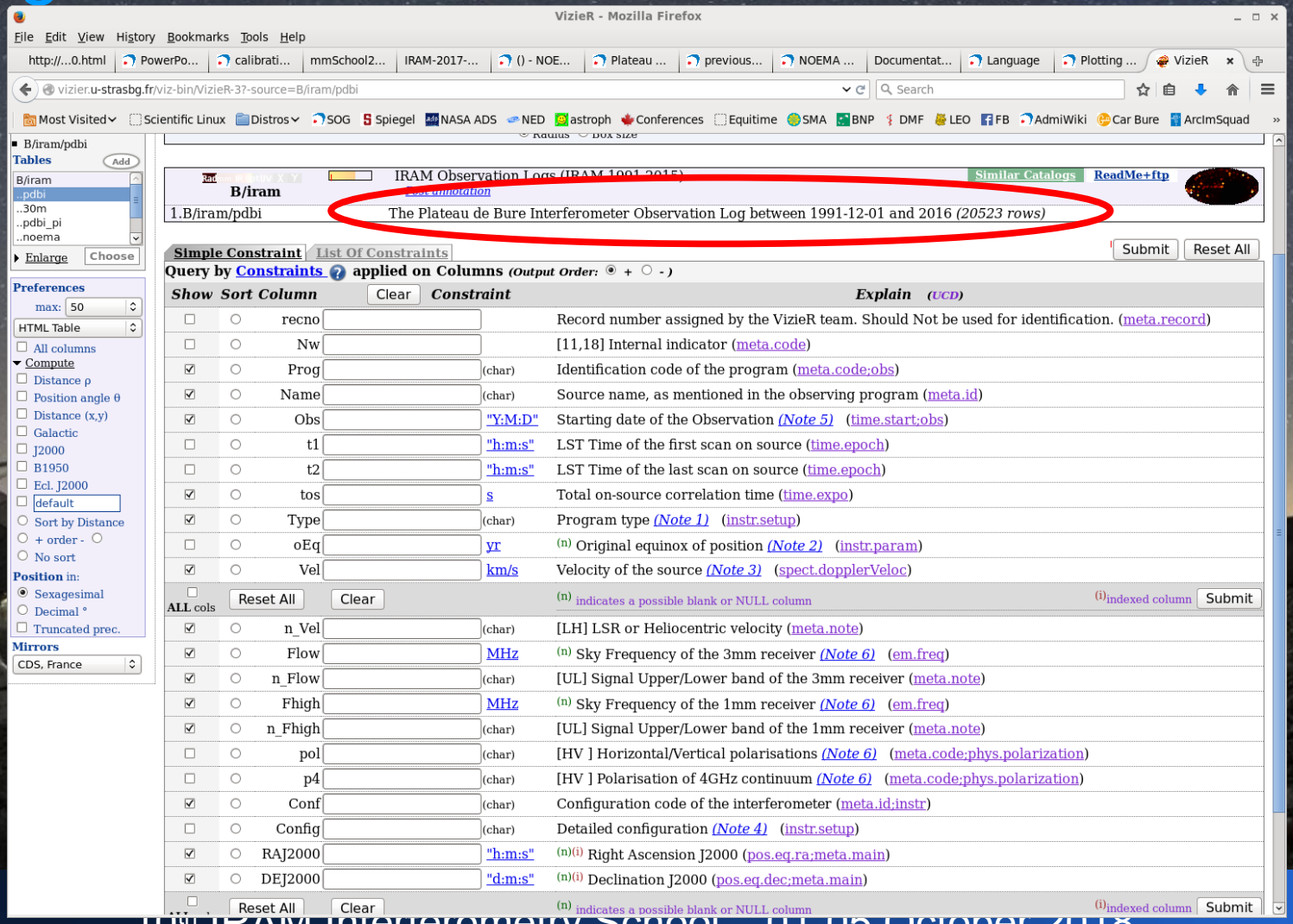

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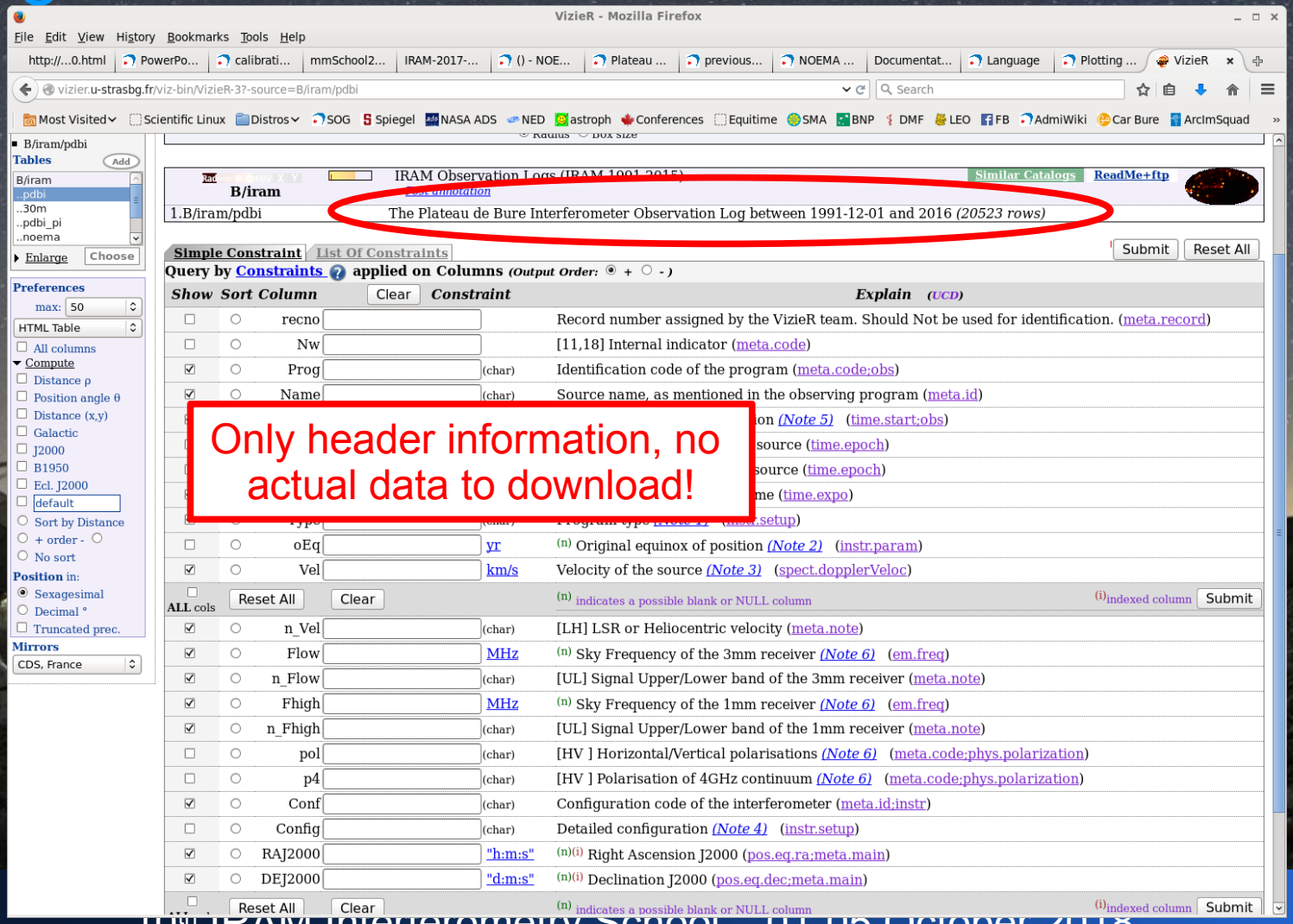

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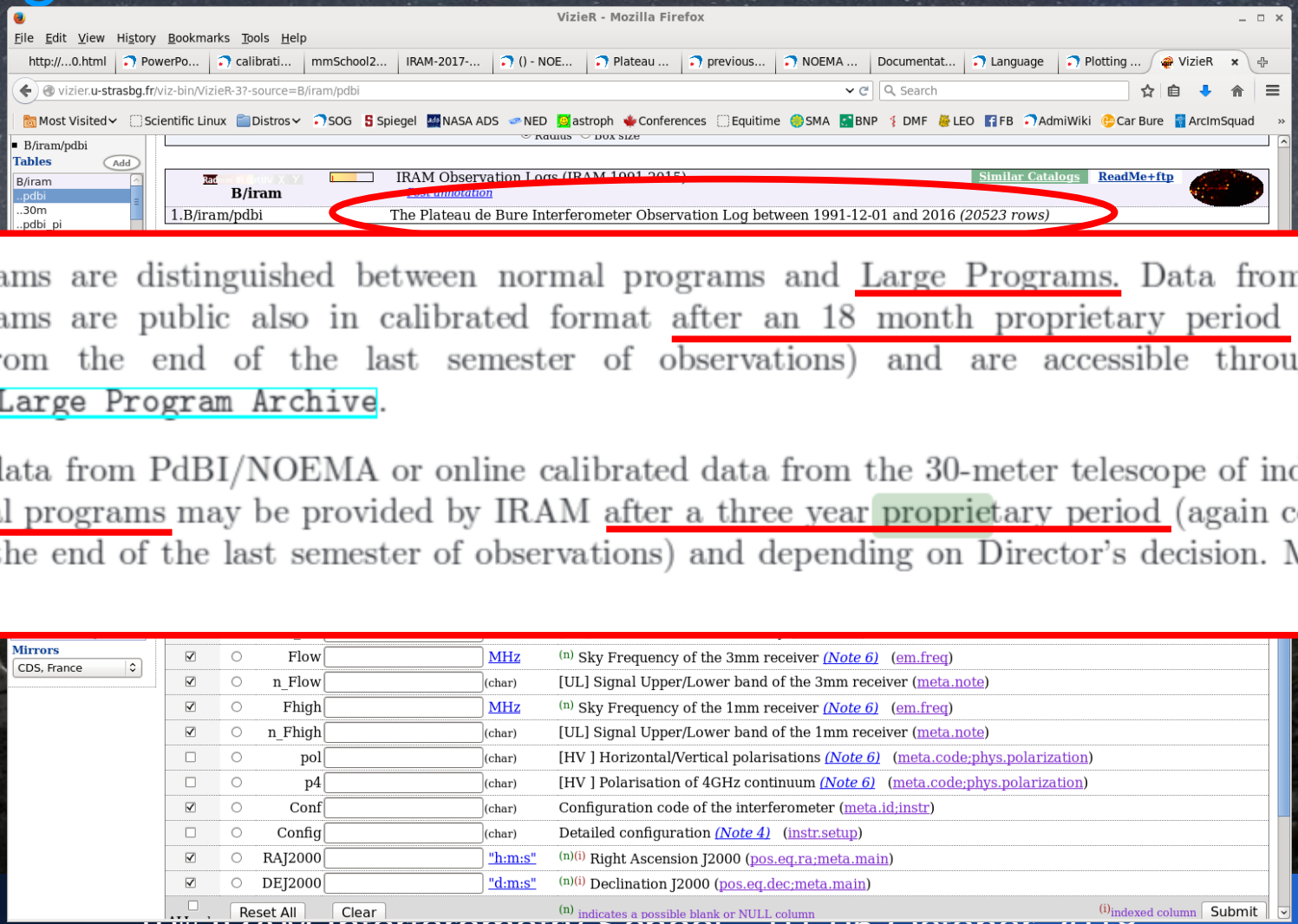

- Programs are distinguished between normal programs and Large Programs. Data from Large Programs are public also in calibrated format after an 18 month proprietary period (counting from the end of the last semester of observations) and are accessible through the IRAM Large Program Archive.
- Raw data from PdBI/NOEMA or online calibrated data from the 30-meter telescope of individual normal programs may be provided by IRAM after a three year proprietary period (again counting from the end of the last semester of observations) and depending on Director's decision. Multiple

How to use NOEMA

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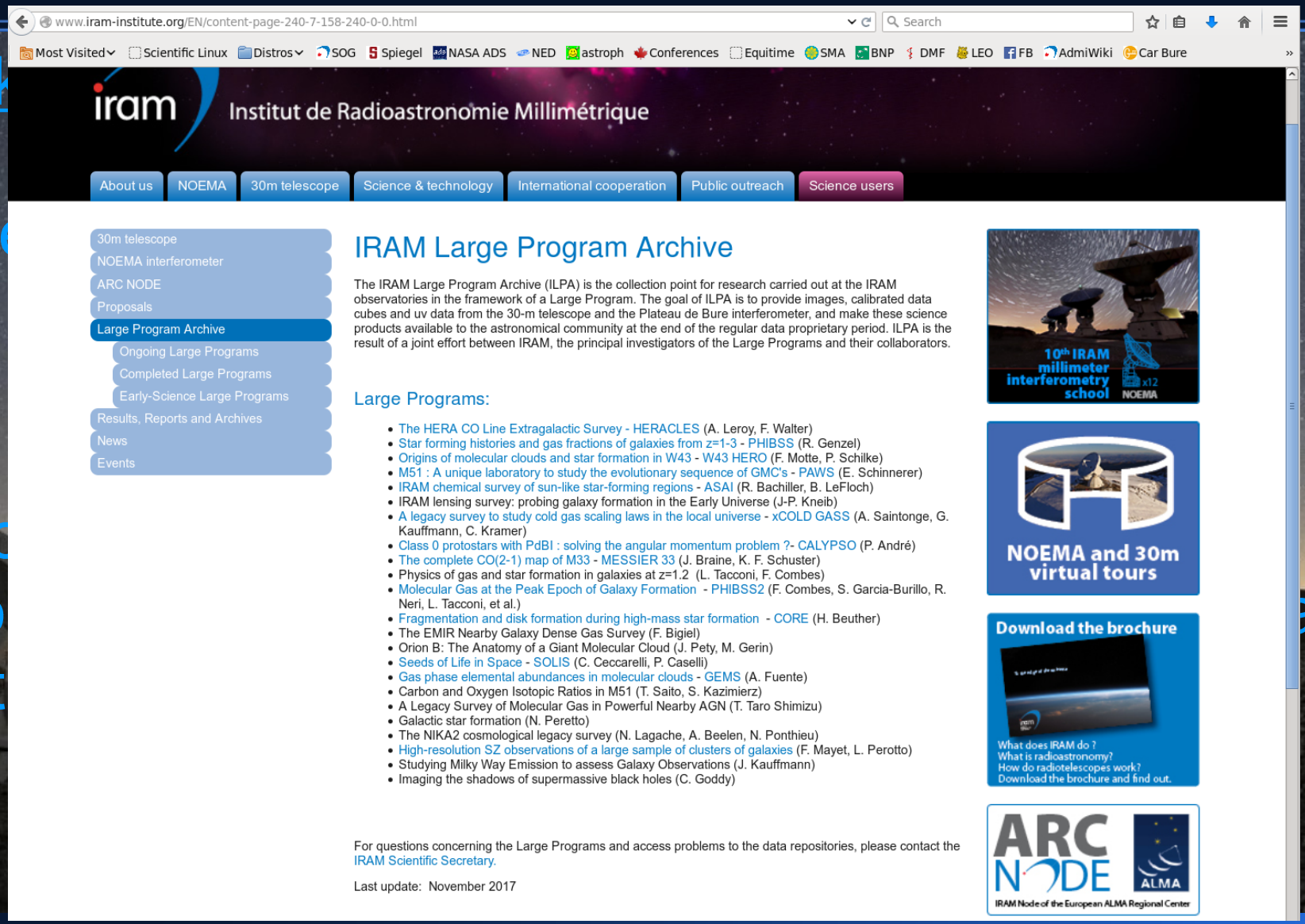

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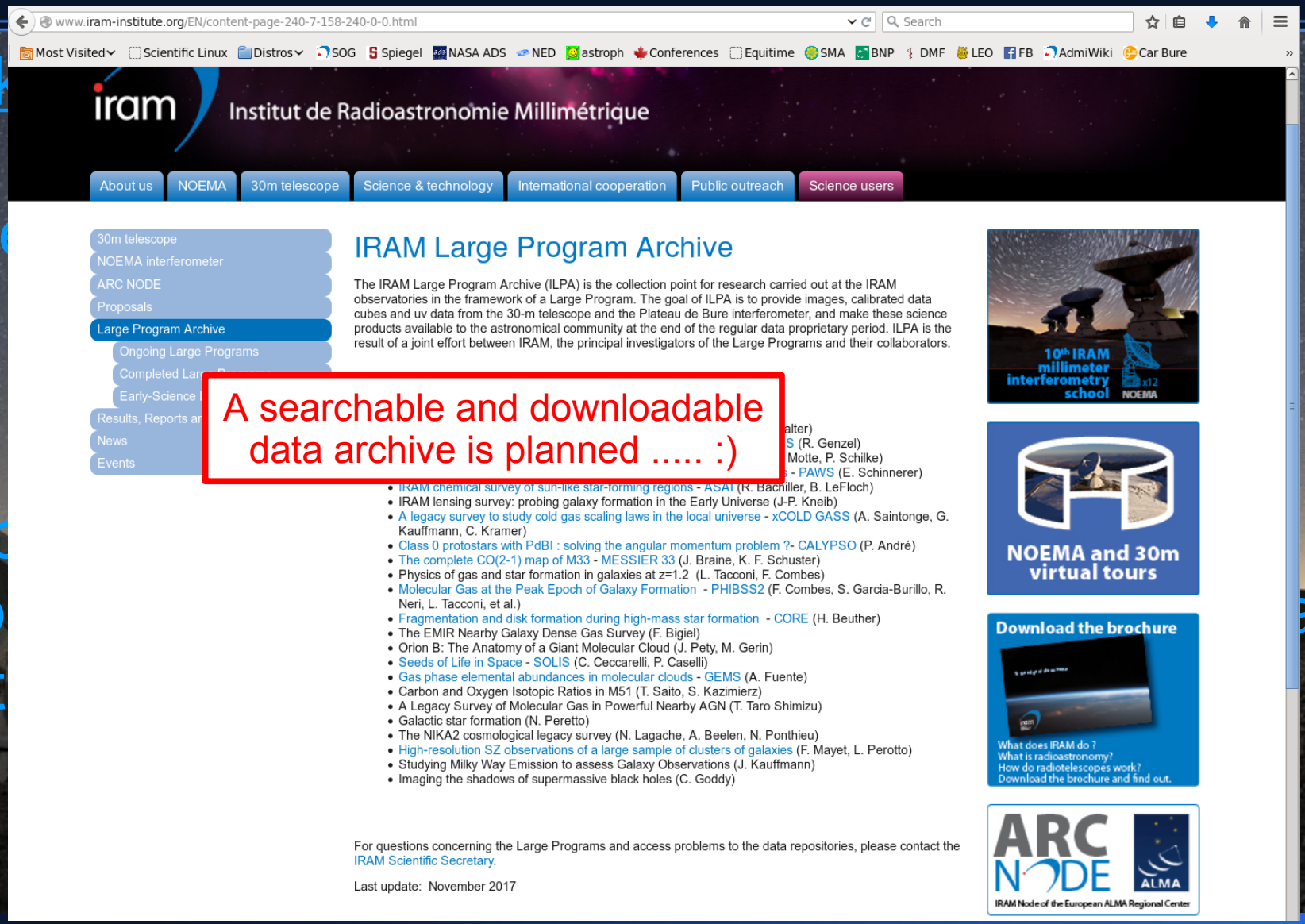


How to use NOEMA

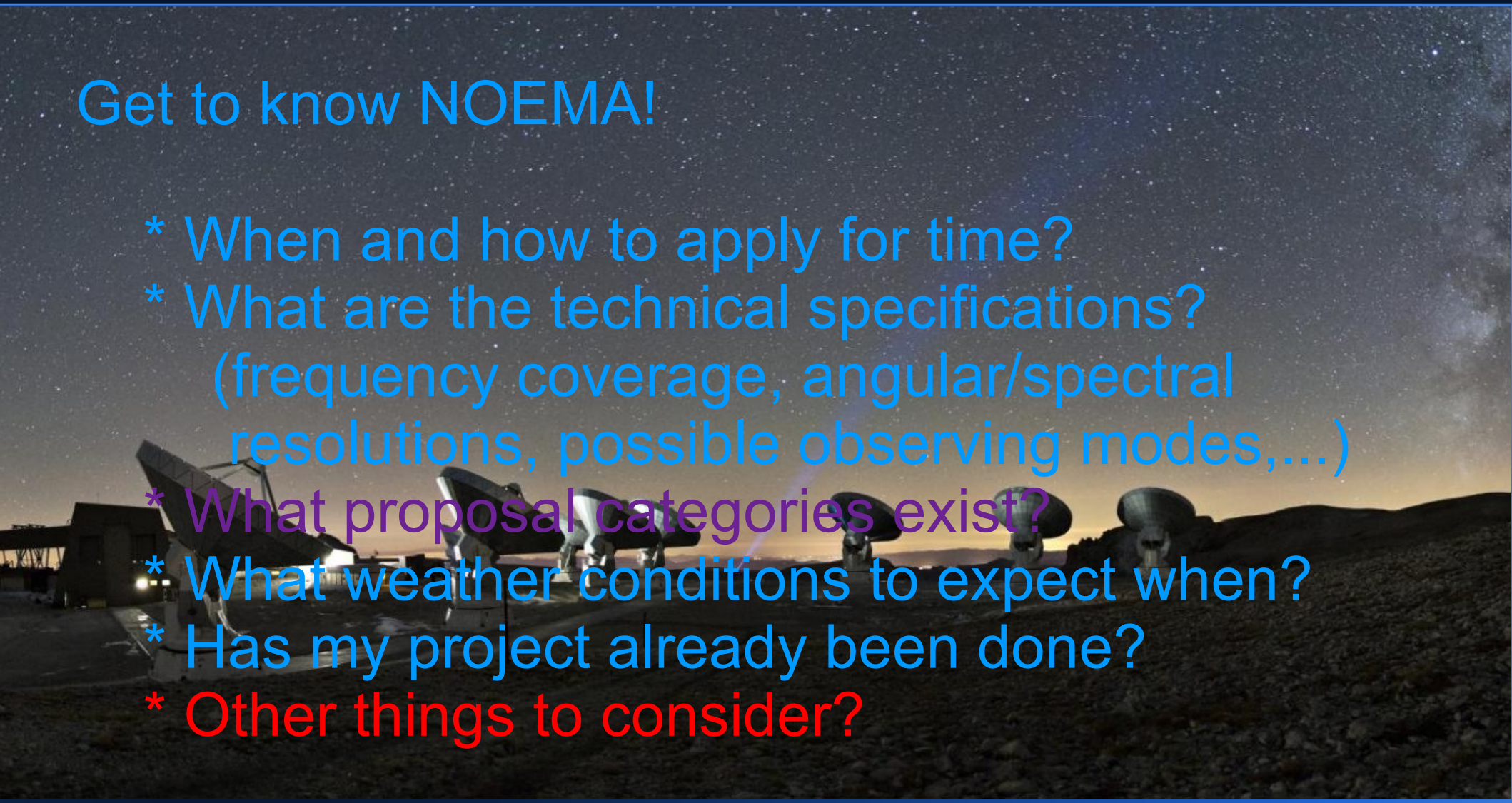
Checking Archival Data:

If observed lathen~2016, please contact us at:

sog@iram.fr

Duplication checks will be done during the proposal reviewing process, also against proposal from the same semester

How to use NOEMA



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How to use NOEMA

What else to consider?

* Non-standard projects:
  - very large scale mosaics
  - projects requiring special calibration
    + detecting weak lines over strong continuum (~1-2% of bandpass accuracy with standard bandpass calibration; could maybe reach 0.5%)
    + special needs concerning flux calibration
* Specific dates needed for observations? Coordinated Observations?

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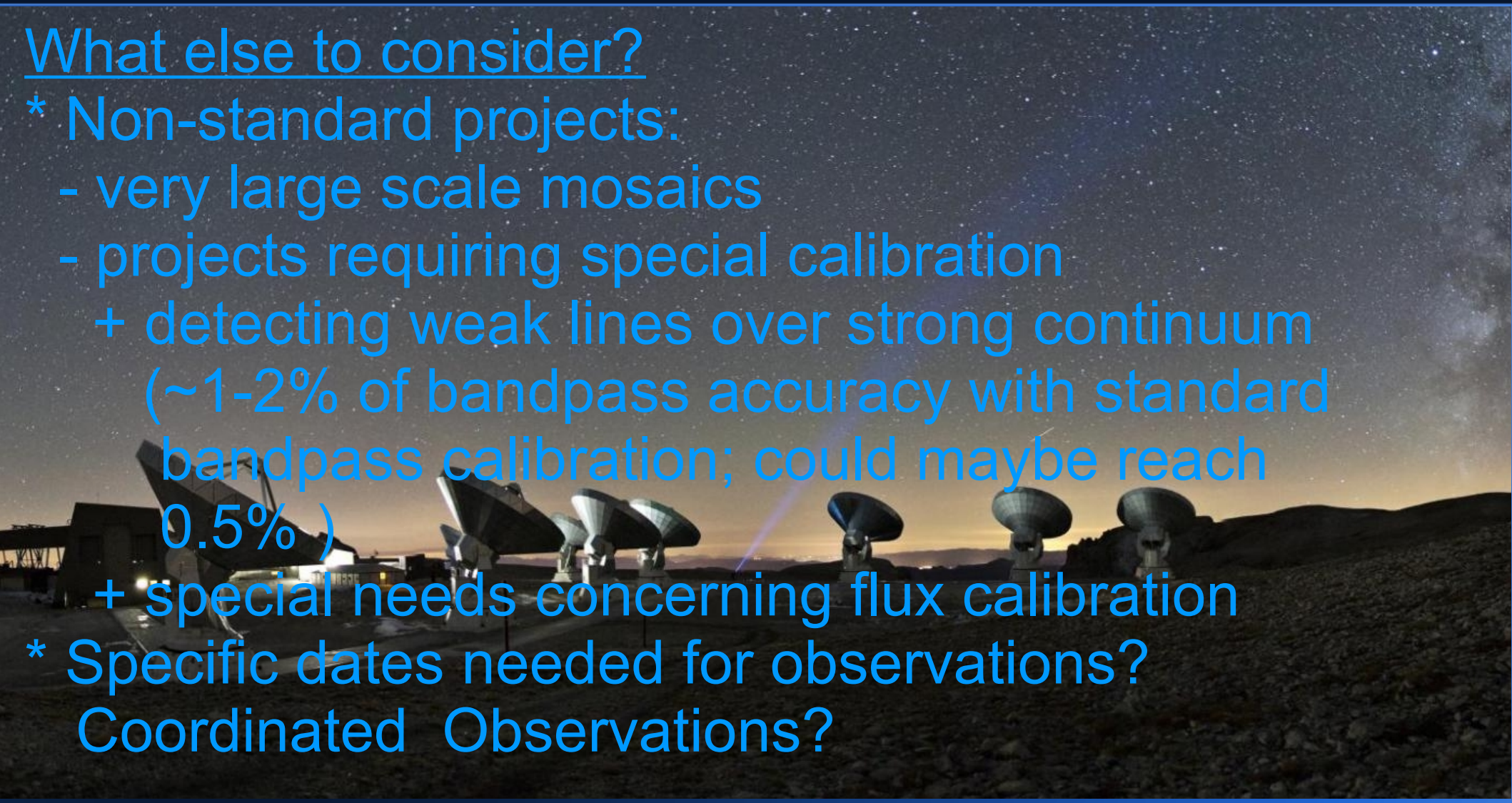
PLEASE CONTACT US IN ADVANCE AT
sog@iram.fr
TO DISCUSS OBSERVING STRATEGIES

How to use NOEMA



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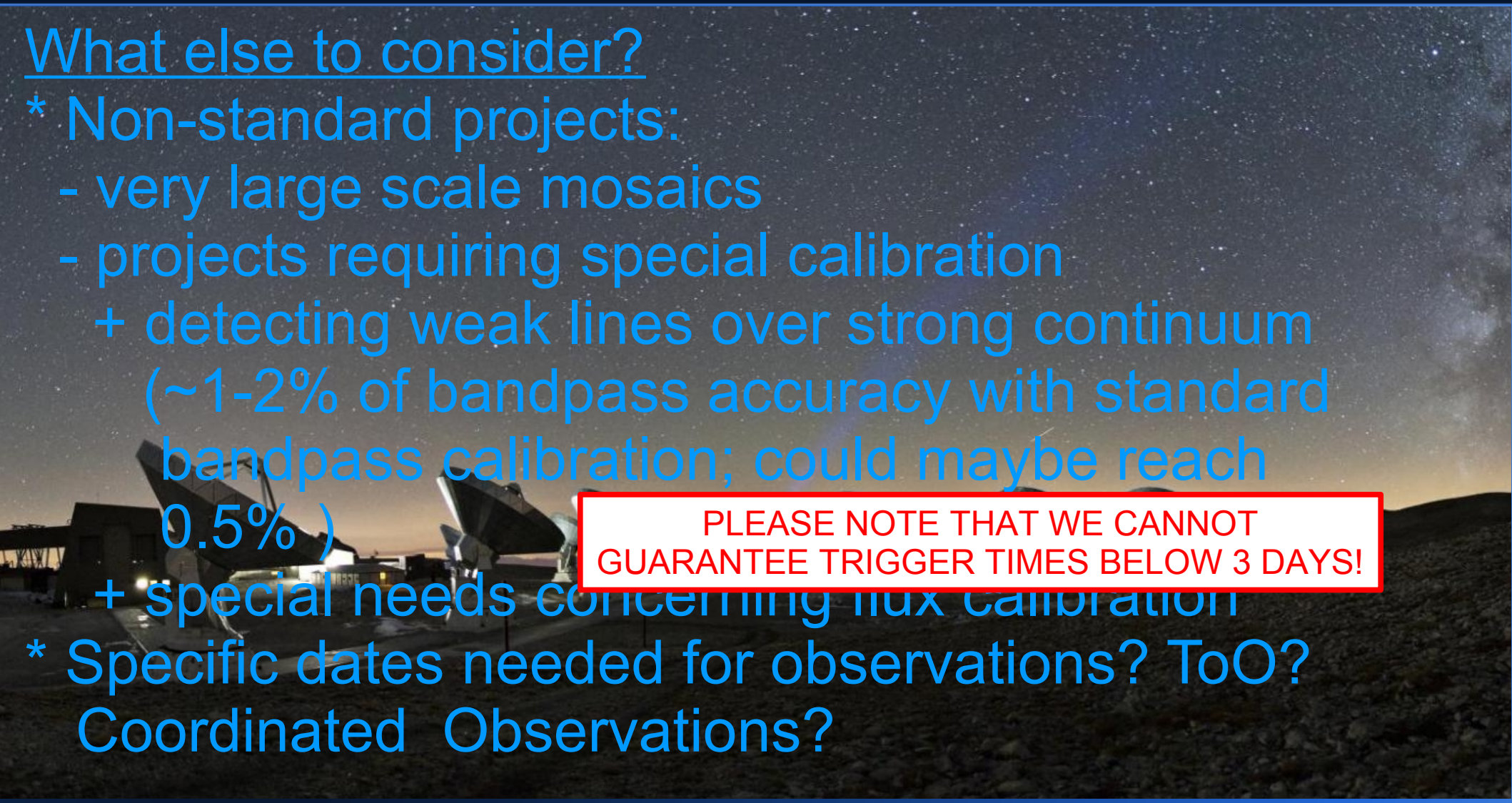

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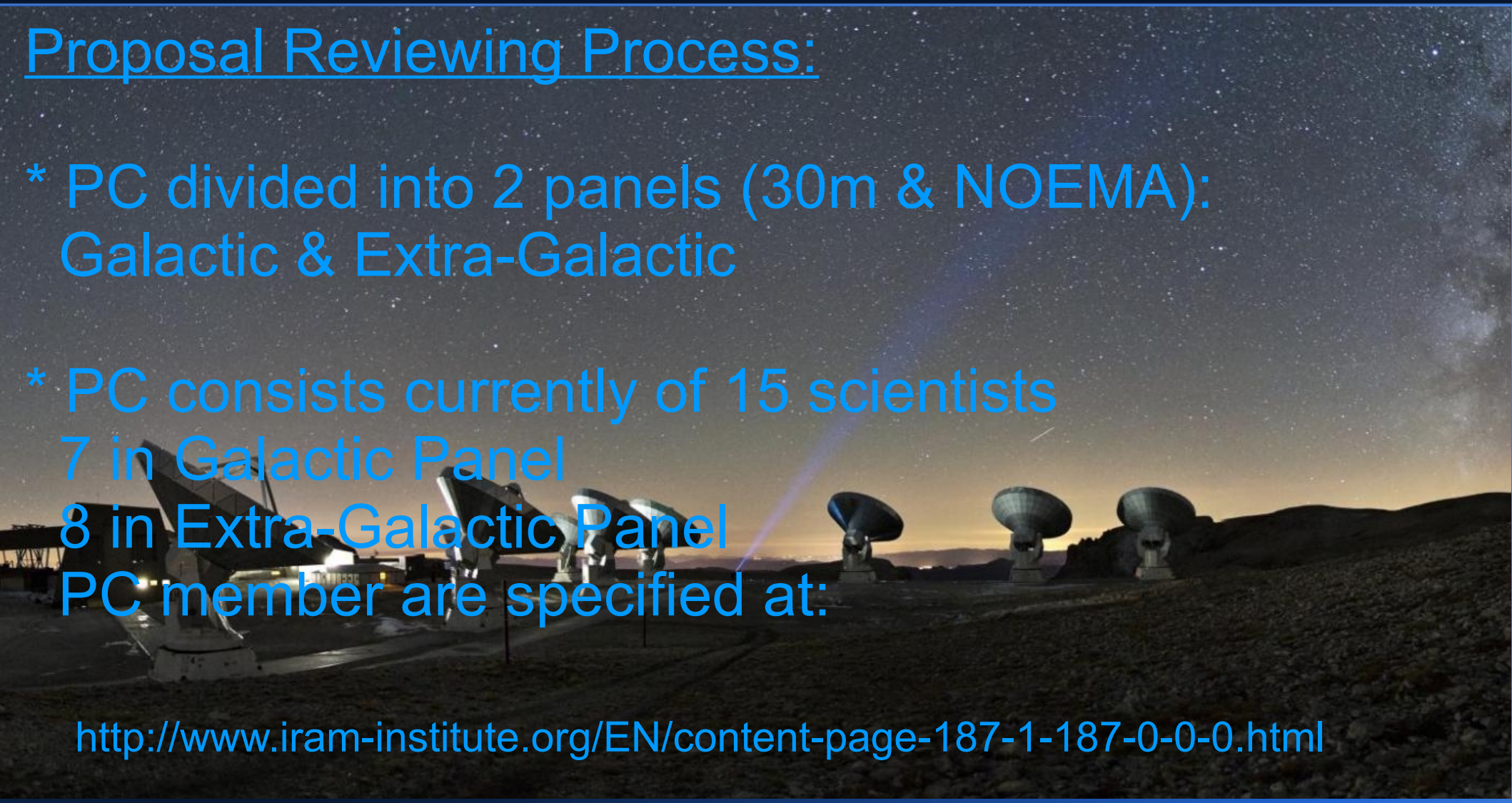
PLEASE NOTE THAT WE CANNOT GUARANTEE TRIGGER TIMES BELOW 3 DAYS!

* Specific dates needed for observations? ToO? Coordinated Observations?

How to use NOEMA

How to use NOEMA



Proposal Reviewing Process:

* PC divided into 2 panels (30m & NOEMA):
  Galactic & Extra-Galactic

* PC consists currently of 15 scientists
  7 in Galactic Panel
  8 in Extra-Galactic Panel
  PC member are specified at:

http://www.iram-institute.org/EN/content-page-187-1-187-0-0-0.html

How to use NOEMA

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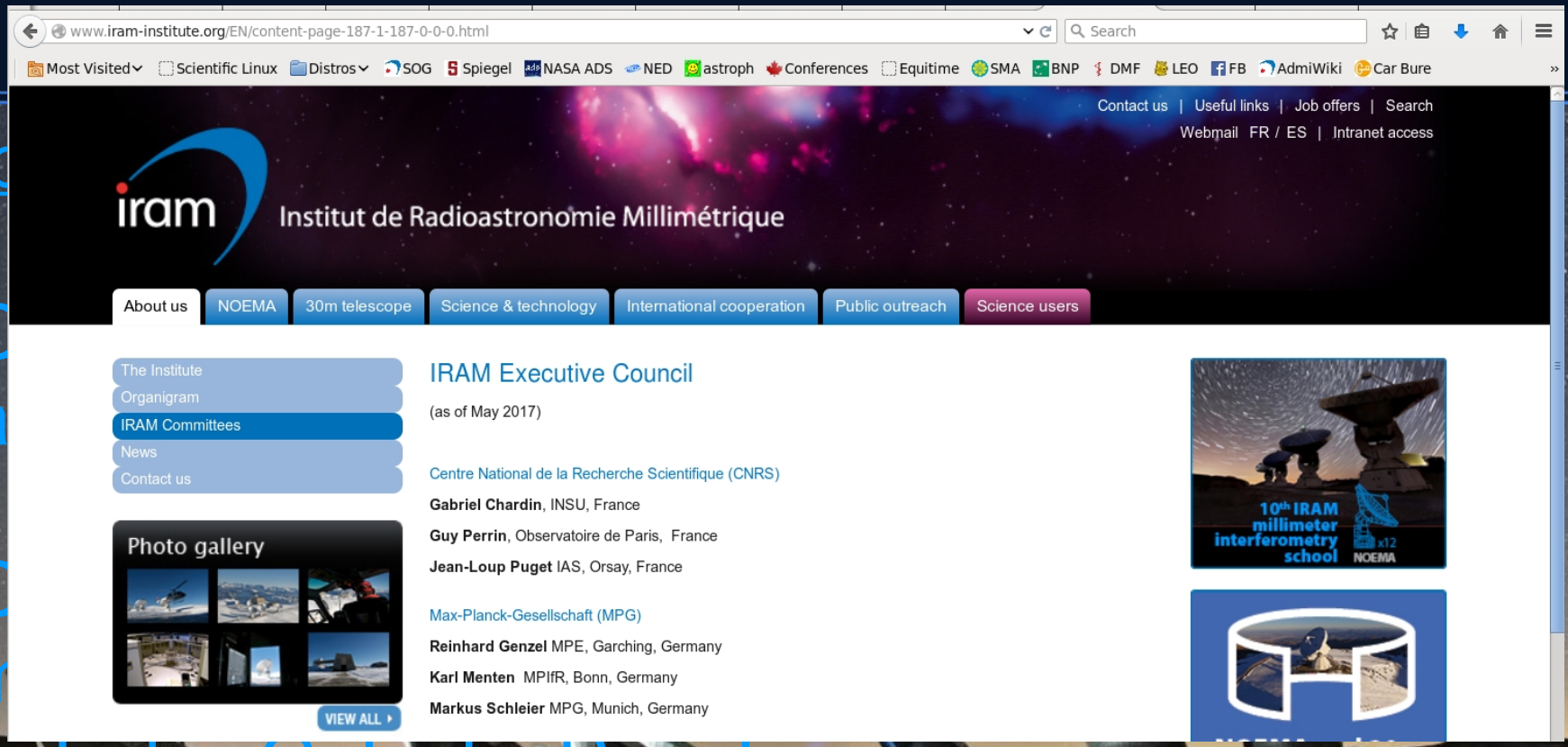
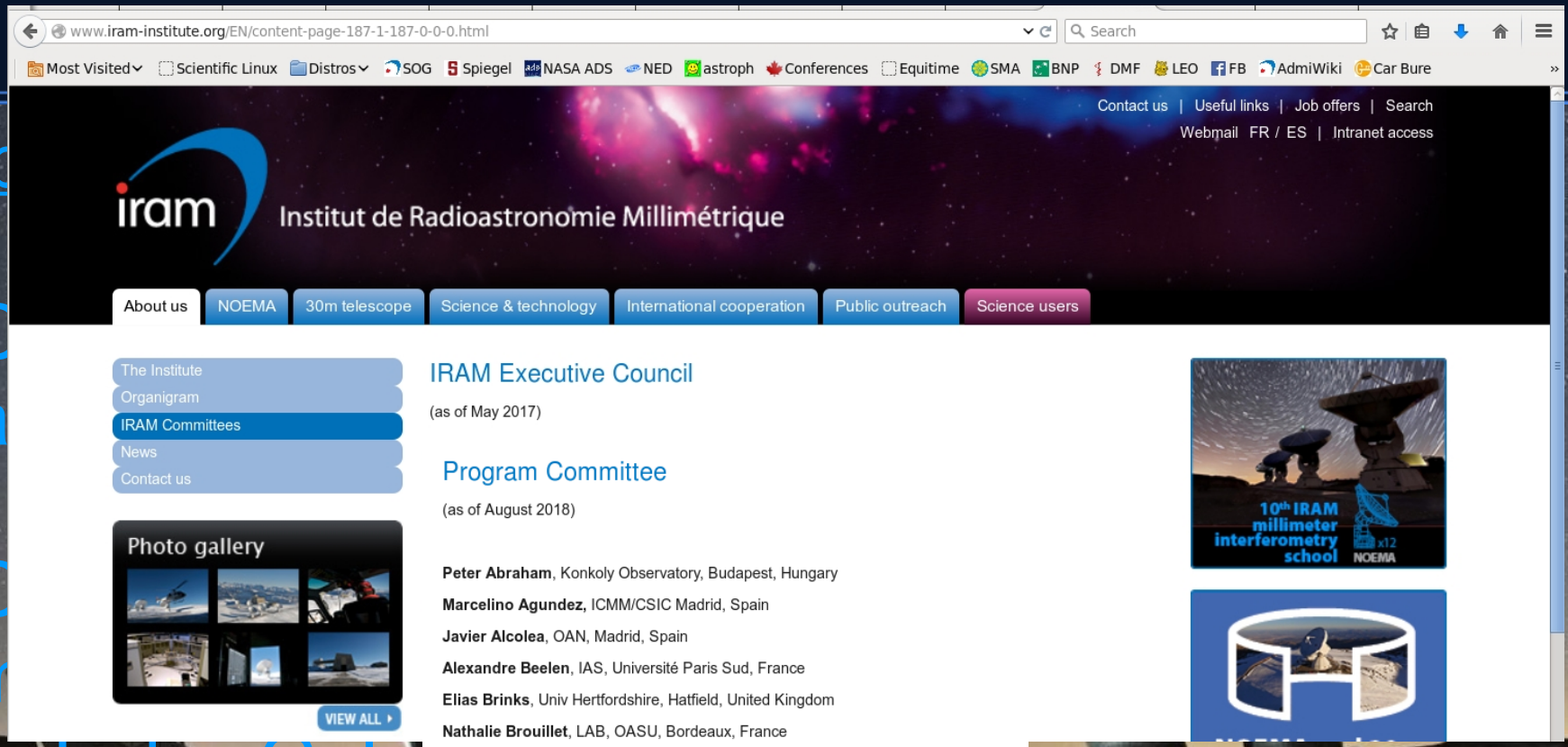
* PC
  Ga

* PC
  7 ir
  8 in Extra-Galactic Panel
  PC member are specified at:



http://www.iram-institute.org/EN/content-page-187-1-187-0-0-0.html

How to use NOEMA



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http://www.iram-institute.org/EN/content-page-187-1-187-0-0-0.html

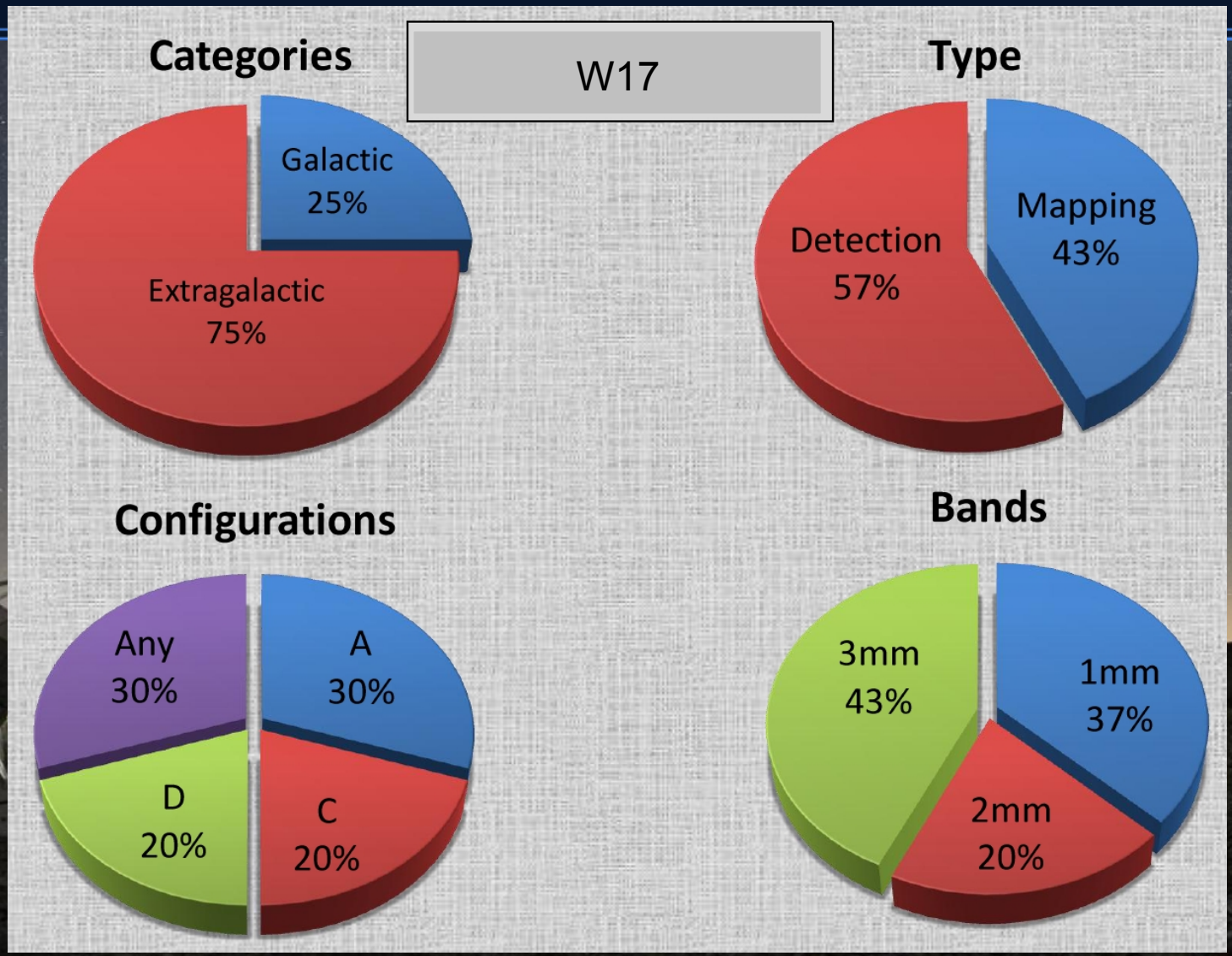
How to use NOEMA

Proposal Reviewing Process:

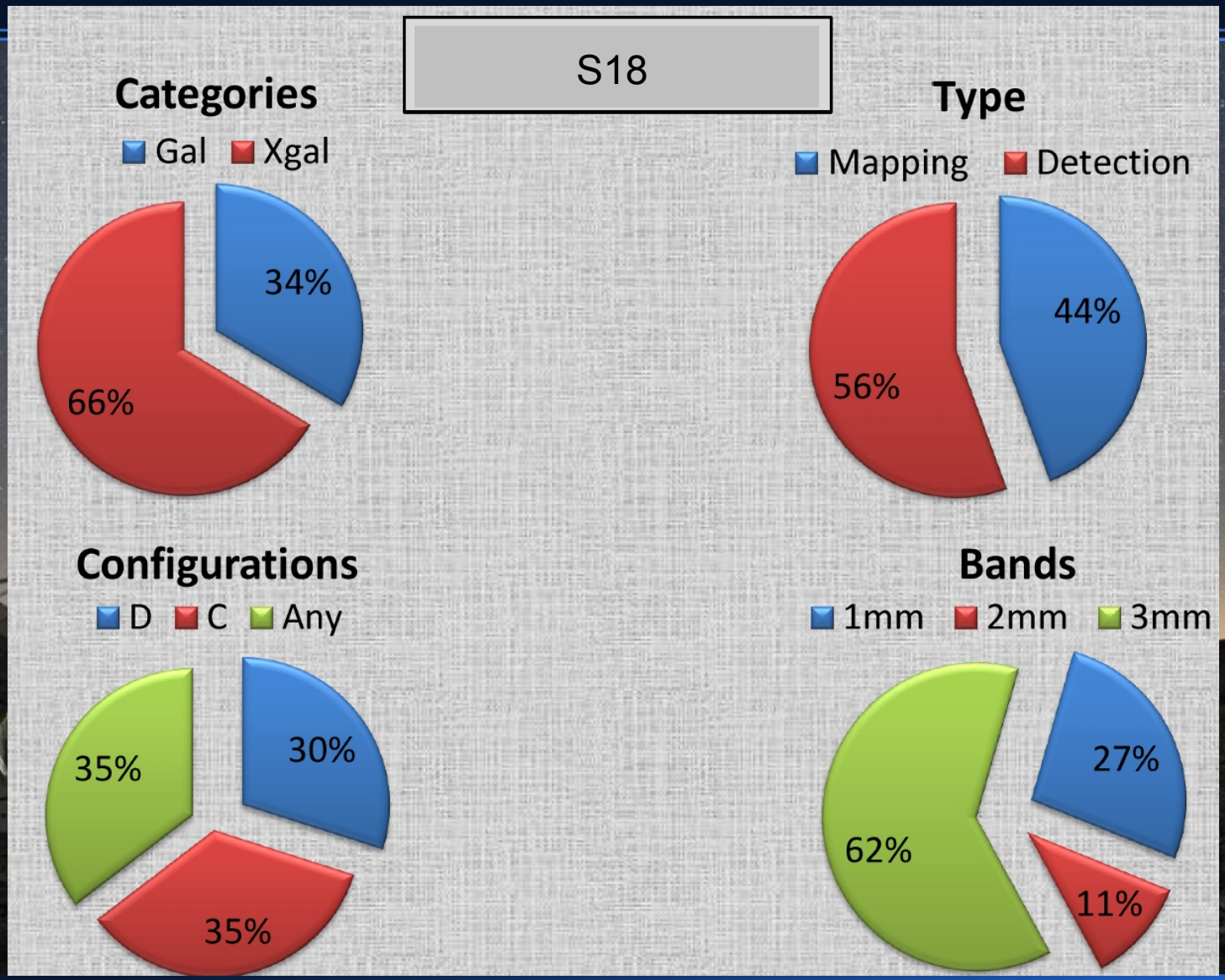
* On average, we receive ~200-250 proposals per year at NOEMA:
  Summer <100
  Winter ~130-150

* 30m receives a bit less, ~200 per year!

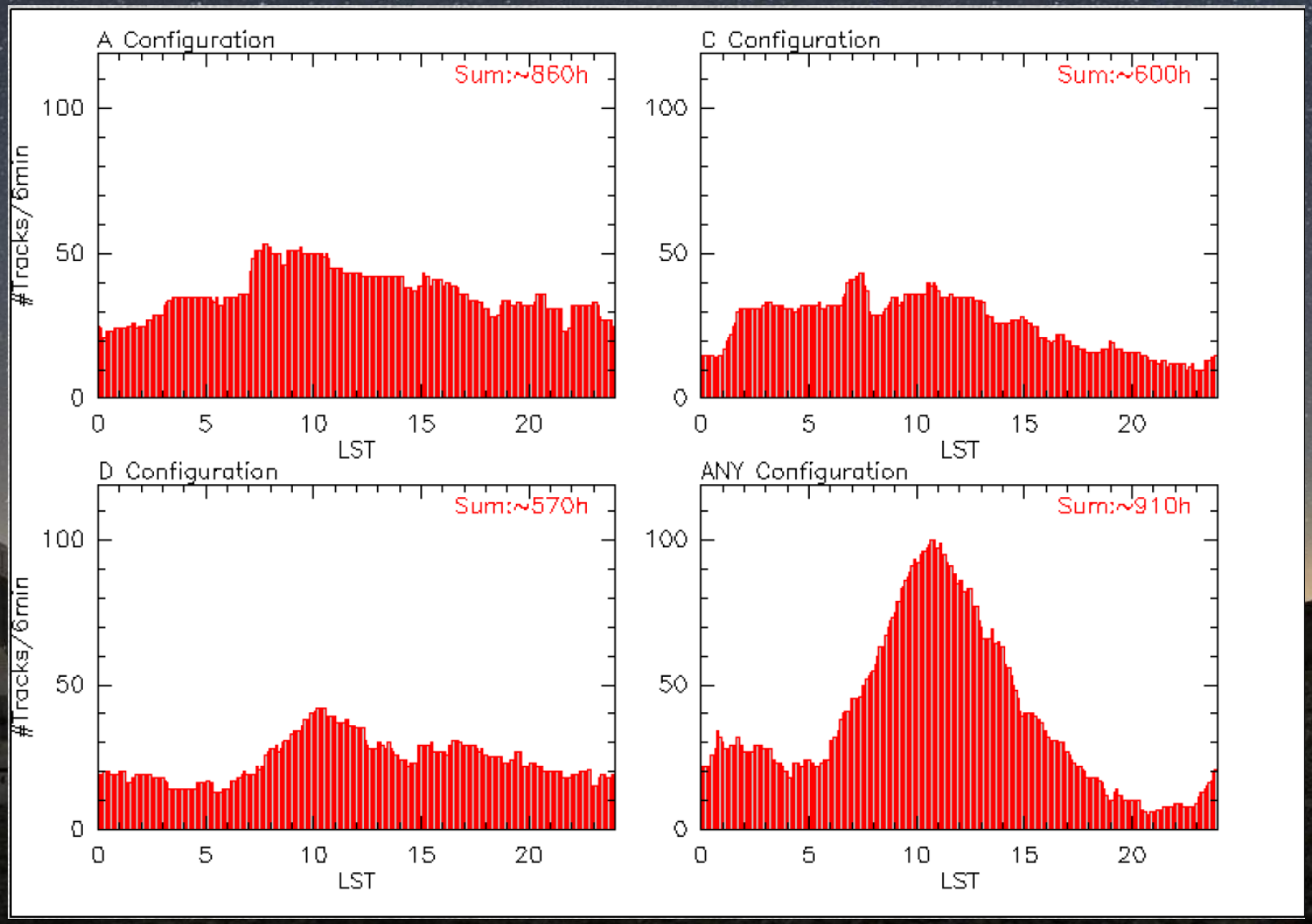
How to use NOEMA



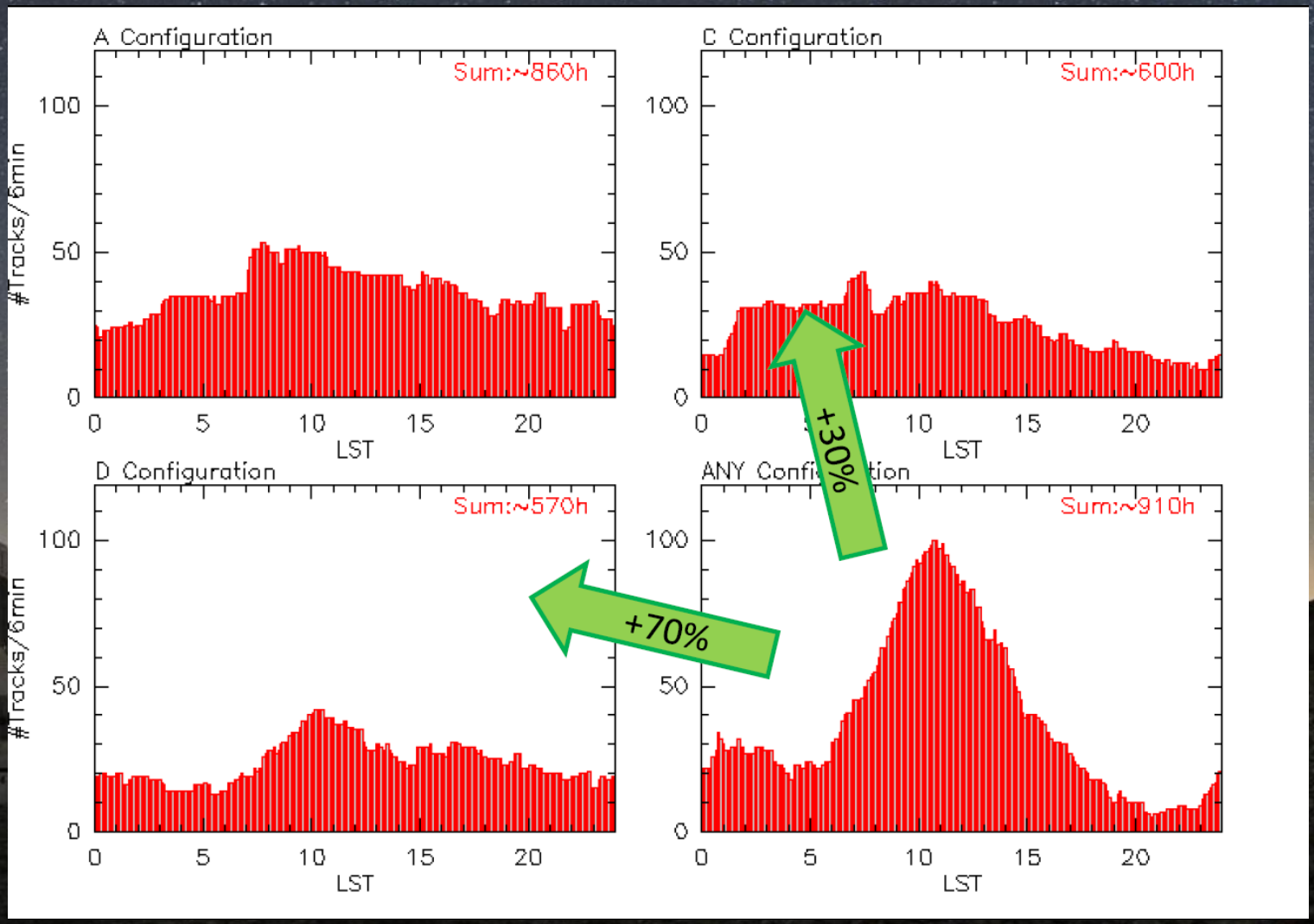
How to use NOEMA



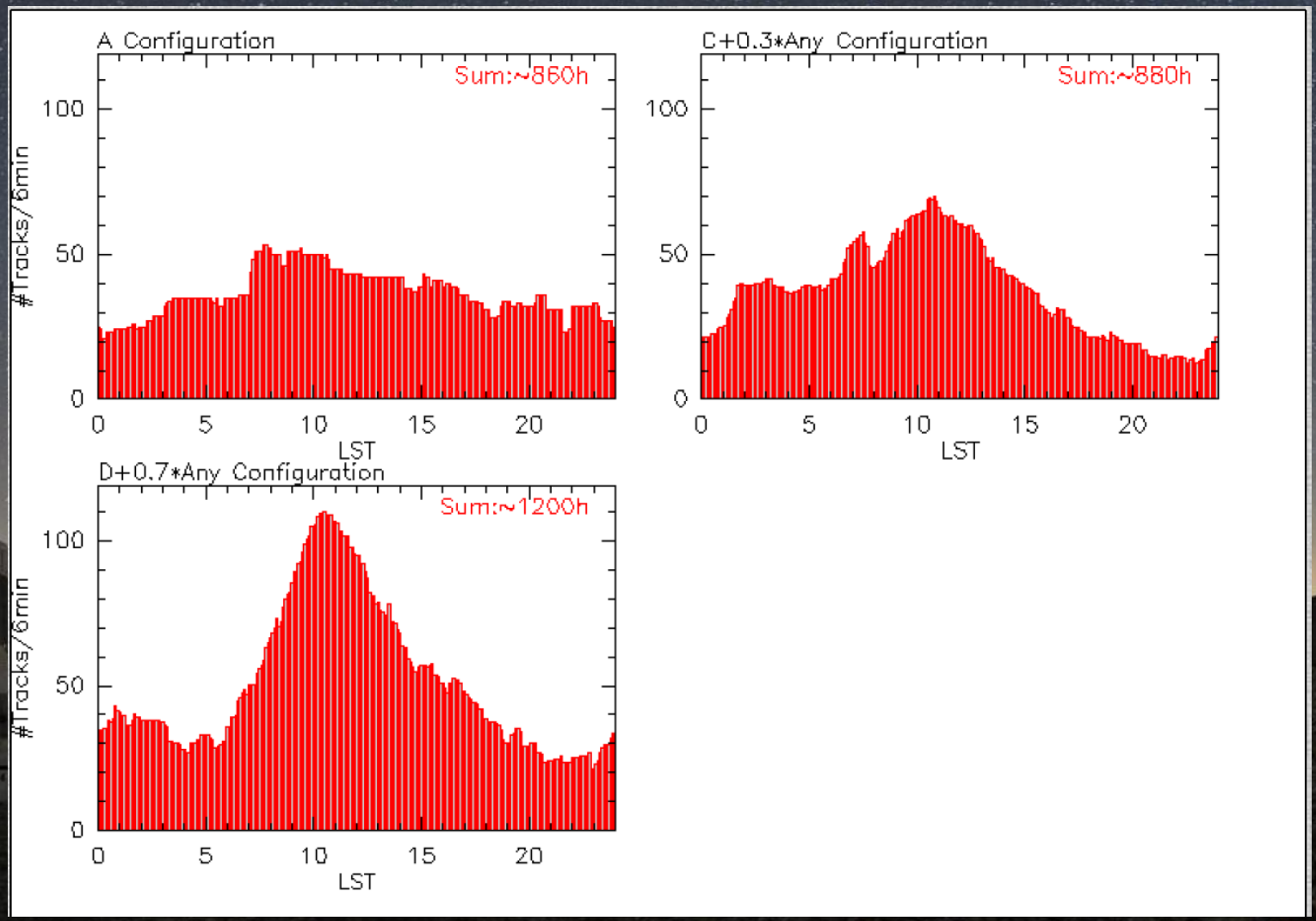
How to use NOEMA



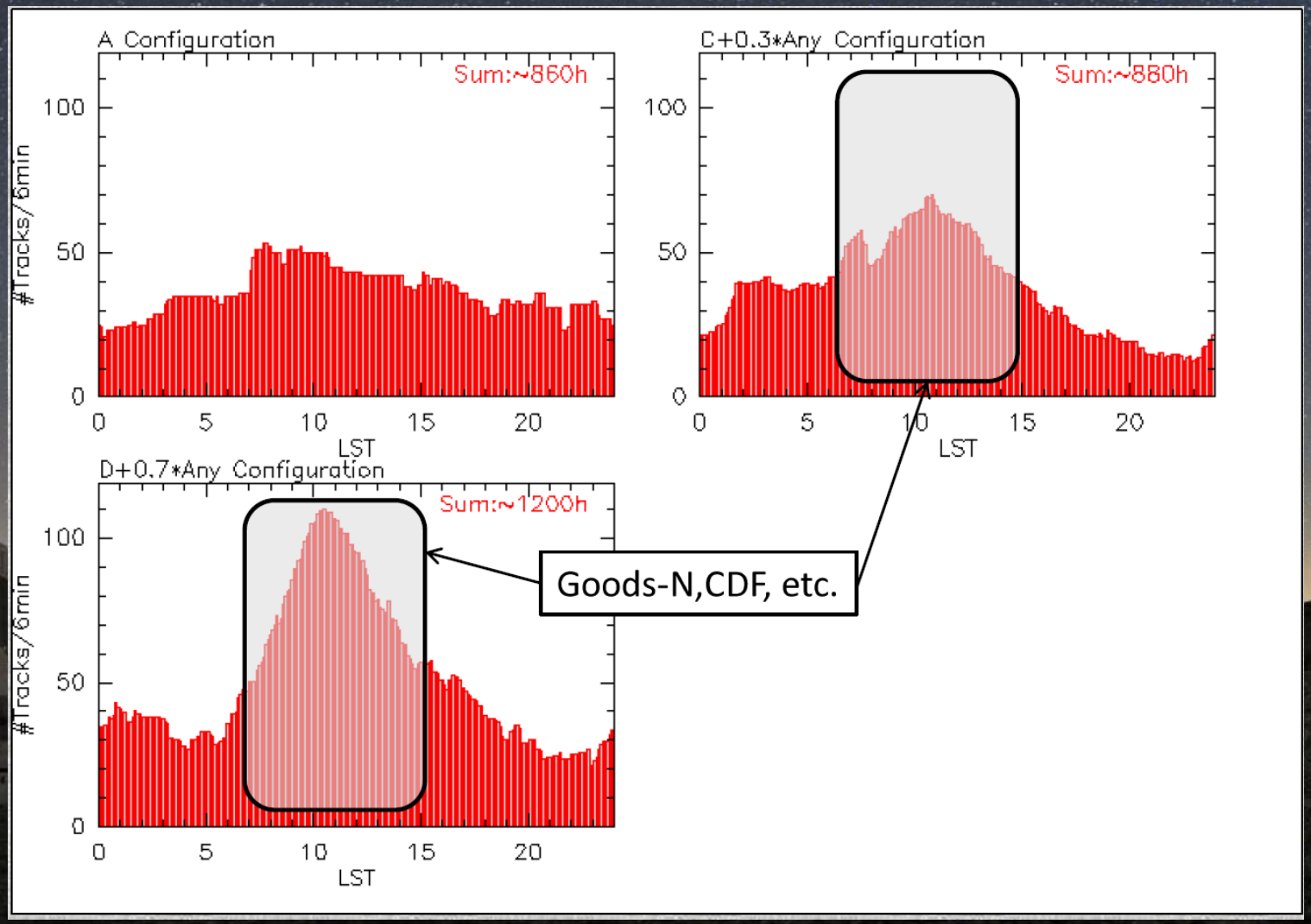
How to use NOEMA



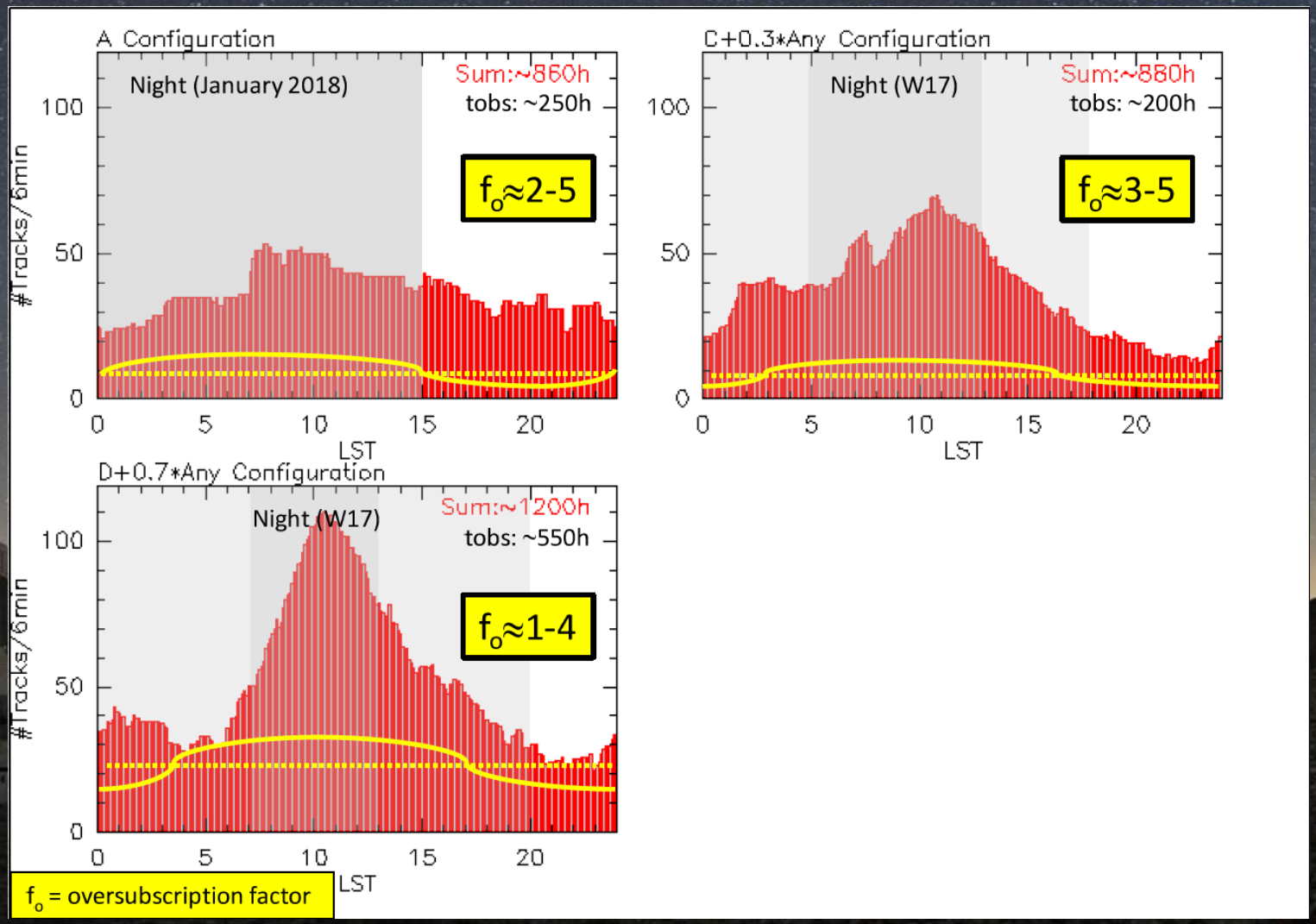
How to use NOEMA



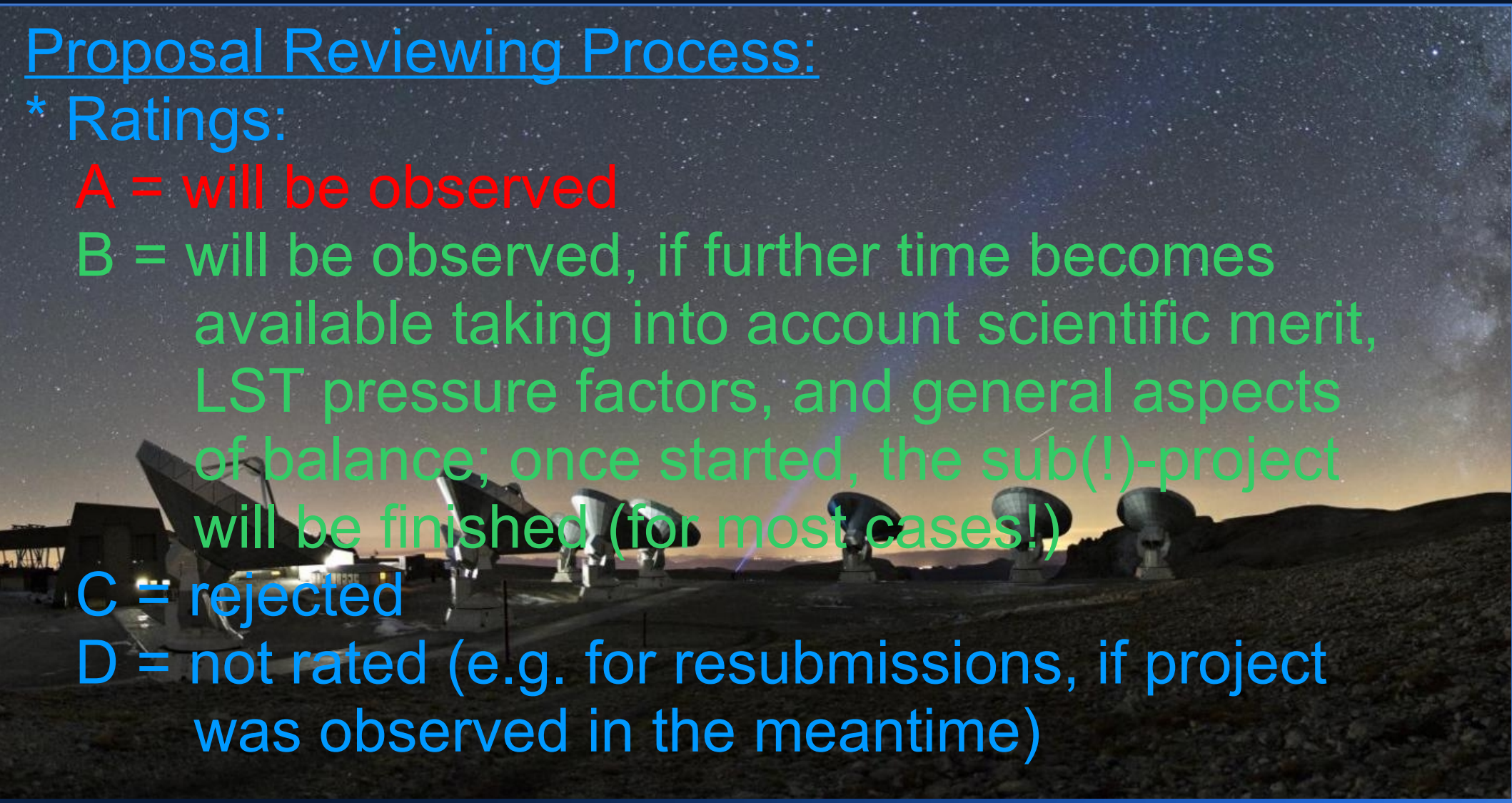
How to use NOEMA



How to use NOEMA



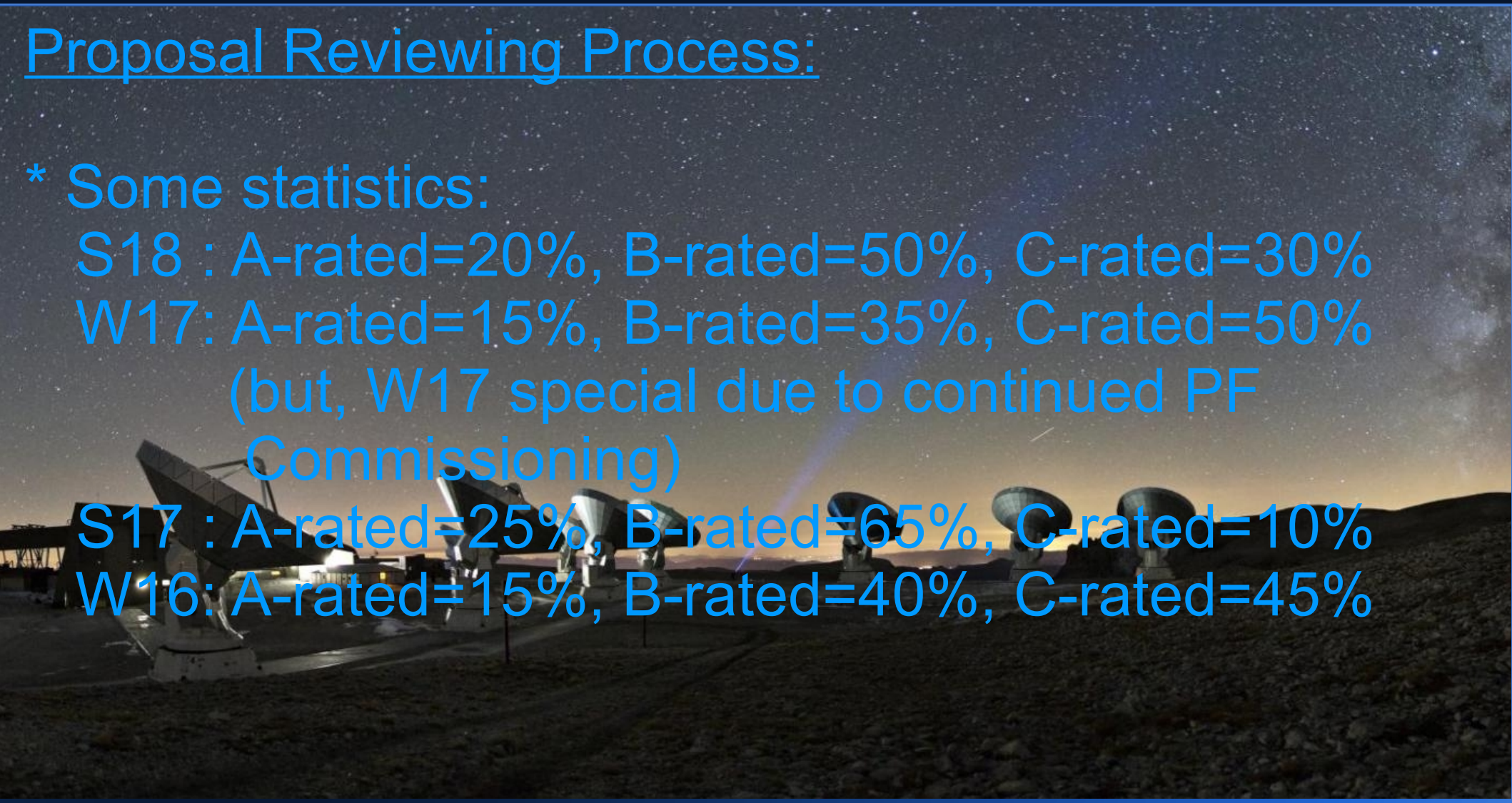
How to use NOEMA



Proposal Reviewing Process:

* Ratings:
  - A = will be observed
  - B = will be observed, if further time becomes available taking into account scientific merit, LST pressure factors, and general aspects of balance; once started, the sub(!)-project will be finished (for most cases!)
  - C = rejected
  - D = not rated (e.g. for resubmissions, if project was observed in the meantime)

How to use NOEMA



Proposal Reviewing Process:

* Some statistics:
  S18 : A-rated=20%, B-rated=50%, C-rated=30%
  W17: A-rated=15%, B-rated=35%, C-rated=50%
  (but, W17 special due to continued PF Commissioning)
  S17 : A-rated=25%, B-rated=65%, C-rated=10%
  W16: A-rated=15%, B-rated=40%, C-rated=45%

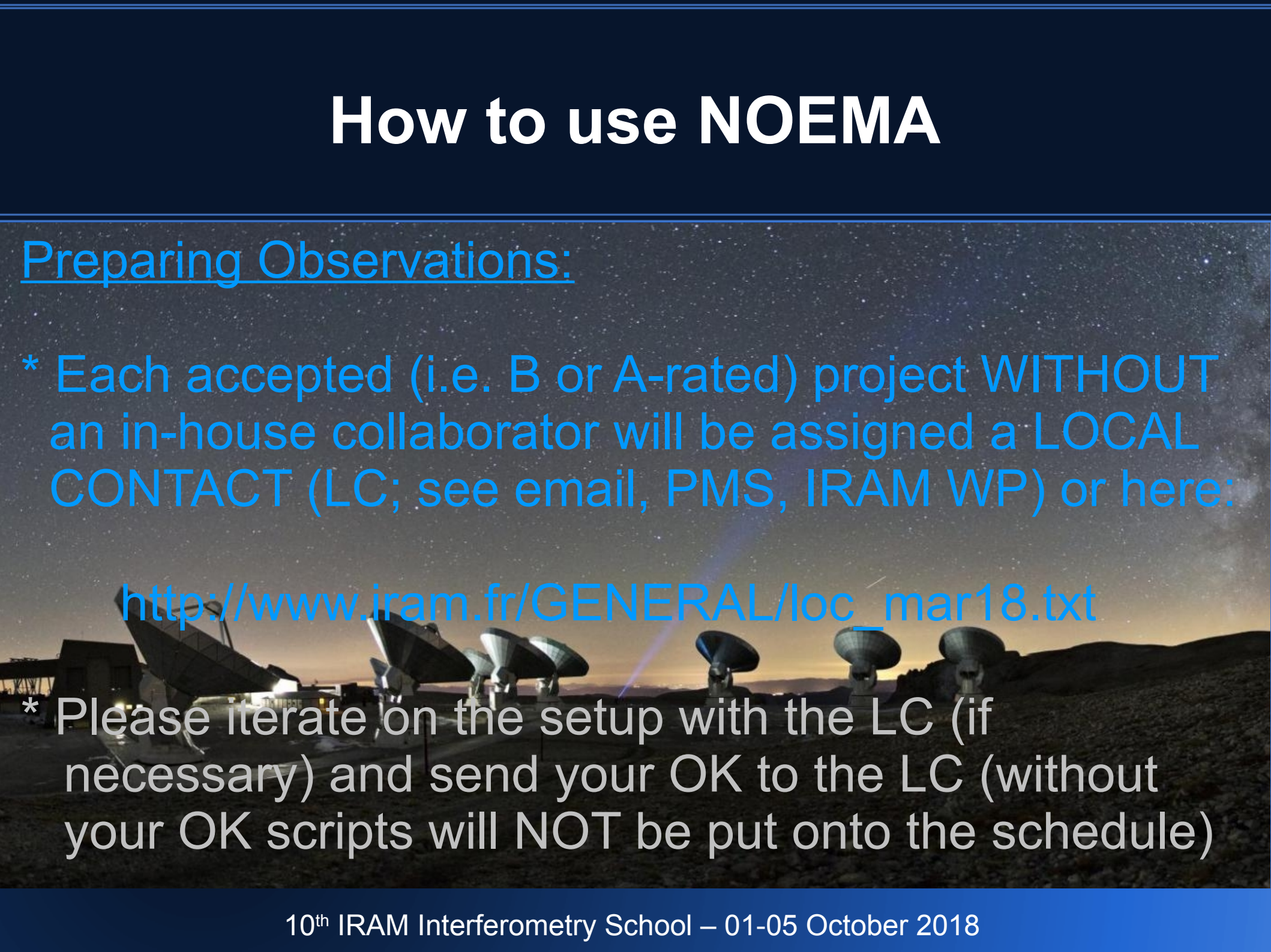
How to use NOEMA

Proposal Reviewing Process:

* Some statistics:
  Since S14, we observed ~65%±10% of all accepted (standard+LP) proposals (not counting DDTs (90-100%))
  -> this means for B-rated projects, ~50% are observed

How to use NOEMA

How to use NOEMA



Preparing Observations:

* Each accepted (i.e. B or A-rated) project WITHOUT an in-house collaborator will be assigned a LOCAL CONTACT (LC; see email, PMS, IRAM WP) or here:

http://www.iram.fr/GENERAL/loc_mar18.txt

* Please iterate on the setup with the LC (if necessary) and send your OK to the LC (without your OK scripts will NOT be put onto the schedule)

How to use NOEMA

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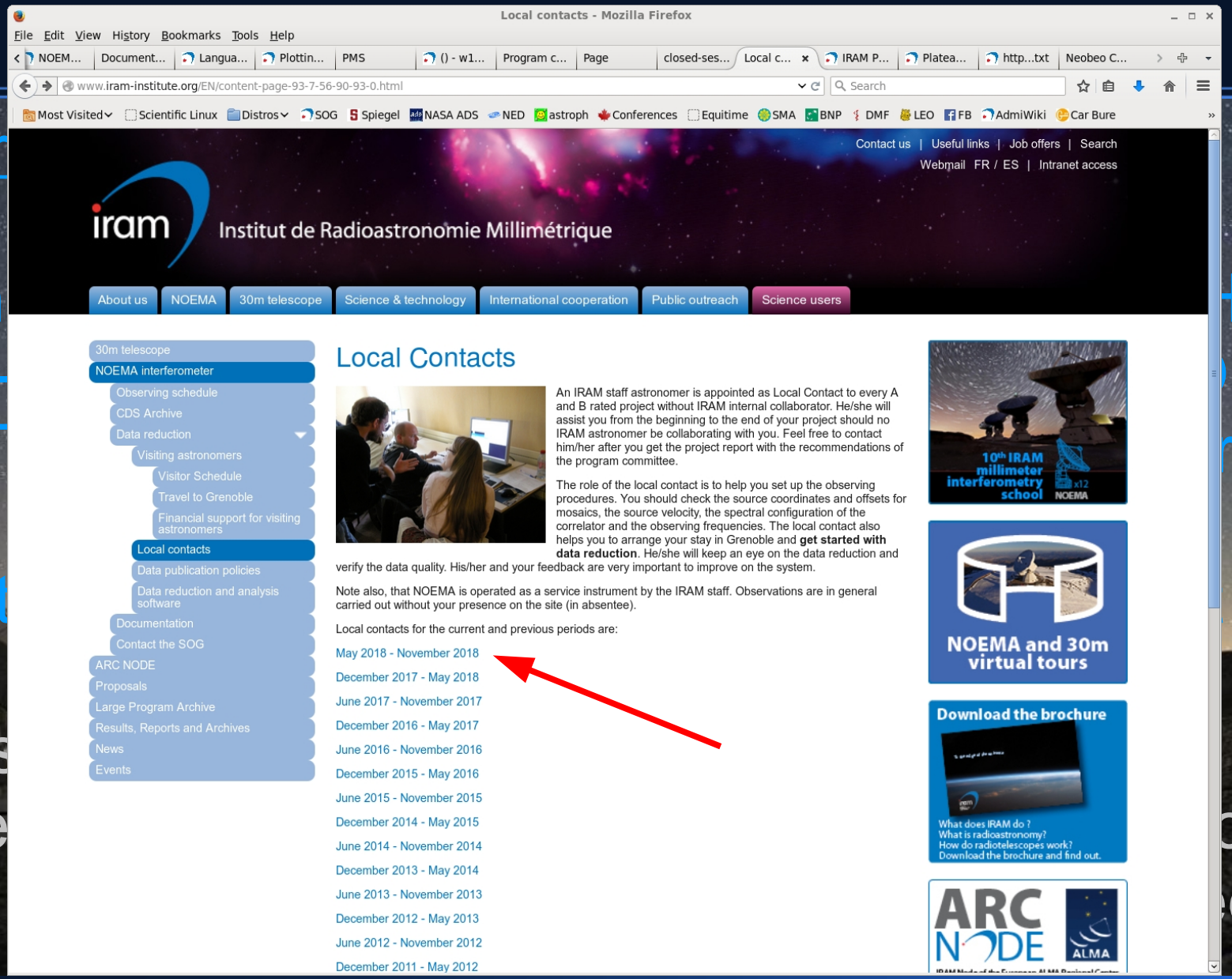

How to use NOEMA

Preparing Ob

* Each accep : WITHOUT
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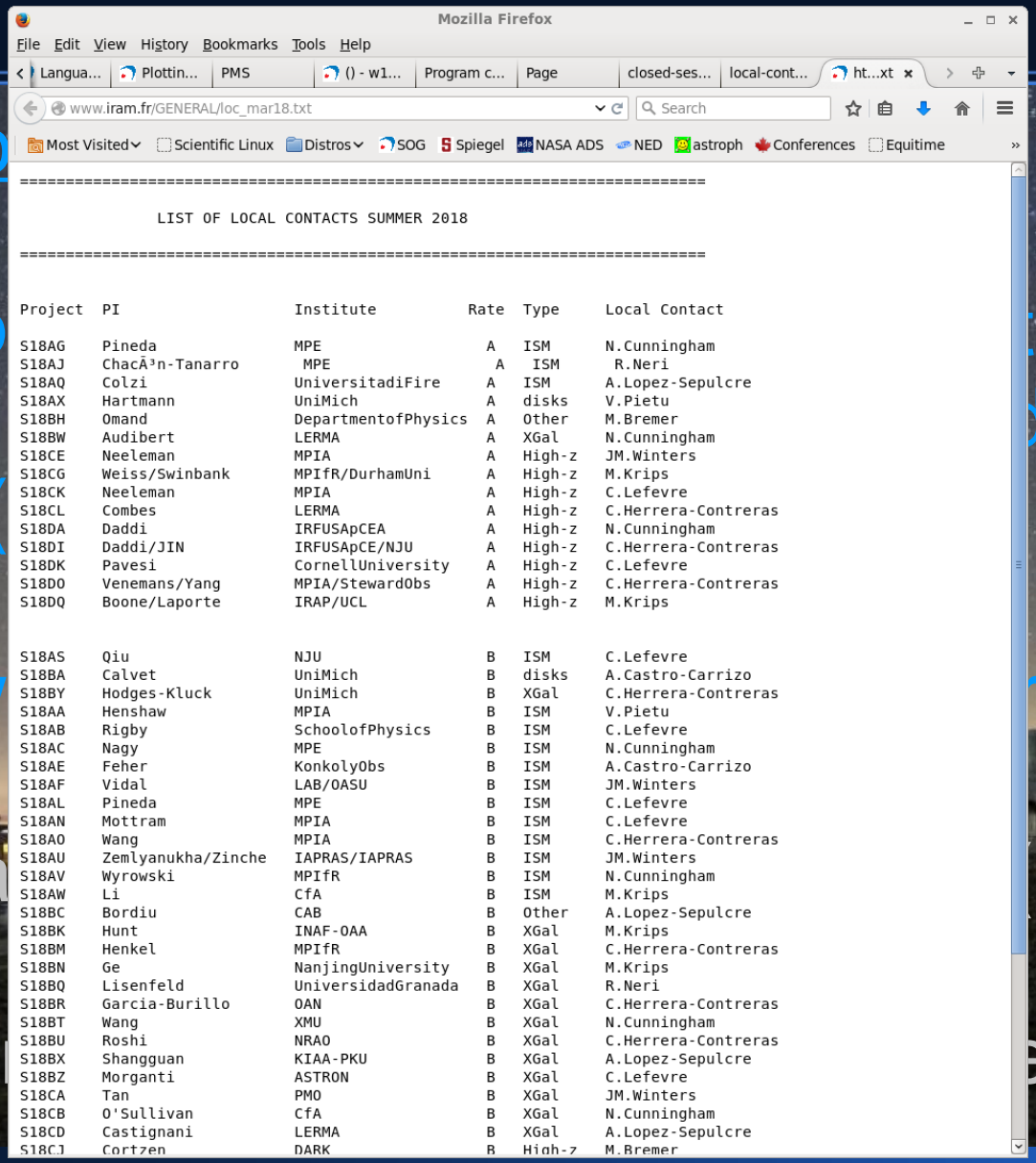
CONTACT (

http://ww r18.txt

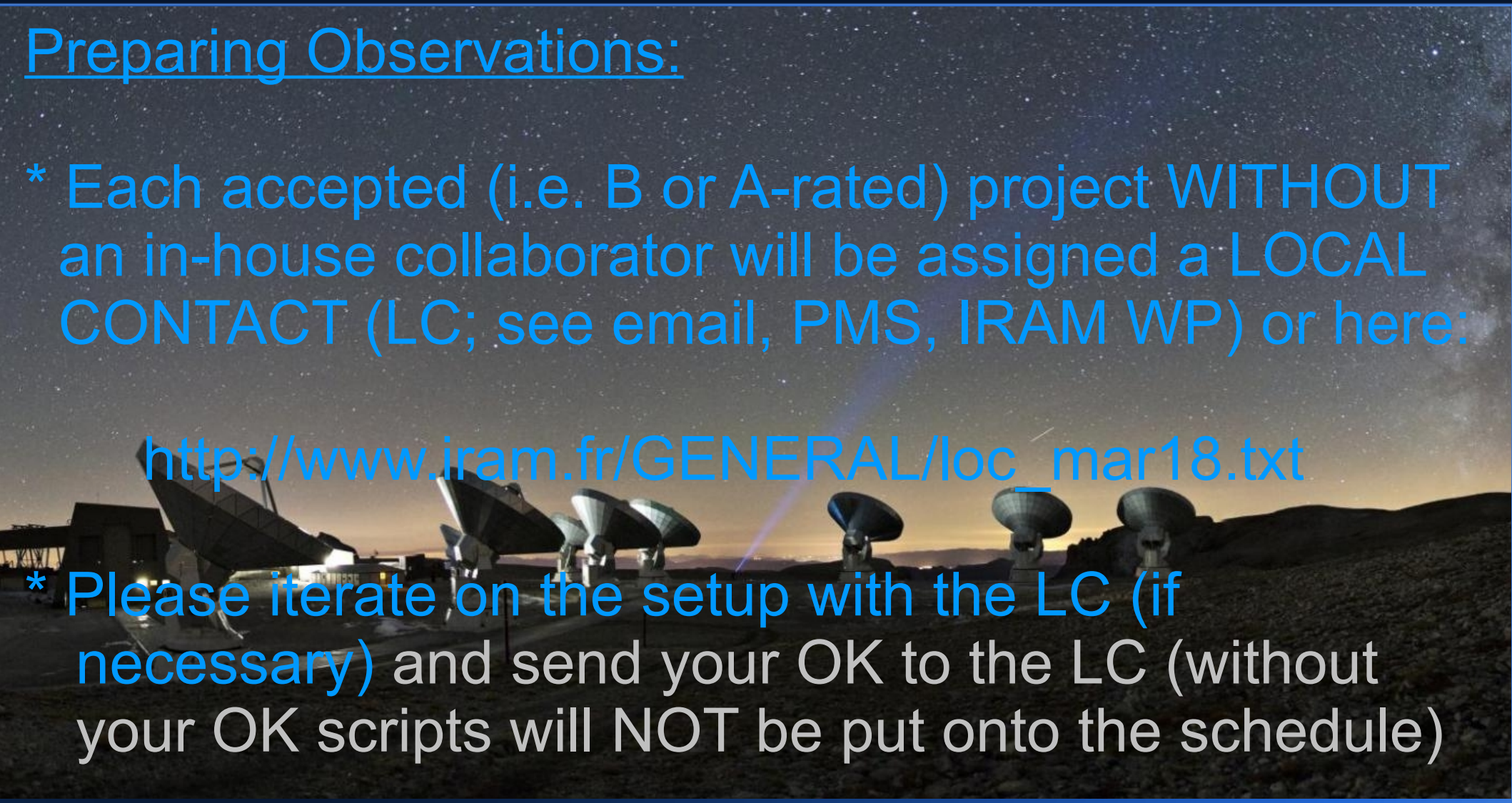
* Please itera if
necessary) (without
your OK sc e schedule)



LIST OF LOCAL CONTACTS SUMMER 2018

Project	PI	Institute	Rate	Type	Local Contact
S18AG	Pineda	MPE	A	ISM	N.Cunningham
S18AJ	ChacÃ³n-Tanarro	MPE	A	ISM	R.Neri
S18AQ	Colzi	UniversitadiFire	A	ISM	A.Lopez-Sepulcre
S18AX	Hartmann	UniMich	A	disks	V.Pietu
S18BH	Omand	DepartmentofPhysics	A	Other	M.Bremer
S18BW	Audibert	LERMA	A	XGal	N.Cunningham
S18CE	Neeleman	MPIA	A	High-z	JM.Winters
S18CG	Weiss/Swinbank	MPIfR/DurhamUni	A	High-z	M.Krips
S18CK	Neeleman	MPIA	A	High-z	C.Lefevre
S18CL	Combes	LERMA	A	High-z	C.Herrera-Contreras
S18DA	Daddi	IRFUSApCEA	A	High-z	N.Cunningham
S18DI	Daddi/JIN	IRFUSApCE/NJU	A	High-z	C.Herrera-Contreras
S18DK	Pavesi	CornellUniversity	A	High-z	C.Lefevre
S18DO	Venemans/Yang	MPIA/StewardObs	A	High-z	C.Herrera-Contreras
S18DQ	Boone/Laporte	IRAP/UCL	A	High-z	M.Krips
S18AS	Qiu	NJU	B	ISM	C.Lefevre
S18BA	Calvet	UniMich	B	disks	A.Castro-Carrizo
S18BY	Hodges-Kluck	UniMich	B	XGal	C.Herrera-Contreras
S18AA	Henshaw	MPIA	B	ISM	V.Pietu
S18AB	Rigby	SchoolofPhysics	B	ISM	C.Lefevre
S18AC	Nagy	MPE	B	ISM	N.Cunningham
S18AE	Feher	KonkolyObs	B	ISM	A.Castro-Carrizo
S18AF	Vidal	LAB/OASU	B	ISM	JM.Winters
S18AL	Pineda	MPE	B	ISM	C.Lefevre
S18AN	Mottram	MPIA	B	ISM	C.Lefevre
S18AO	Wang	MPIA	B	ISM	C.Herrera-Contreras
S18AU	Zemlyanukha/Zinche	IAPRAS/IAPRAS	B	ISM	JM.Winters
S18AV	Wyrowski	MPIfR	B	ISM	N.Cunningham
S18AW	Li	CfA	B	ISM	M.Krips
S18BC	Bordiu	CAB	B	Other	A.Lopez-Sepulcre
S18BK	Hunt	INAF-OAA	B	XGal	M.Krips
S18BM	Henkel	MPIfR	B	XGal	C.Herrera-Contreras
S18BN	Ge	NanjingUniversity	B	XGal	M.Krips
S18BQ	Lisenfeld	UniversidadGranada	B	XGal	R.Neri
S18BR	Garcia-Burillo	OAN	B	XGal	C.Herrera-Contreras
S18BT	Wang	XMU	B	XGal	N.Cunningham
S18BU	Roshi	NRAO	B	XGal	C.Herrera-Contreras
S18BX	Shangguan	KIAA-PKU	B	XGal	A.Lopez-Sepulcre
S18BZ	Morganti	ASTRON	B	XGal	C.Lefevre
S18CA	Tan	PMO	B	XGal	JM.Winters
S18CB	O'Sullivan	CfA	B	XGal	N.Cunningham
S18CD	Castignani	LERMA	B	XGal	A.Lopez-Sepulcre
S18CJ	Cortzen	DARK	B	High-z	M.Bremer

How to use NOEMA



Preparing Observations:

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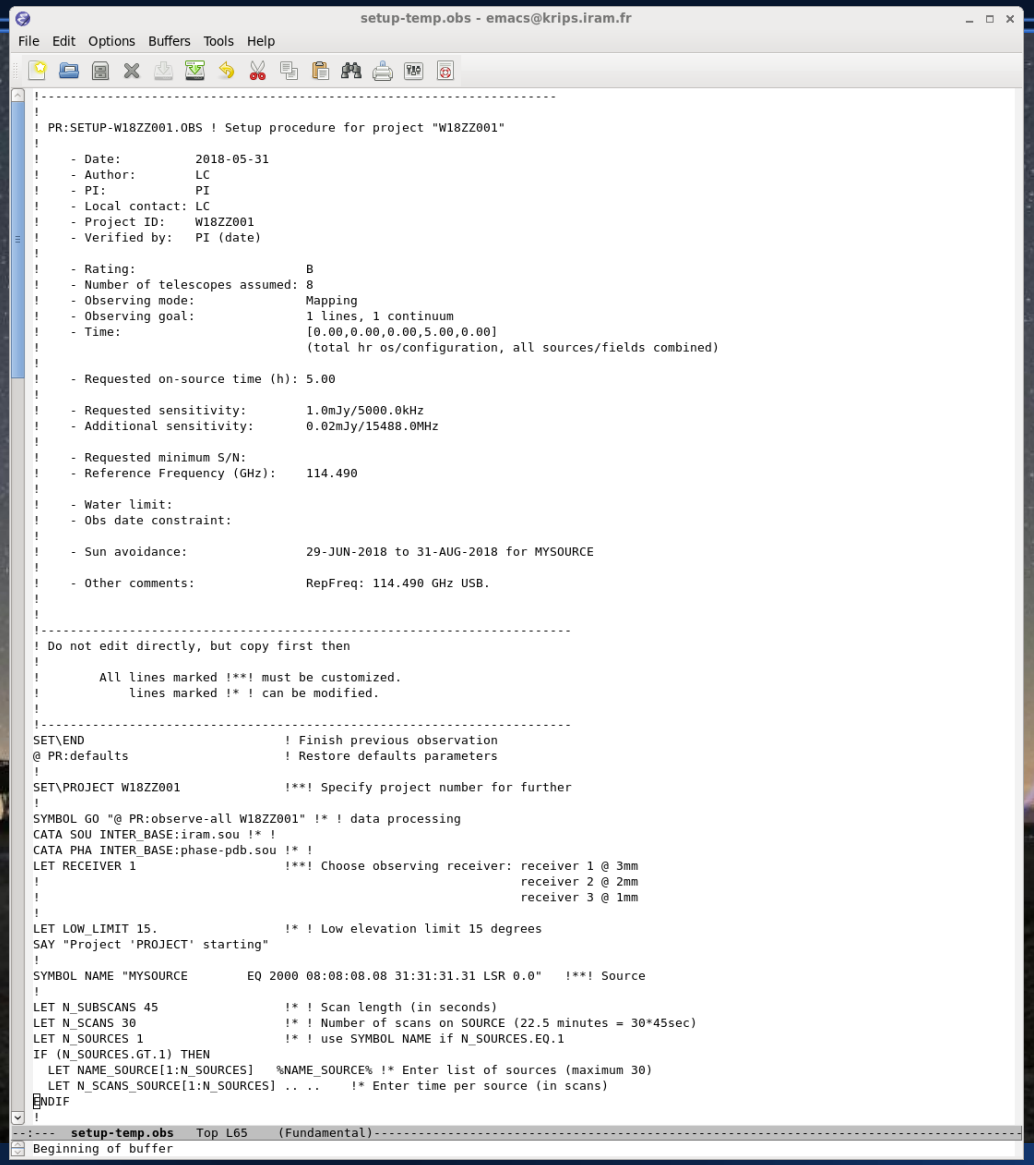
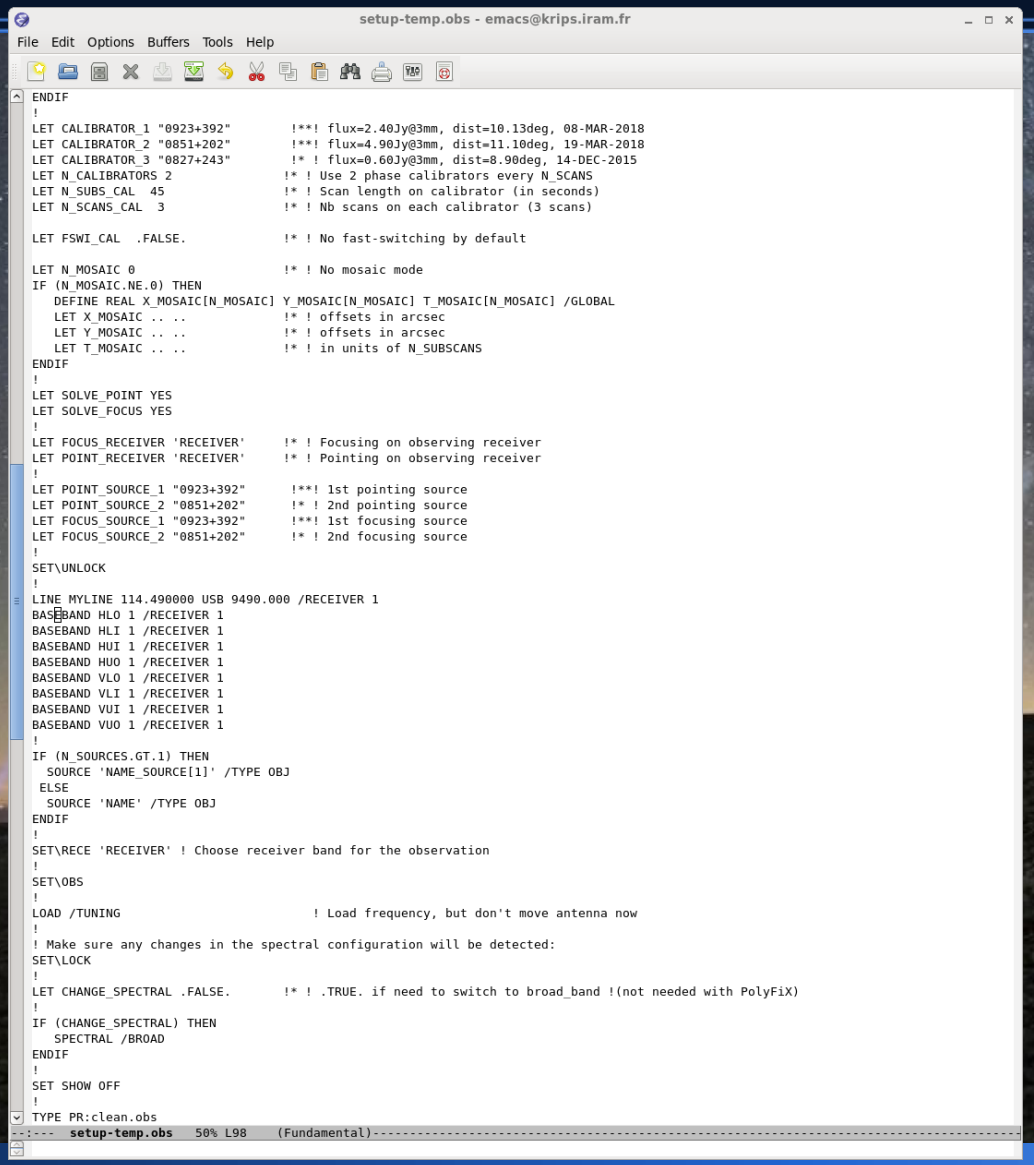
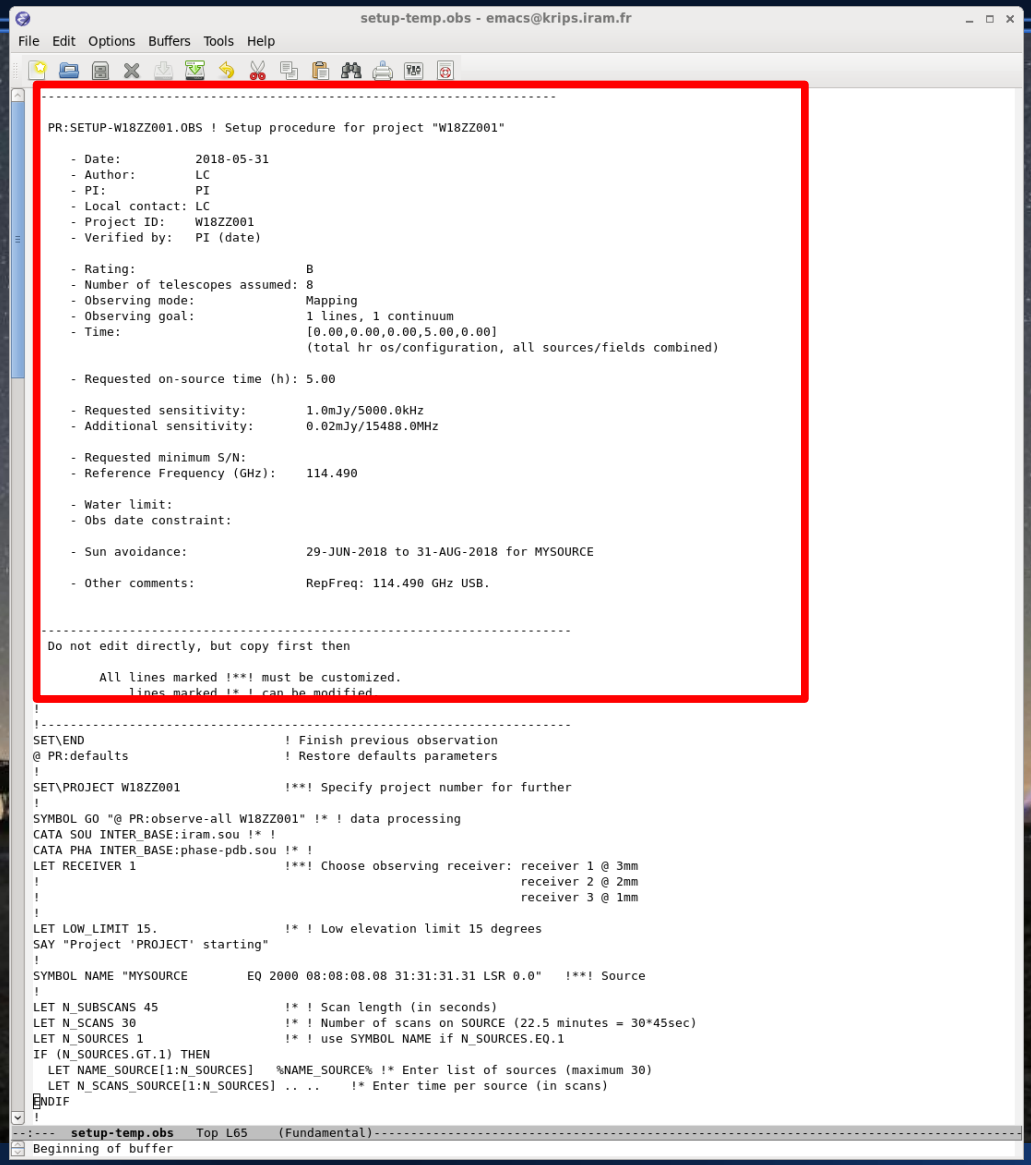
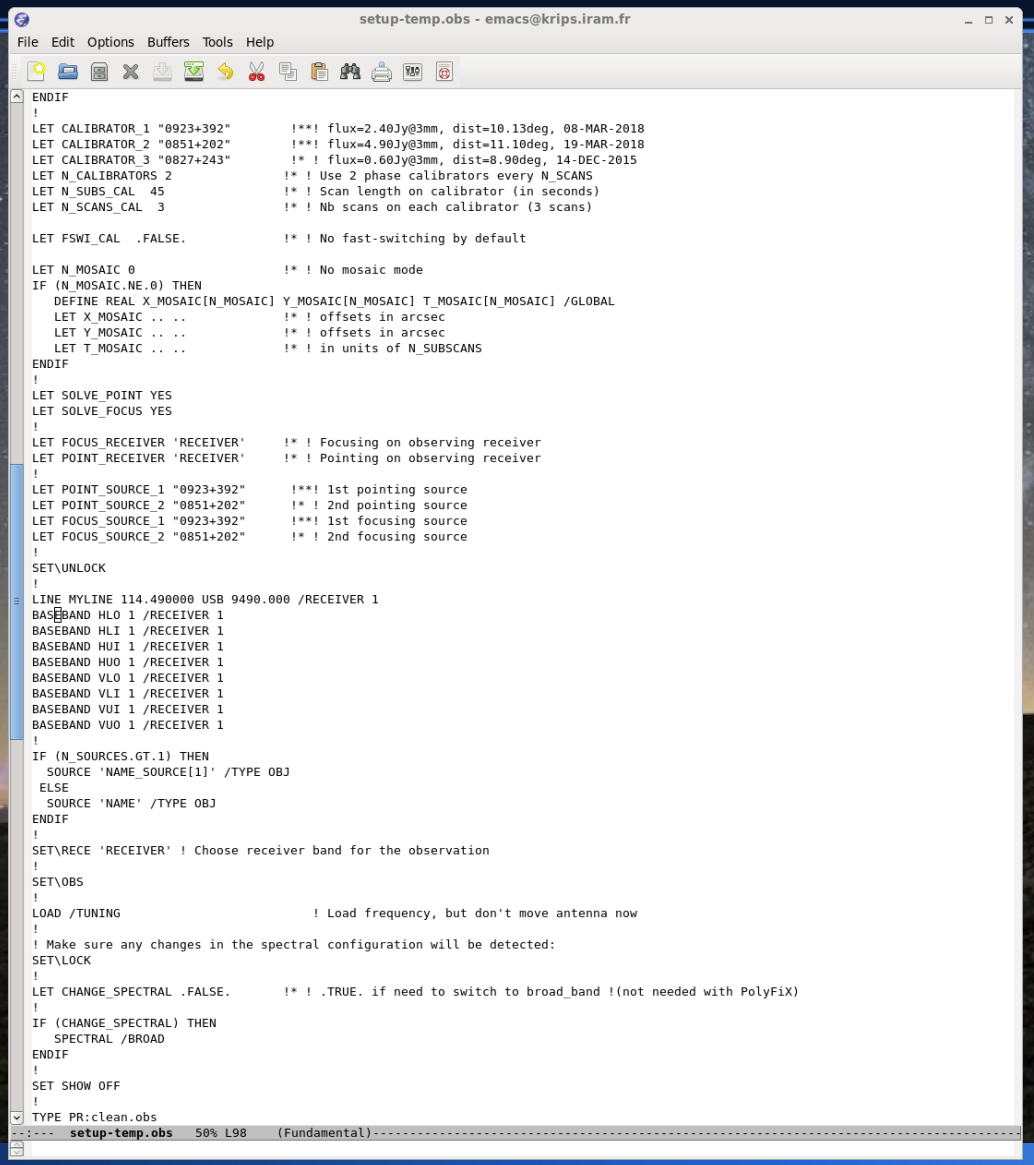
How to use NOEMA



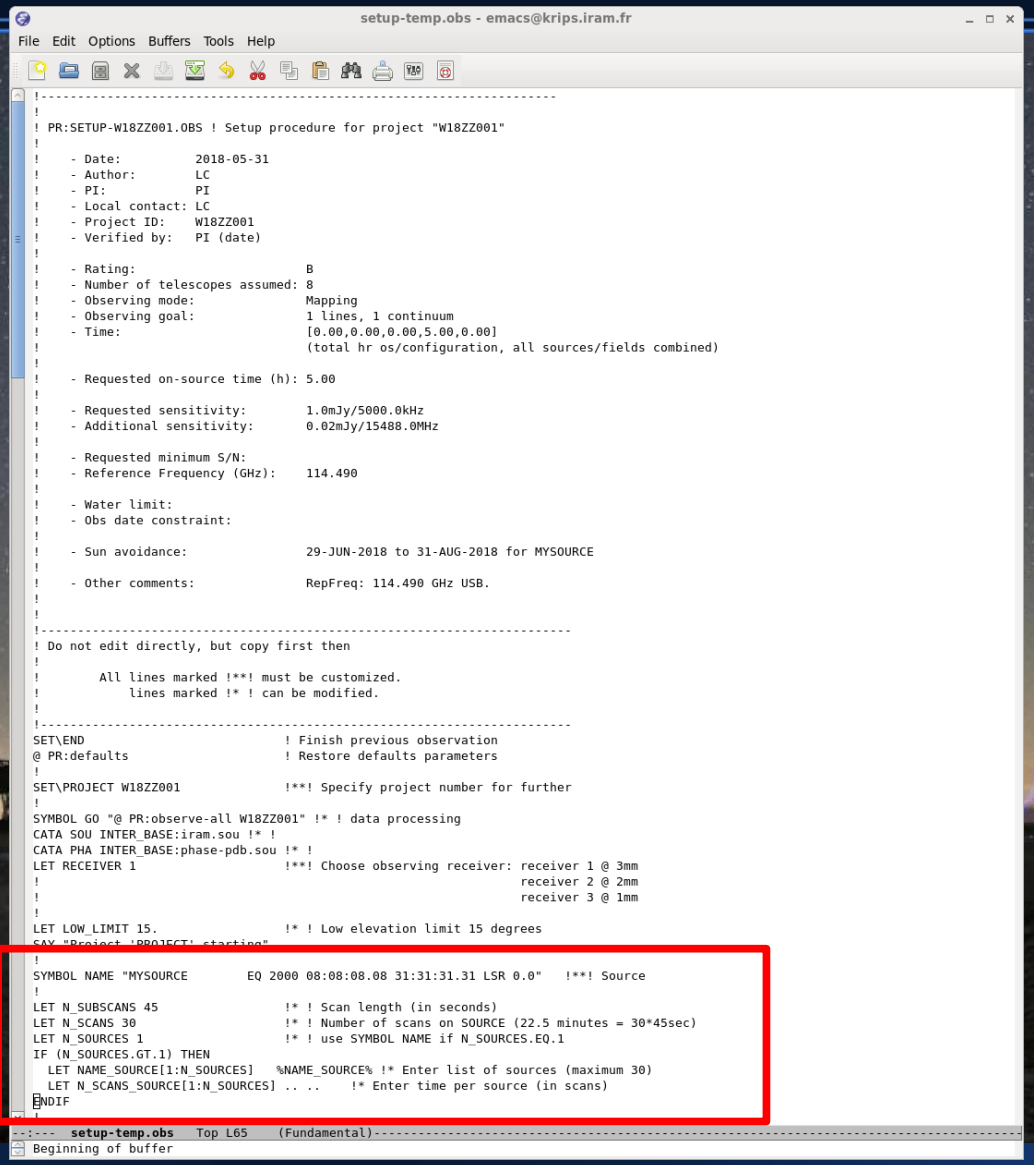


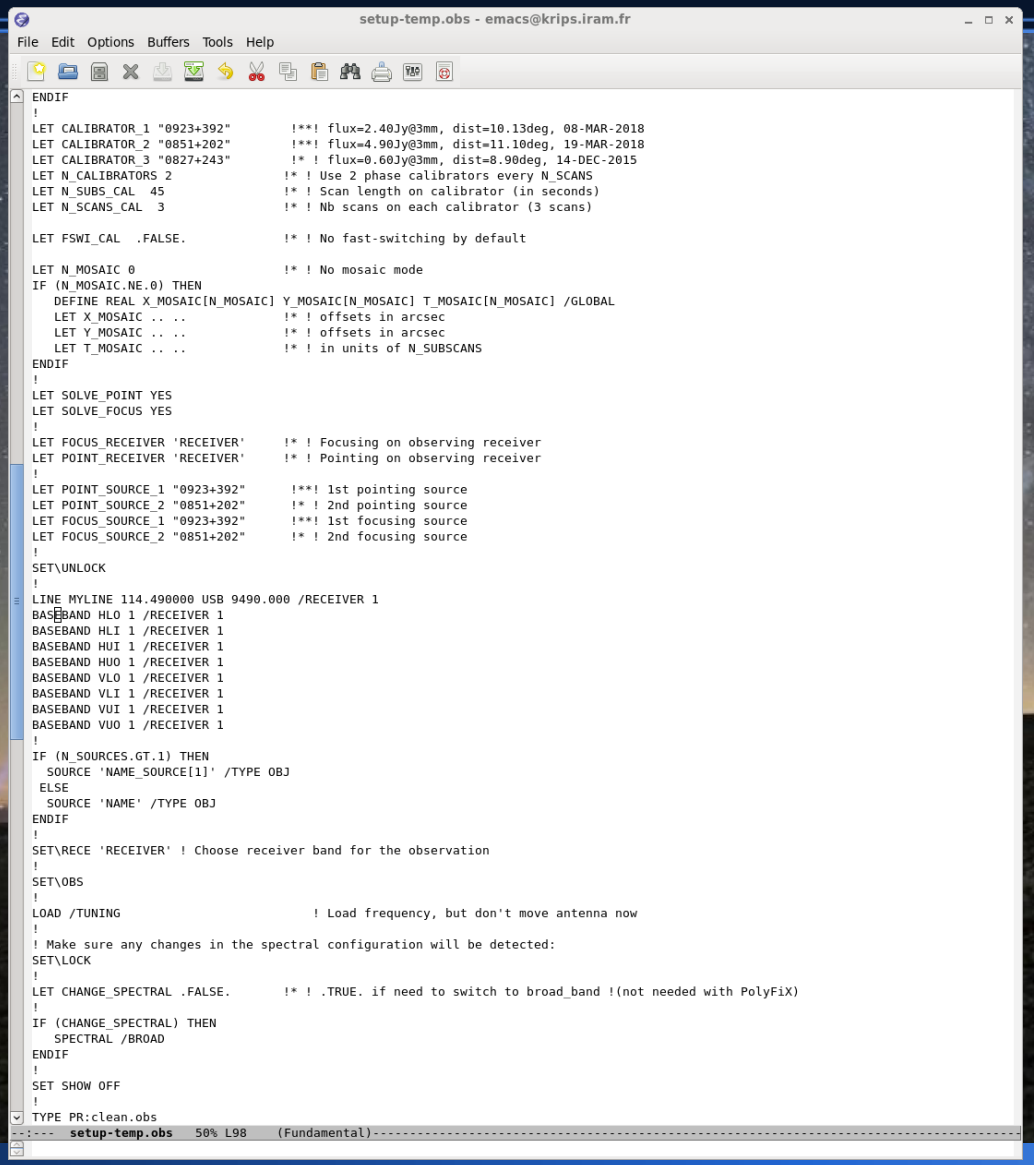
How to use NOEMA



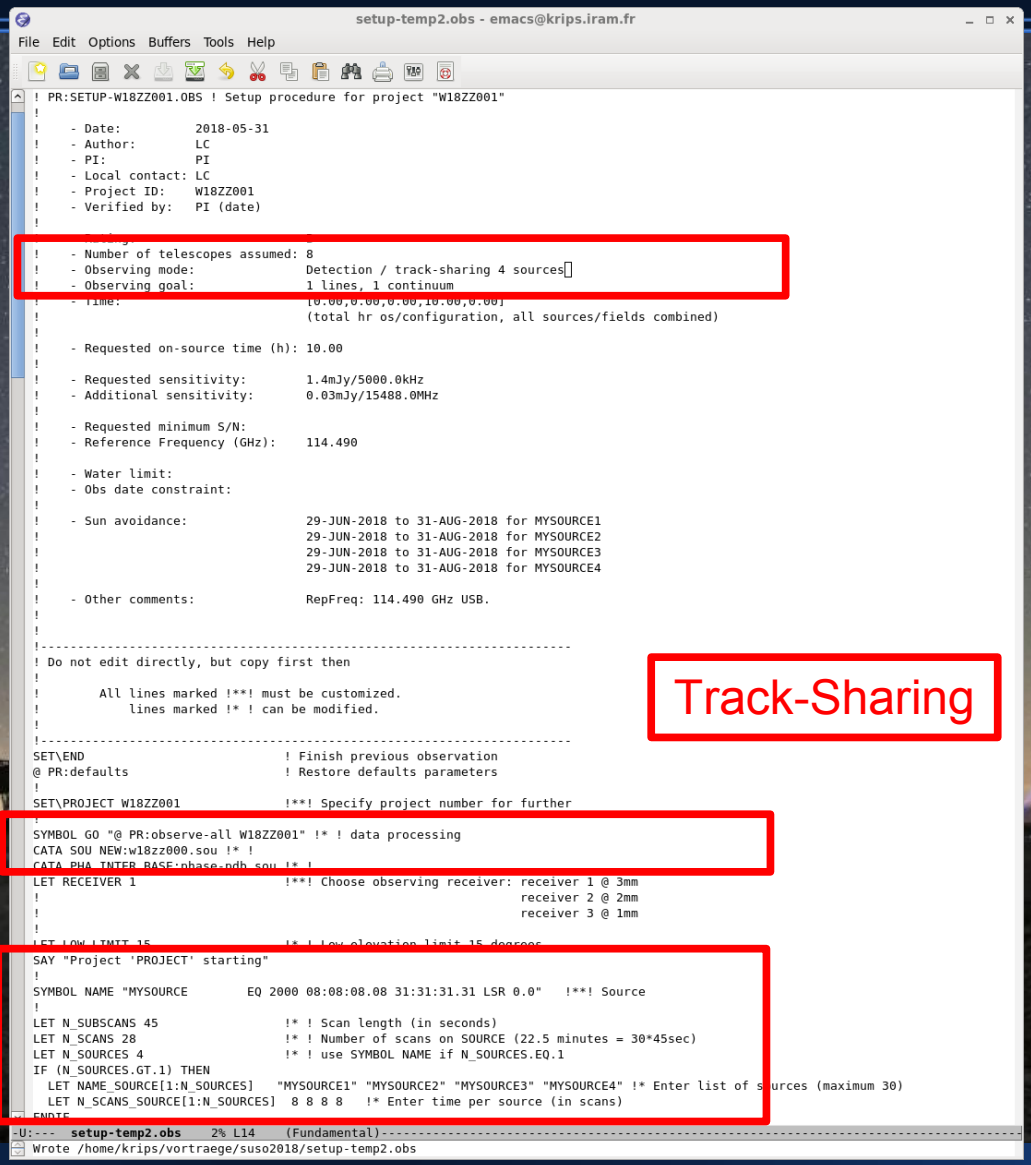


How to use NOEMA

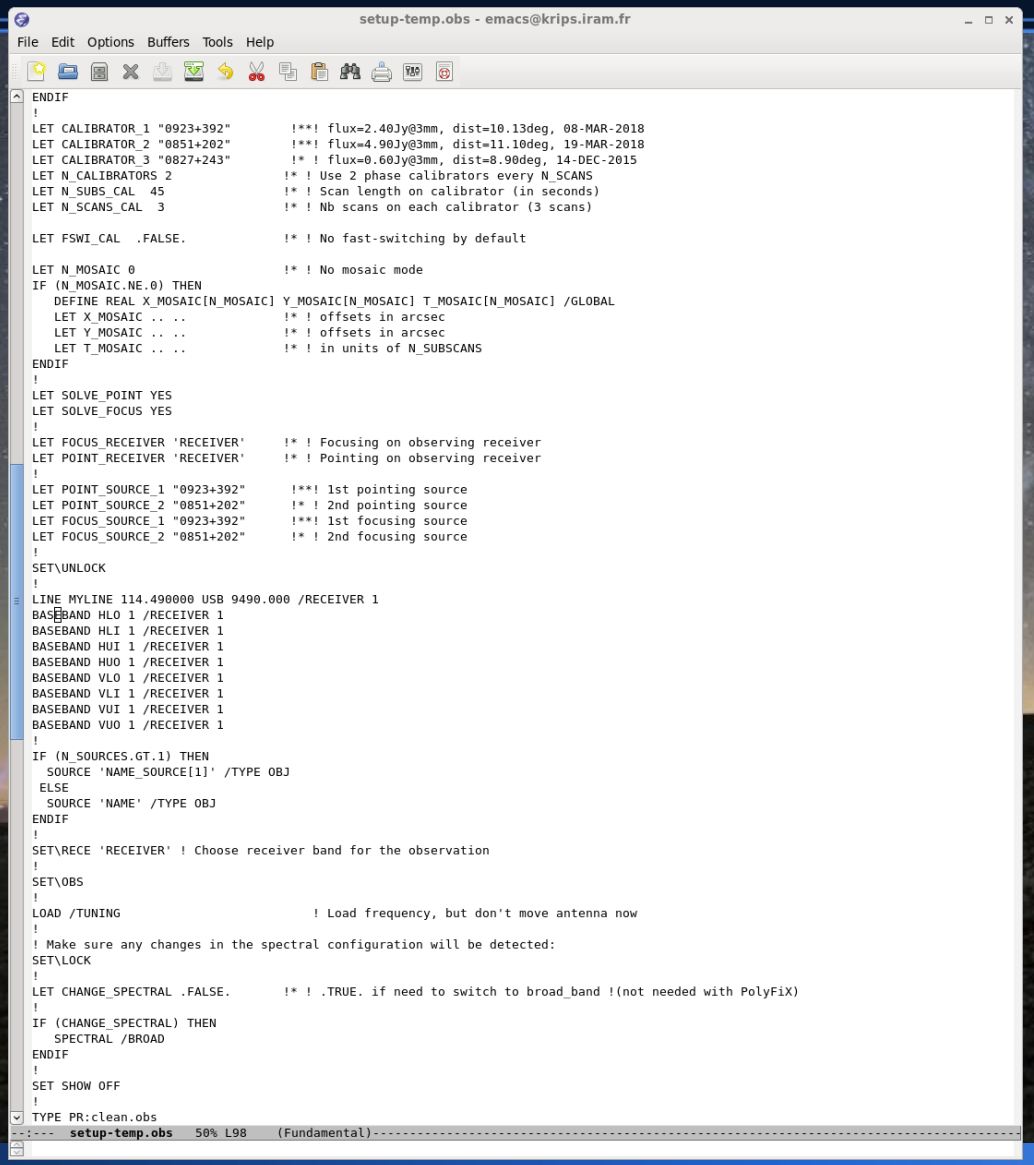




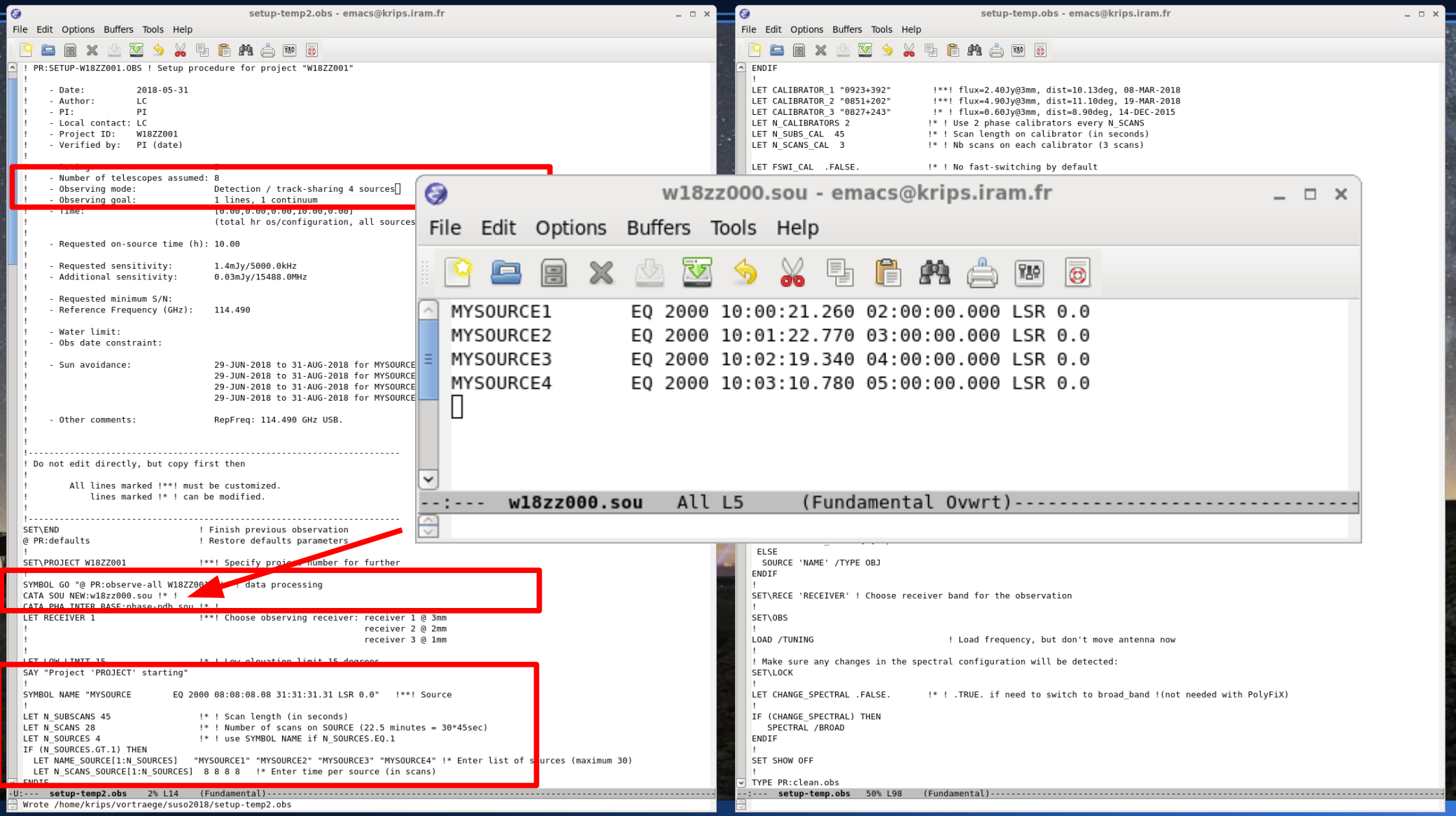
How to use NOEMA



Track-Sharing



How to use NOEMA



How to use NOEMA

setup-temp.obs - emacs@krips.iram.fr

File Edit Options Buffers Tools Help

```
!------------------------------------------------------------------------
!
! PR:SETUP-W18ZZ001.OBS ! Setup procedure for project "W18ZZ001"
!
!    - Date:          2018-05-31
!    - Author:        LC
!    - PI:            PI
!    - Local contact: LC
!    - Project ID:    W18ZZ001
!    - Verified by:   PI (date)
!
!    - Rating:                    B
!    - Observing mode:            Mapping / mosaic 2 fields
!    - Observing goal:            1 lines, 1 continuum
!    - Time:                      [0.00,0.00,0.00,5.00,0.00]
!                                 (total hr os/configuration, all sources/fields combined)
!
!    - Requested on-source time (h): 5.00
!
!    - Requested sensitivity:     1.0mJy/5000.0kHz
!    - Additional sensitivity:    0.02mJy/15488.0MHz
!
!    - Requested minimum S/N:
!    - Reference Frequency (GHz):    114.490
!
!    - Water limit:
!    - Obs date constraint:
!
!    - Sun avoidance:             29-JUN-2018 to 31-AUG-2018 for MYSOURCE
!
!    - Other comments:            RepFreq: 114.490 GHz USB.
!
!
!------------------------------------------------------------------------
! Do not edit directly, but copy first then
!
!       All lines marked !**! must be customized.
!           lines marked !* ! can be modified.
!
!------------------------------------------------------------------------
SET\END                          ! Finish previous observation
@ PR:defaults                    ! Restore defaults parameters
!
SET\PROJECT W18ZZ001             !**! Specify project number for further
!
SYMBOL GO "@ PR:observe-all W18ZZ001" !* ! data processing
CATA SOU INTER_BASE:iram.sou !* !
CATA PHA INTER_BASE:phase-pdb.sou !* !
LET RECEIVER 1                   !**! Choose observing receiver: receiver 1 @ 3mm
!                                                                receiver 2 @ 2mm
!                                                                receiver 3 @ 1mm
!
LET LOW_LIMIT 15.                !* ! Low elevation limit 15 degrees
SAY "Project 'PROJECT' starting"
!
SYMBOL NAME "MYSOURCE        EQ 2000 08:08:08.08 31:31:31.31 LSR 0.0"   !**! Source
!
LET N_SUBSCANS 45                !* ! Scan length (in seconds)
LET N_SCANS 30                   !* ! Number of scans on SOURCE (22.5 minutes = 30*45sec)
LET N_SOURCES 1                  !* ! use SYMBOL NAME if N_SOURCES.EQ.1
IF (N_SOURCES.GT.1) THEN
  LET NAME_SOURCE[1:N_SOURCES]   %NAME_SOURCE% !* Enter list of sources (maximum 30)
  LET N_SCANS_SOURCE[1:N_SOURCES] .. ..    !* Enter time per source (in scans)
ENDIF
!
```

-:--- **setup-temp.obs** Top L65 (Fundamental)

Beginning of buffer

MOSAIC!

setup-temp.obs - emacs@krips.iram.fr

File Edit Options Buffers Tools Help

```
ENDIF
!
LET CALIBRATOR_1 "0923+392"        !**! flux=2.40Jy@3mm, dist=10.13deg, 08-MAR-2018
LET CALIBRATOR_2 "0851+202"        !**! flux=4.90Jy@3mm, dist=11.10deg, 19-MAR-2018
LET CALIBRATOR_3 "0827+243"        !* ! flux=0.60Jy@3mm, dist=8.90deg, 14-DEC-2015
LET N_CALIBRATORS 2                !* ! Use 2 phase calibrators every N_SCANS
LET N_SUBS_CAL  45                 !* ! Scan length on calibrator (in seconds)
LET N_SCANS_CAL  3                 !* ! Nb scans on each calibrator (3 scans)

LET N_MOSAIC 2                     !* ! No mosaic mode
IF (N_MOSAIC.NE.0) THEN
   DEFINE REAL X_MOSAIC[N_MOSAIC] Y_MOSAIC[N_MOSAIC] T_MOSAIC[N_MOSAIC] /GLOBAL
   LET X_MOSAIC     9.00   -14.00              !* ! offsets in arcsec
   LET Y_MOSAIC     1.00    -4.00              !* ! offsets in arcsec
   LET T_MOSAIC       15       15              !* ! in units of N_SUBSCANS
ENDIF
!
LET SOLVE_POINT YES
LET SOLVE_FOCUS YES
!
LET FOCUS_RECEIVER 'RECEIVER'      !* ! Focusing on observing receiver
LET POINT_RECEIVER 'RECEIVER'      !* ! Pointing on observing receiver
!
LET POINT_SOURCE_1 "0923+392"       !**! 1st pointing source
LET POINT_SOURCE_2 "0851+202"       !* ! 2nd pointing source
LET FOCUS_SOURCE_1 "0923+392"       !**! 1st focusing source
LET FOCUS_SOURCE_2 "0851+202"       !* ! 2nd focusing source
!
SET\UNLOCK
!
LINE MYLINE 114.490000 USB 9490.000 /RECEIVER 1
BASEBAND HLO 1 /RECEIVER 1
BASEBAND HLI 1 /RECEIVER 1
BASEBAND HUI 1 /RECEIVER 1
BASEBAND HUO 1 /RECEIVER 1
BASEBAND VLO 1 /RECEIVER 1
BASEBAND VLI 1 /RECEIVER 1
BASEBAND VUI 1 /RECEIVER 1
BASEBAND VUO 1 /RECEIVER 1
!
IF (N_SOURCES.GT.1) THEN
  SOURCE 'NAME_SOURCE[1]' /TYPE OBJ
 ELSE
  SOURCE 'NAME' /TYPE OBJ
ENDIF
!
SET\RECE 'RECEIVER' ! Choose receiver band for the observation
!
SET\OBS
!
LOAD /TUNING                          ! Load frequency, but don't move antenna now
!
! Make sure any changes in the spectral configuration will be detected:
SET\LOCK
!
LET CHANGE_SPECTRAL .FALSE.       !* ! .TRUE. if need to switch to broad_band !(not needed with PolyFiX)
!
IF (CHANGE_SPECTRAL) THEN
   SPECTRAL /BROAD
ENDIF
!
SET SHOW OFF
!
TYPE PR:clean.obs
```

-:--- **setup-temp.obs** 50% L98 (Fundamental)

How to use NOEMA

setup-temp.obs - emacs@krips.iram.fr

File Edit Options Buffers Tools Help

```
!------------------------------------------------------------------------
!
! PR:SETUP-W18ZZ001.OBS ! Setup procedure for project "W18ZZ001"
!
!    - Date:          2018-05-31
!    - Author:        LC
!    - PI:            PI
!    - Local contact: LC
!    - Project ID:    W18ZZ001
!    - Verified by:   PI (date)
!
!    - Rating:                       B
!    - Number of telescopes assumed: 8
!    - Observing mode:               Mapping
!    - Observing goal:               1 lines, 1 continuum
!    - Time:                         [0.00,0.00,0.00,5.00,0.00]
!                                    (total hr os/configuration, all sources/fields combined)
!
!    - Requested on-source time (h): 5.00
!
!    - Requested sensitivity:        1.0mJy/5000.0kHz
!    - Additional sensitivity:       0.02mJy/15488.0MHz
!
!    - Requested minimum S/N:
!    - Reference Frequency (GHz):    114.490
!
!    - Water limit:
!    - Obs date constraint:
!
!    - Sun avoidance:
!
!    - Other comments:
!
!
!---------------------------------
! Do not edit directly, but copy first then
!
!       All lines marked !**! must be customized.
!           lines marked !* ! can be modified.
!
!------------------------------------------------------------------------
SET\END                         ! Finish previous observation
@ PR:defaults                   ! Restore defaults parameters
!
SET\PROJECT W18ZZ001            !**! Specify project number for further
!
SYMBOL GO "@ PR:observe-all W18ZZ001" !* ! data processing
CATA SOU INTER_BASE:iram.sou !* !
CATA PHA INTER_BASE:phase-pdb.sou !* !
LET RECEIVER 1                  !**! Choose observing receiver: receiver 1 @ 3mm
!                                                               receiver 2 @ 2mm
!                                                               receiver 3 @ 1mm
!
LET LOW_LIMIT 15.               !* ! Low elevation limit 15 degrees
SAY "Project 'PROJECT' starting"
!
SYMBOL NAME "MYSOURCE        EQ 2000 08:08:08.08 31:31:31.31 LSR 0.0"   !**! Source
!
LET N_SUBSCANS 45               !* ! Scan length (in seconds)
LET N_SCANS 30                  !* ! Number of scans on SOURCE (22.5 minutes = 30*45sec)
LET N_SOURCES 1                 !* ! use SYMBOL NAME if N_SOURCES.EQ.1
IF (N_SOURCES.GT.1) THEN
  LET NAME_SOURCE[1:N_SOURCES]   %NAME_SOURCE% !* Enter list of sources (maximum 30)
  LET N_SCANS_SOURCE[1:N_SOURCES] .. ..    !* Enter time per source (in scans)
ENDIF
!
```

-:--- **setup-temp.obs** Top L65 (Fundamental)

Beginning of buffer

Calibrators

2 for Mapping Projects
1 for Detection Projects

setup-temp.obs - emacs@krips.iram.fr

File Edit Options Buffers Tools Help

```
ENDIF
!
LET CALIBRATOR_1 "0923+392"        !**! flux=2.40Jy@3mm, dist=10.13deg, 08-MAR-2018
LET CALIBRATOR_2 "0851+202"        !**! flux=4.90Jy@3mm, dist=11.10deg, 19-MAR-2018
LET CALIBRATOR_3 "0827+243"        !* ! flux=0.60Jy@3mm, dist=8.90deg, 14-DEC-2015
LET N_CALIBRATORS 2                !* ! Use 2 phase calibrators every N_SCANS
LET N_SUBS_CAL  45                 !* ! Scan length on calibrator (in seconds)
LET N_SCANS_CAL  3                 !* ! Nb scans on each calibrator (3 scans)

LET N_MOSAIC 0                     !* ! No mosaic mode
IF (N_MOSAIC.NE.0) THEN
   DEFINE REAL X_MOSAIC[N_MOSAIC] Y_MOSAIC[N_MOSAIC] T_MOSAIC[N_MOSAIC] /GLOBAL
   LET X_MOSAIC .. ..              !* ! offsets in arcsec
   LET Y_MOSAIC .. ..              !* ! offsets in arcsec
   LET T_MOSAIC .. ..              !* ! in units of N_SUBSCANS
ENDIF
!
LET SOLVE_POINT YES
LET SOLVE_FOCUS YES
!
LET FOCUS_RECEIVER 'RECEIVER'      !* ! Focusing on observing receiver
LET POINT_RECEIVER 'RECEIVER'      !* ! Pointing on observing receiver
!
LET POINT_SOURCE_1 "0923+392"       !**! 1st pointing source
LET POINT_SOURCE_2 "0851+202"       !* ! 2nd pointing source
LET FOCUS_SOURCE_1 "0923+392"       !**! 1st focusing source
LET FOCUS_SOURCE_2 "0851+202"       !* ! 2nd focusing source
!
!
LINE MYLINE 114.490000 USB 9490.000 /RECEIVER 1
BASEBAND HLO 1 /RECEIVER 1
BASEBAND HLI 1 /RECEIVER 1
BASEBAND HUI 1 /RECEIVER 1
BASEBAND HUO 1 /RECEIVER 1
BASEBAND VLO 1 /RECEIVER 1
BASEBAND VLI 1 /RECEIVER 1
BASEBAND VUI 1 /RECEIVER 1
BASEBAND VUO 1 /RECEIVER 1
!
IF (N_SOURCES.GT.1) THEN
  SOURCE 'NAME_SOURCE[1]' /TYPE OBJ
 ELSE
  SOURCE 'NAME' /TYPE OBJ
ENDIF
!
SET\RECE 'RECEIVER' ! Choose receiver band for the observation
!
SET\OBS
!
LOAD /TUNING                          ! Load frequency, but don't move antenna now
!
! Make sure any changes in the spectral configuration will be detected:
SET\LOCK
!
LET CHANGE_SPECTRAL .FALSE.       !* ! .TRUE. if need to switch to broad_band !(not needed with PolyFiX)
!
IF (CHANGE_SPECTRAL) THEN
   SPECTRAL /BROAD
ENDIF
!
SET SHOW OFF
!
TYPE PR:clean.obs
```

-:--- **setup-temp.obs** 50% L98 (Fundamental)

How to use NOEMA

setup-temp.obs - emacs@krips.iram.fr

File Edit Options Buffers Tools Help

```
!------------------------------------------------------------------------
!
! PR:SETUP-W18ZZ001.OBS ! Setup procedure for project "W18ZZ001"
!
!    - Date:          2018-05-31
!    - Author:        LC
!    - PI:            PI
!    - Local contact: LC
!    - Project ID:    W18ZZ001
!    - Verified by:   PI (date)
!
!    - Rating:                     B
!    - Number of telescopes assumed: 8
!    - Observing mode:             Mapping
!    - Observing goal:             1 lines, 1 continuum
!    - Time:                       [0.00,0.00,0.00,5.00,0.00]
!                                  (total hr os/configuration, all sources/fields combined)
!
!    - Requested on-source time (h): 5.00
!
!    - Requested sensitivity:      1.0mJy/5000.0kHz
!    - Additional sensitivity:     0.02mJy/15488.0MHz
!
!    - Requested minimum S/N:
!    - Reference Frequency (GHz):  114.490
!
!    - Water limit:
!    - Obs date constraint:
!
!    - Sun avoidance:              29-JUN-2018 to 31-AUG-2018 for MYSOURCE
!
!    - Other comments:             RepFreq: 114.490 GHz USB.
!
!
!------------------------------------------------------------------------
! Do not edit directly, but copy first then
!
!       All lines marked !**! must be customized.
!           lines marked !* ! can be modified.
!
!------------------------------------------------------------------------
SET\END                         ! Finish previous observation
@ PR:defaults                   ! Restore defaults parameters
!
SET\PROJECT W18ZZ001            !**! Specify project number for further
!
SYMBOL GO "@ PR:observe-all W18ZZ001" !* ! data processing
CATA SOU INTER_BASE:iram.sou !* !
CATA PHA INTER_BASE:phase-pdb.sou !* !
LET RECEIVER 1                  !**! Choose observing receiver: receiver 1 @ 3mm
!                                                               receiver 2 @ 2mm
!                                                               receiver 3 @ 1mm
!
LET LOW_LIMIT 15.               !* ! Low elevation limit 15 degrees
SAY "Project 'PROJECT' starting"
!
SYMBOL NAME "MYSOURCE        EQ 2000 08:08:08.08 31:31:31.31 LSR 0.0"   !**! Source
!
LET N_SUBSCANS 45               !* ! Scan length (in seconds)
LET N_SCANS 30                  !* ! Number of scans on SOURCE (22.5 minutes = 30*45sec)
LET N_SOURCES 1                 !* ! use SYMBOL NAME if N_SOURCES.EQ.1
IF (N_SOURCES.GT.1) THEN
  LET NAME_SOURCE[1:N_SOURCES]   %NAME_SOURCE% !* Enter list of sources (maximum 30)
  LET N_SCANS_SOURCE[1:N_SOURCES] .. ..    !* Enter time per source (in scans)
ENDIF
!
```

-:--- **setup-temp.obs** Top L65 (Fundamental)

Beginning of buffer

setup-temp.obs - emacs@krips.iram.fr

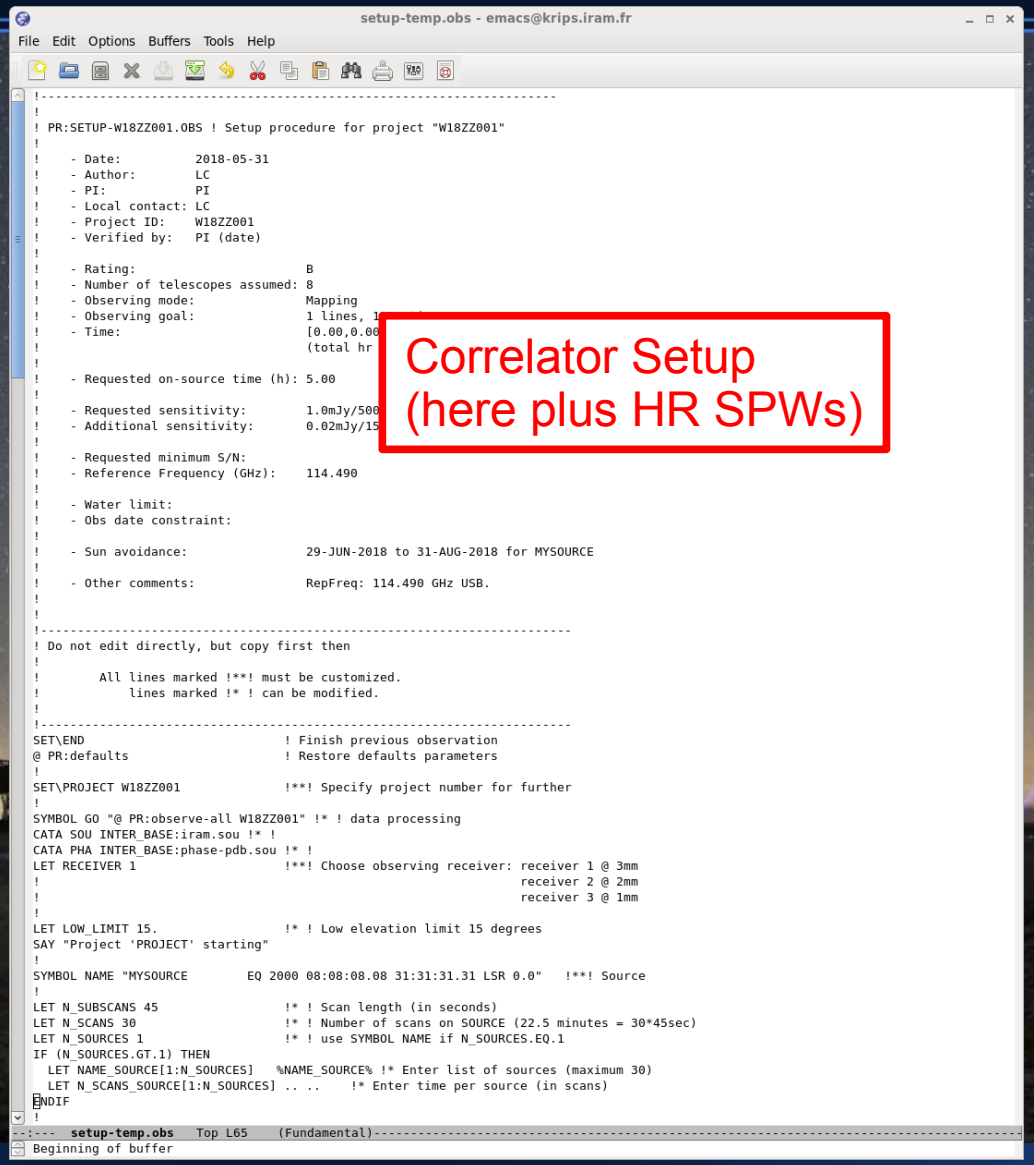
File Edit Options Buffers Tools Help

```
ENDIF
!
LET CALIBRATOR_1 "0923+392"        !**! flux=2.40Jy@3mm, dist=10.13deg, 08-MAR-2018
LET CALIBRATOR_2 "0851+202"        !**! flux=4.90Jy@3mm, dist=11.10deg, 19-MAR-2018
LET CALIBRATOR_3 "0827+243"        !* ! flux=0.60Jy@3mm, dist=8.90deg, 14-DEC-2015
LET N_CALIBRATORS 2                !* ! Use 2 phase calibrators every N_SCANS
LET N_SUBS_CAL  45                 !* ! Scan length on calibrator (in seconds)
LET N_SCANS_CAL  3                 !* ! Nb scans on each calibrator (3 scans)

LET FSWI_CAL  .FALSE.              !* ! No fast-switching by default

LET N_MOSAIC 0                     !* ! No mosaic mode
IF (N_MOSAIC.NE.0) THEN
                                                 IC[N_MOSAIC] T_MOSAIC[N_MOSAIC] /GLOBAL
                                                 offsets in arcsec
                                                 offsets in arcsec
                                                 in units of N_SUBSCANS

LET FOCUS_RECEIVER 'RECEIVER'      !* ! Focusing on observing receiver
LET POINT_RECEIVER 'RECEIVER'      !* ! Pointing on observing receiver
!
LET POINT_SOURCE_1 "0923+392"       !**! 1st pointing source
LET POINT_SOURCE_2 "0851+202"       !* ! 2nd pointing source
LET FOCUS_SOURCE_1 "0923+392"       !**! 1st focusing source
LET FOCUS_SOURCE_2 "0851+202"       !* ! 2nd focusing source
!
SET\UNLOCK
!
LINE MYLINE 114.490000 USB 9490.000 /RECEIVER 1
BASEBAND HLO 1 /RECEIVER 1
BASEBAND HLI 1 /RECEIVER 1
BASEBAND HUI 1 /RECEIVER 1
BASEBAND HUO 1 /RECEIVER 1
BASEBAND VLO 1 /RECEIVER 1
BASEBAND VLI 1 /RECEIVER 1
BASEBAND VUI 1 /RECEIVER 1
BASEBAND VUO 1 /RECEIVER 1
!
  SOURCE 'NAME_SOURCE[1]' /TYPE OBJ
 ELSE
  SOURCE 'NAME' /TYPE OBJ
ENDIF
!
SET\RECE 'RECEIVER' ! Choose receiver band for the observation
!
SET\OBS
!
LOAD /TUNING                           ! Load frequency, but don't move antenna now
!
! Make sure any changes in the spectral configuration will be detected:
SET\LOCK
!
LET CHANGE_SPECTRAL .FALSE.        !* ! .TRUE. if need to switch to broad_band !(not needed with PolyFiX)
!
IF (CHANGE_SPECTRAL) THEN
   SPECTRAL /BROAD
ENDIF
!
SET SHOW OFF
!
TYPE PR:clean.obs
```

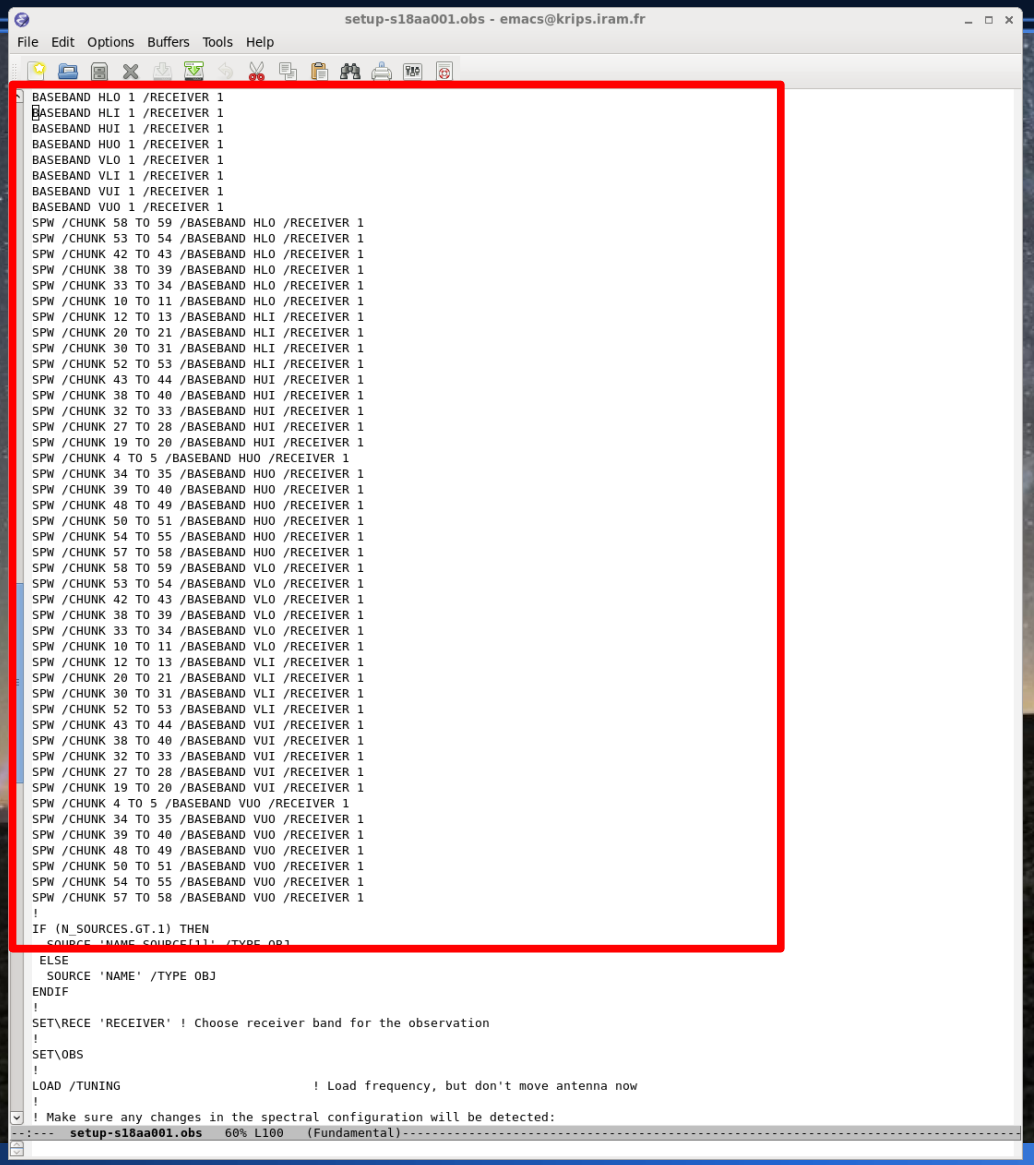
-:--- **setup-temp.obs** 50% L98 (Fundamental)

Correlator Setup
(here only LR basebands)

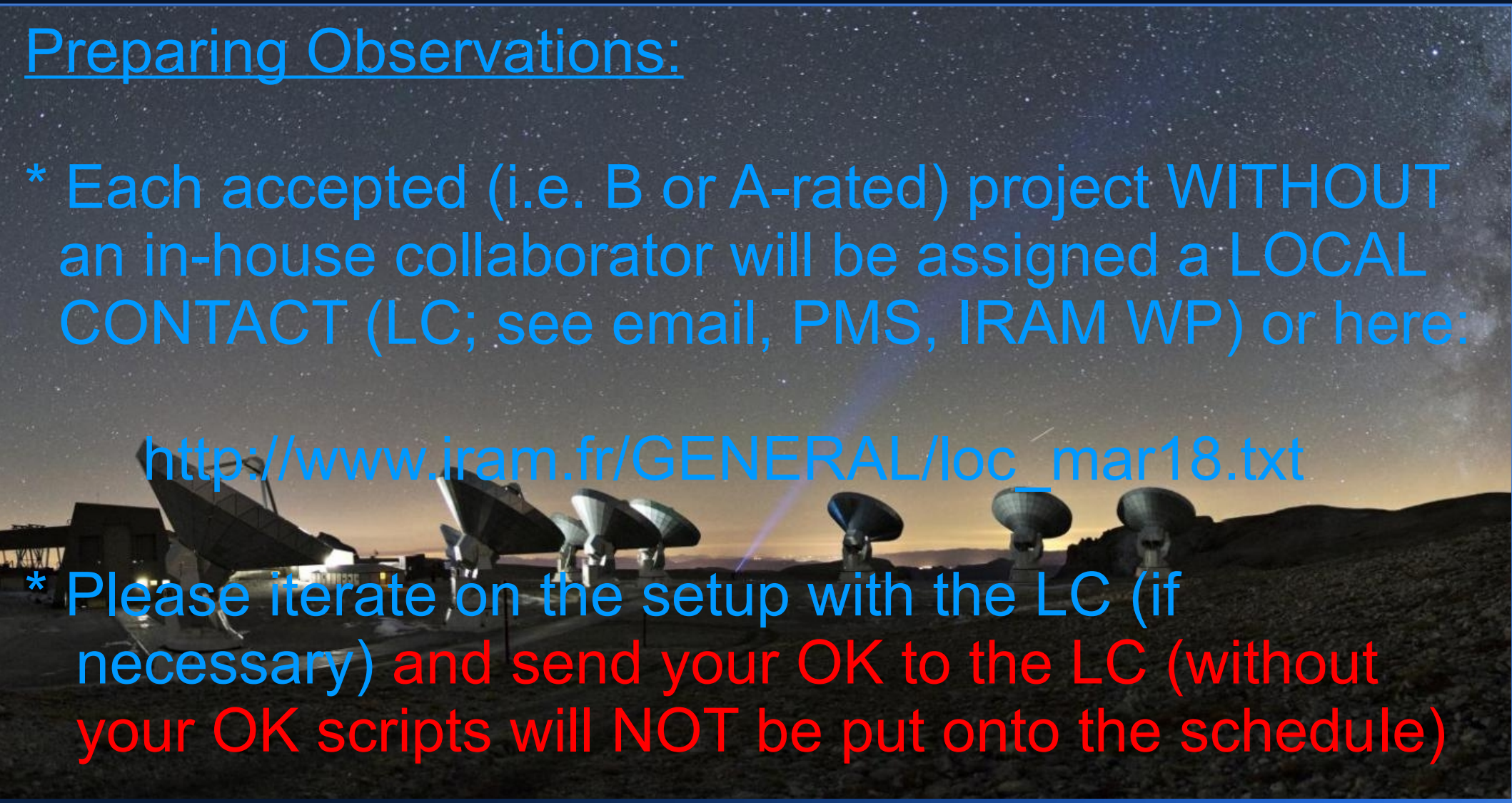
How to use NOEMA



Correlator Setup
(here plus HR SPWs)



How to use NOEMA



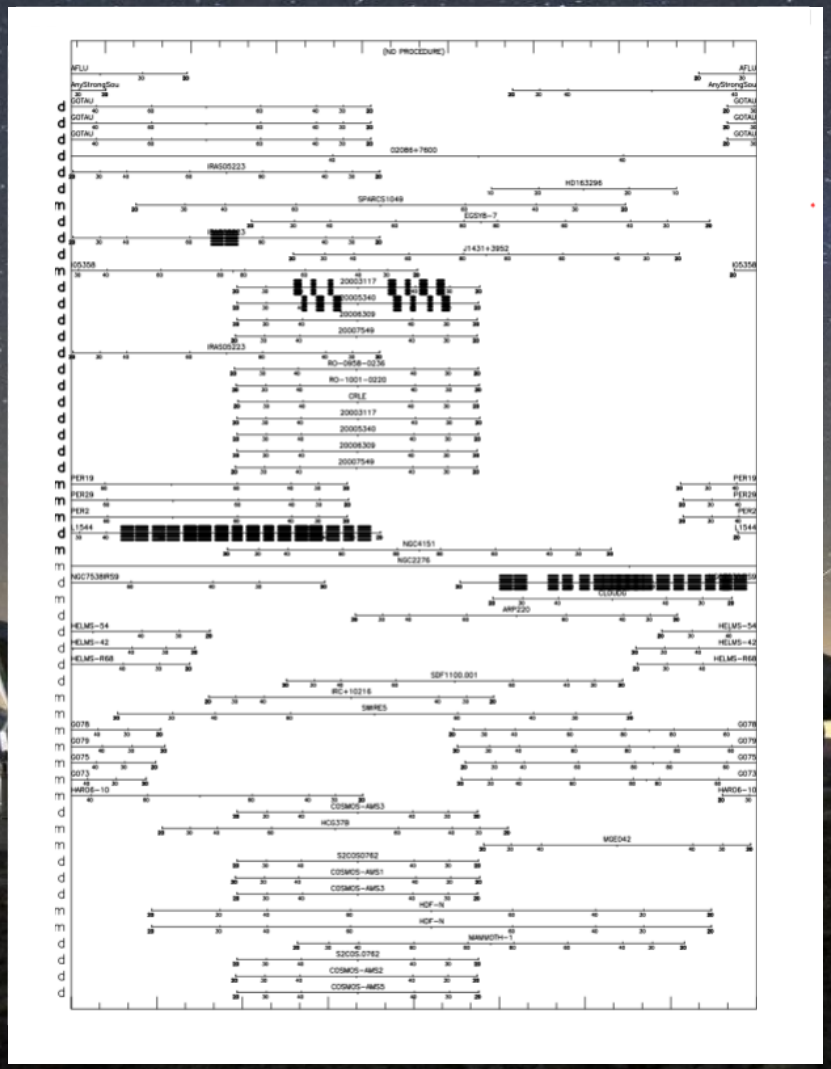
Preparing Observations:

* Each accepted (i.e. B or A-rated) project WITHOUT an in-house collaborator will be assigned a LOCAL CONTACT (LC; see email, PMS, IRAM WP) or here:

http://www.iram.fr/GENERAL/loc_mar18.txt

* Please iterate on the setup with the LC (if necessary) and send your OK to the LC (without your OK scripts will NOT be put onto the schedule)

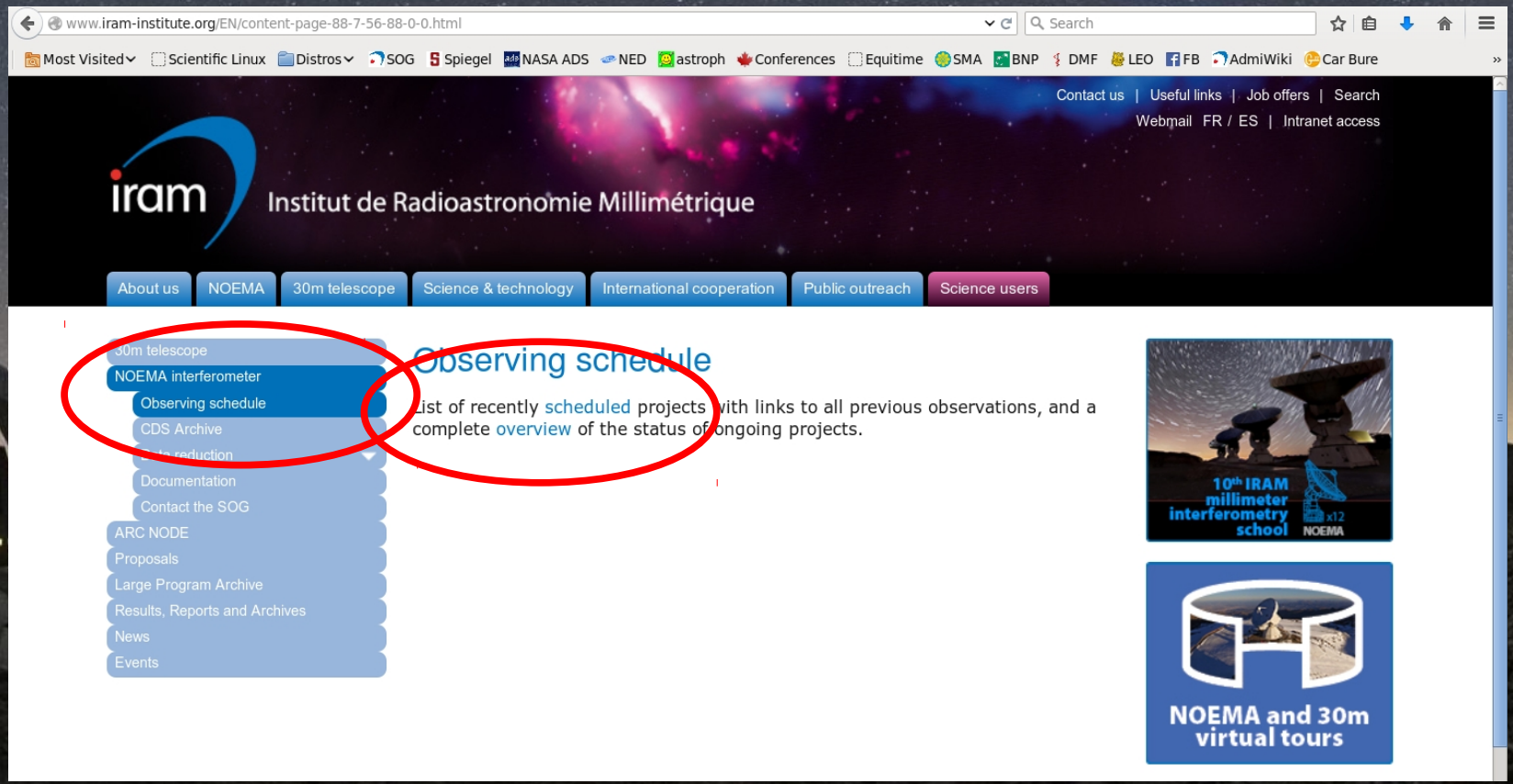
How to use NOEMA



How to use NOEMA

Observations:

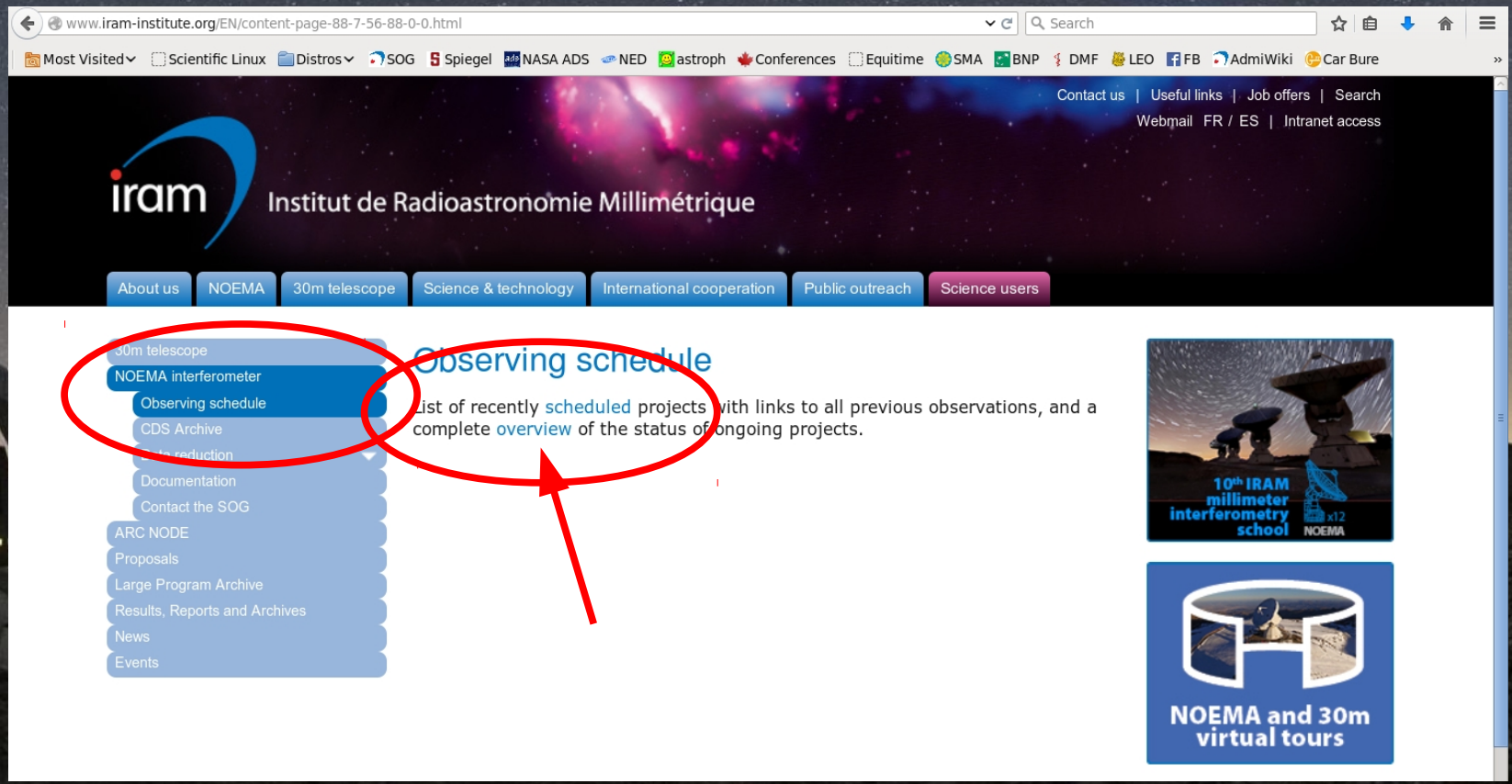
* Check Status of your project



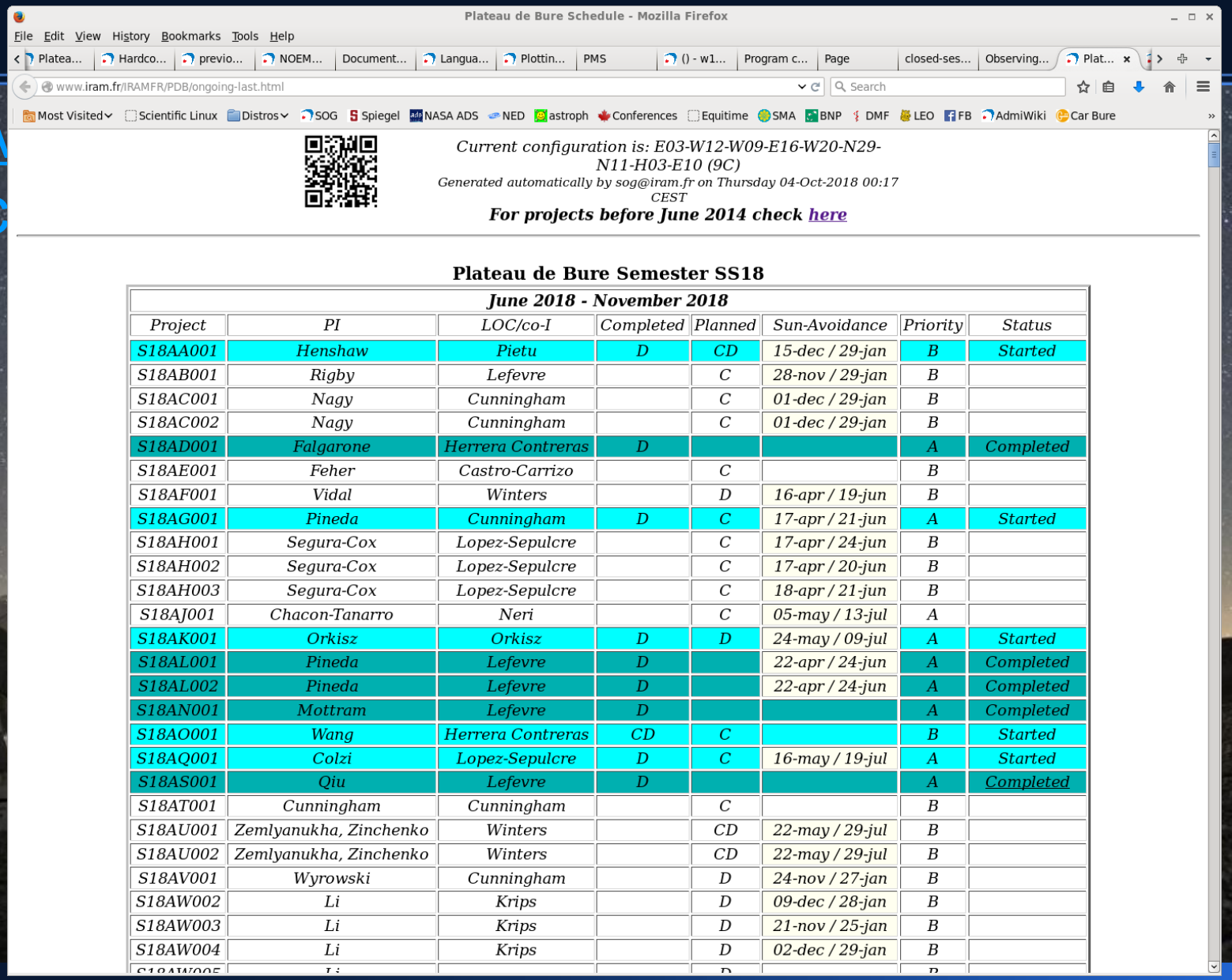
How to use NOEMA

Observations:

* Check Status of your project



How to use NOEMA



Current configuration is: E03-W12-W09-E16-W20-N29-N11-H03-E10 (9C)

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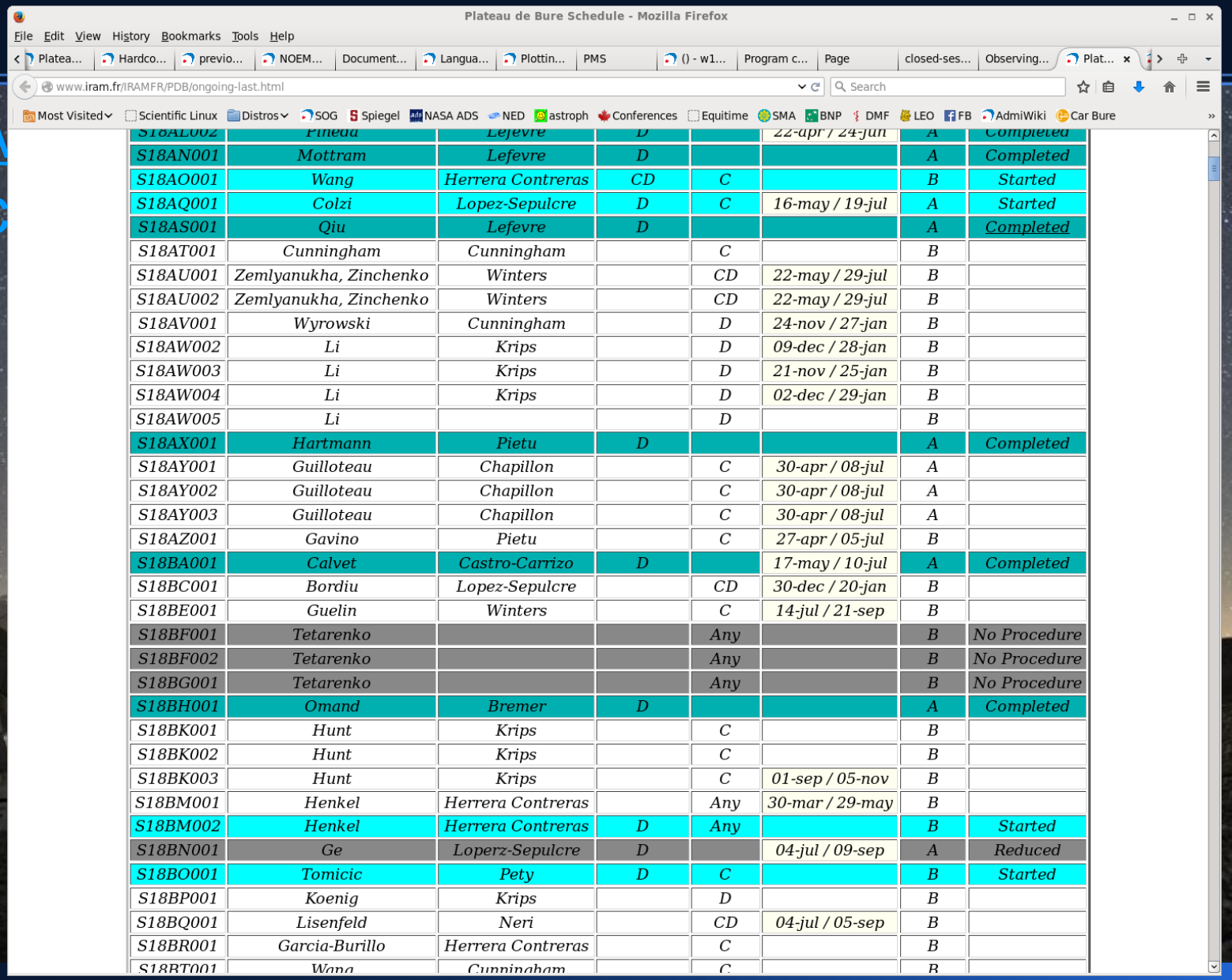
For projects before June 2014 check here

Plateau de Bure Semester SS18

June 2018 - November 2018							
Project	*PI*	*LOC/co-I*	*Completed*	*Planned*	*Sun-Avoidance*	*Priority*	*Status*
S18AA001	*Henshaw*	*Pietu*	*D*	*CD*	*15-dec / 29-jan*	*B*	*Started*
S18AB001	*Rigby*	*Lefevre*		*C*	*28-nov / 29-jan*	*B*	
S18AC001	*Nagy*	*Cunningham*		*C*	*01-dec / 29-jan*	*B*	
S18AC002	*Nagy*	*Cunningham*		*C*	*01-dec / 29-jan*	*B*	
S18AD001	*Falgarone*	*Herrera Contreras*	*D*			*A*	*Completed*
S18AE001	*Feher*	*Castro-Carrizo*		*C*		*B*	
S18AF001	*Vidal*	*Winters*		*D*	*16-apr / 19-jun*	*B*	
S18AG001	*Pineda*	*Cunningham*	*D*	*C*	*17-apr / 21-jun*	*A*	*Started*
S18AH001	*Segura-Cox*	*Lopez-Sepulcre*		*C*	*17-apr / 24-jun*	*B*	
S18AH002	*Segura-Cox*	*Lopez-Sepulcre*		*C*	*17-apr / 20-jun*	*B*	
S18AH003	*Segura-Cox*	*Lopez-Sepulcre*		*C*	*18-apr / 21-jun*	*B*	
S18AJ001	*Chacon-Tanarro*	*Neri*		*C*	*05-may / 13-jul*	*A*	
S18AK001	*Orkisz*	*Orkisz*	*D*	*D*	*24-may / 09-jul*	*A*	*Started*
S18AL001	*Pineda*	*Lefevre*	*D*		*22-apr / 24-jun*	*A*	*Completed*
S18AL002	*Pineda*	*Lefevre*	*D*		*22-apr / 24-jun*	*A*	*Completed*
S18AN001	*Mottram*	*Lefevre*	*D*			*A*	*Completed*
S18AO001	*Wang*	*Herrera Contreras*	*CD*	*C*		*B*	*Started*
S18AQ001	*Colzi*	*Lopez-Sepulcre*	*D*	*C*	*16-may / 19-jul*	*A*	*Started*
S18AS001	*Qiu*	*Lefevre*	*D*			*A*	*Completed*
S18AT001	*Cunningham*	*Cunningham*		*C*		*B*	
S18AU001	*Zemlyanukha, Zinchenko*	*Winters*		*CD*	*22-may / 29-jul*	*B*	
S18AU002	*Zemlyanukha, Zinchenko*	*Winters*		*CD*	*22-may / 29-jul*	*B*	
S18AV001	*Wyrowski*	*Cunningham*		*D*	*24-nov / 27-jan*	*B*	
S18AW002	*Li*	*Krips*		*D*	*09-dec / 28-jan*	*B*	
S18AW003	*Li*	*Krips*		*D*	*21-nov / 25-jan*	*B*	
S18AW004	*Li*	*Krips*		*D*	*02-dec / 29-jan*	*B*	

How to use NOEMA

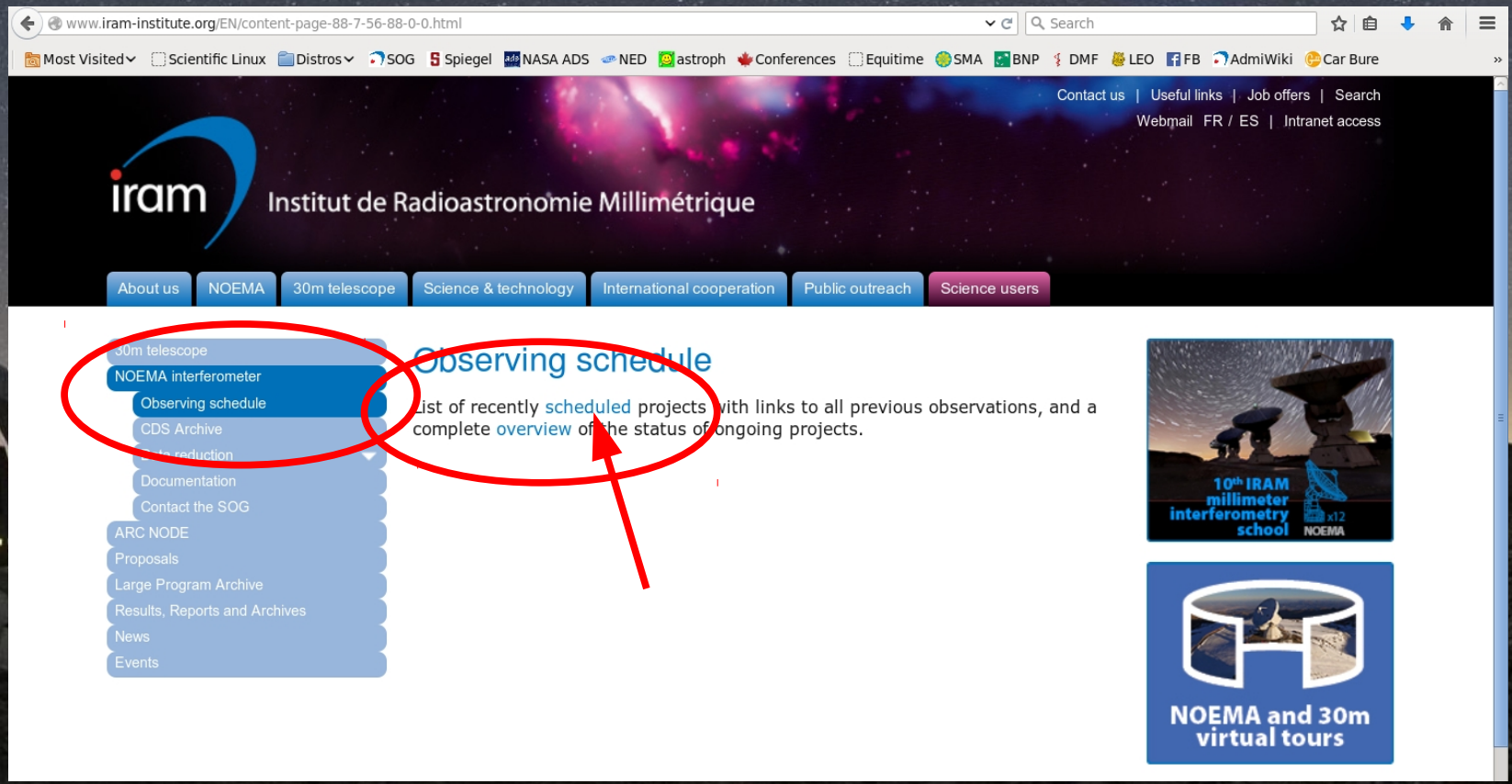


S18AL002	Pineda	Lefevre	D		22-apr / 24-jun	A	Completed
S18AN001	Mottram	Lefevre	D			A	Completed
S18AO001	Wang	Herrera Contreras	CD	C		B	Started
S18AQ001	Colzi	Lopez-Sepulcre	D	C	16-may / 19-jul	A	Started
S18AS001	Qiu	Lefevre	D			A	Completed
S18AT001	Cunningham	Cunningham		C		B	
S18AU001	Zemlyanukha, Zinchenko	Winters		CD	22-may / 29-jul	B	
S18AU002	Zemlyanukha, Zinchenko	Winters		CD	22-may / 29-jul	B	
S18AV001	Wyrowski	Cunningham		D	24-nov / 27-jan	B	
S18AW002	Li	Krips		D	09-dec / 28-jan	B	
S18AW003	Li	Krips		D	21-nov / 25-jan	B	
S18AW004	Li	Krips		D	02-dec / 29-jan	B	
S18AW005	Li			D		B	
S18AX001	Hartmann	Pietu	D			A	Completed
S18AY001	Guilloteau	Chapillon		C	30-apr / 08-jul	A	
S18AY002	Guilloteau	Chapillon		C	30-apr / 08-jul	A	
S18AY003	Guilloteau	Chapillon		C	30-apr / 08-jul	A	
S18AZ001	Gavino	Pietu		C	27-apr / 05-jul	B	
S18BA001	Calvet	Castro-Carrizo	D		17-may / 10-jul	A	Completed
S18BC001	Bordiu	Lopez-Sepulcre		CD	30-dec / 20-jan	B	
S18BE001	Guelin	Winters		C	14-jul / 21-sep	B	
S18BF001	Tetarenko			Any		B	No Procedure
S18BF002	Tetarenko			Any		B	No Procedure
S18BG001	Tetarenko			Any		B	No Procedure
S18BH001	Omand	Bremer	D			A	Completed
S18BK001	Hunt	Krips		C		B	
S18BK002	Hunt	Krips		C		B	
S18BK003	Hunt	Krips		C	01-sep / 05-nov	B	
S18BM001	Henkel	Herrera Contreras		Any	30-mar / 29-may	B	
S18BM002	Henkel	Herrera Contreras	D	Any		B	Started
S18BN001	Ge	Loperz-Sepulcre	D		04-jul / 09-sep	A	Reduced
S18BO001	Tomicic	Pety	D	C		B	Started
S18BP001	Koenig	Krips		D		B	
S18BQ001	Lisenfeld	Neri		CD	04-jul / 05-sep	B	
S18BR001	Garcia-Burillo	Herrera Contreras		C		B	
S18BT001	Wang	Cunningham		C		B	

How to use NOEMA

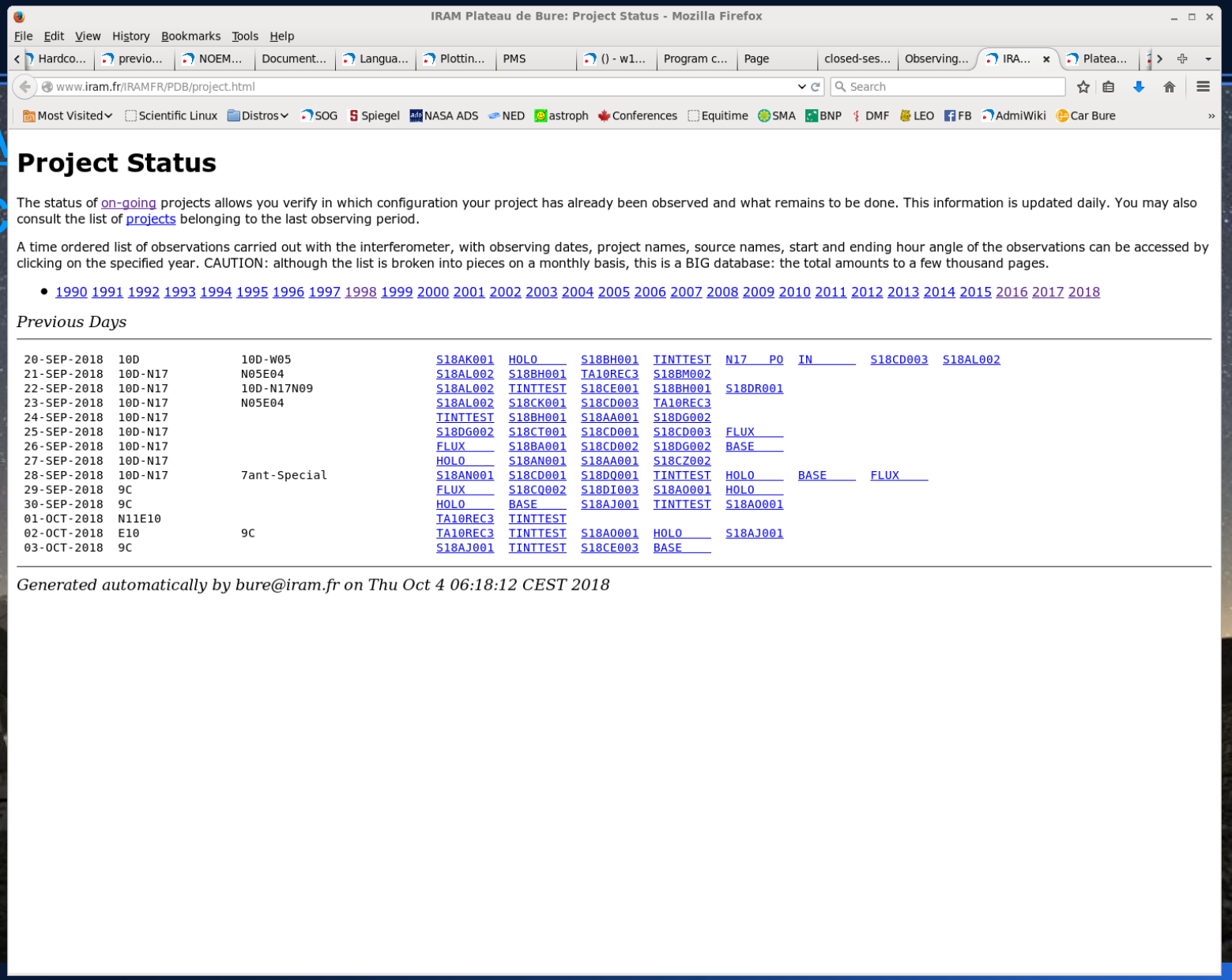
Observations:

* Check Status of your project



How to use NOEMA

Observ
* Chec



Project Status

The status of on-going projects allows you verify in which configuration your project has already been observed and what remains to be done. This information is updated daily. You may also consult the list of projects belonging to the last observing period.

A time ordered list of observations carried out with the interferometer, with observing dates, project names, source names, start and ending hour angle of the observations can be accessed by clicking on the specified year. CAUTION: although the list is broken into pieces on a monthly basis, this is a BIG database: the total amounts to a few thousand pages.

- 1990 1991 1992 1993 1994 1995 1996 1997 1998 1999 2000 2001 2002 2003 2004 2005 2006 2007 2008 2009 2010 2011 2012 2013 2014 2015 2016 2017 2018

Previous Days

20-SEP-2018	10D	10D-W05	S18AK001	HOLO___	S18BH001	TINTTEST	N17__PO	IN_____	S18CD003	S18AL002
21-SEP-2018	10D-N17	N05E04	S18AL002	S18BH001	TA10REC3	S18BM002				
22-SEP-2018	10D-N17	10D-N17N09	S18AL002	TINTTEST	S18CE001	S18BH001	S18DR001			
23-SEP-2018	10D-N17	N05E04	S18AL002	S18CK001	S18CD003	TA10REC3				
24-SEP-2018	10D-N17		TINTTEST	S18BH001	S18AA001	S18DG002				
25-SEP-2018	10D-N17		S18DG002	S18CT001	S18CD001	S18CD003	FLUX___			
26-SEP-2018	10D-N17		FLUX___	S18BA001	S18CD002	S18DG002	BASE___			
27-SEP-2018	10D-N17		HOLO___	S18AN001	S18AA001	S18CZ002				
28-SEP-2018	10D-N17	7ant-Special	S18AN001	S18CD001	S18DQ001	TINTTEST	HOLO___	BASE___	FLUX___	
29-SEP-2018	9C		FLUX___	S18CQ002	S18DI003	S18AO001	HOLO___			
30-SEP-2018	9C		HOLO___	BASE___	S18AJ001	TINTTEST	S18AO001			
01-OCT-2018	N11E10		TA10REC3	TINTTEST						
02-OCT-2018	E10	9C	TA10REC3	TINTTEST	S18AO001	HOLO___	S18AJ001			
03-OCT-2018	9C		S18AJ001	TINTTEST	S18CE003	BASE___				

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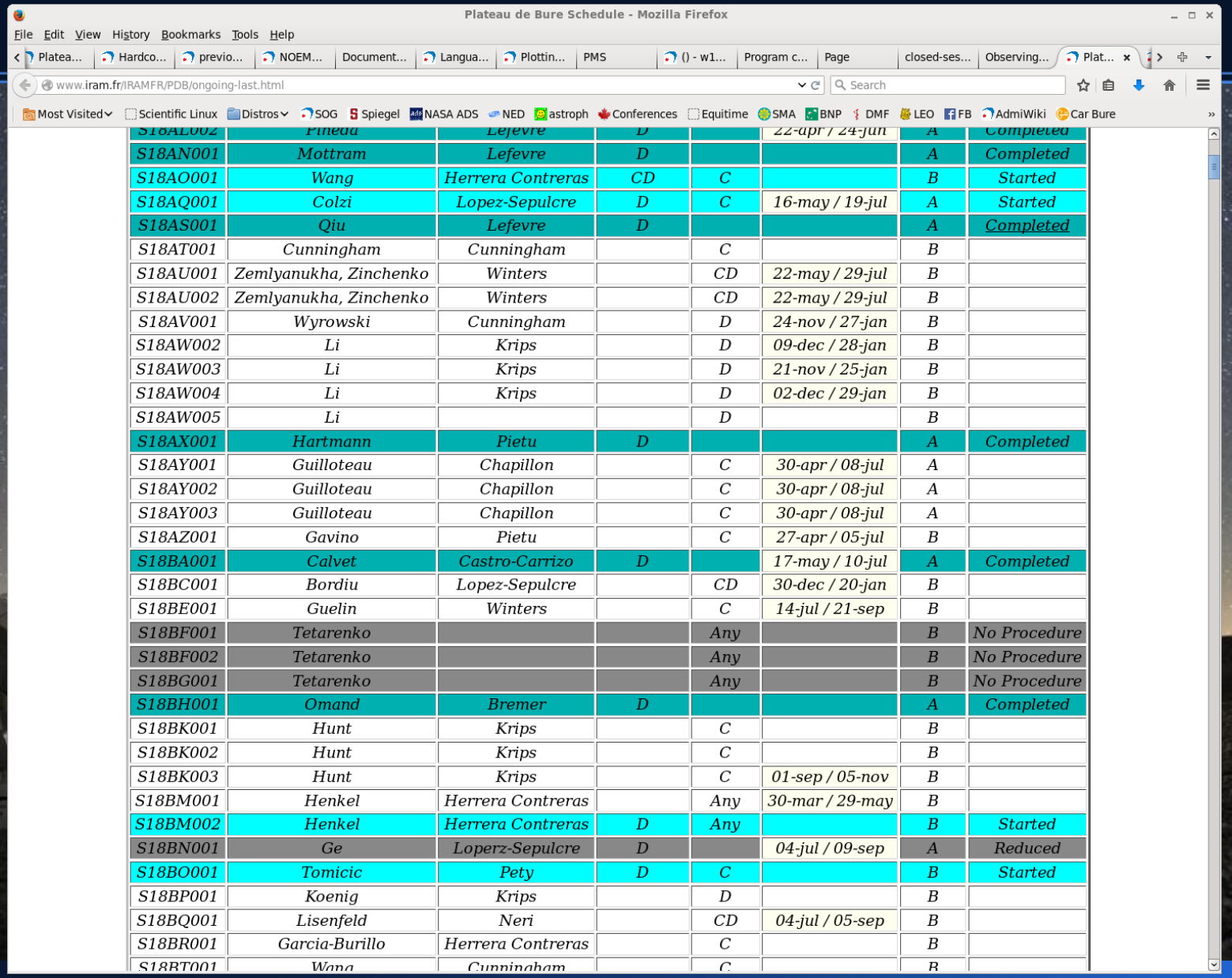
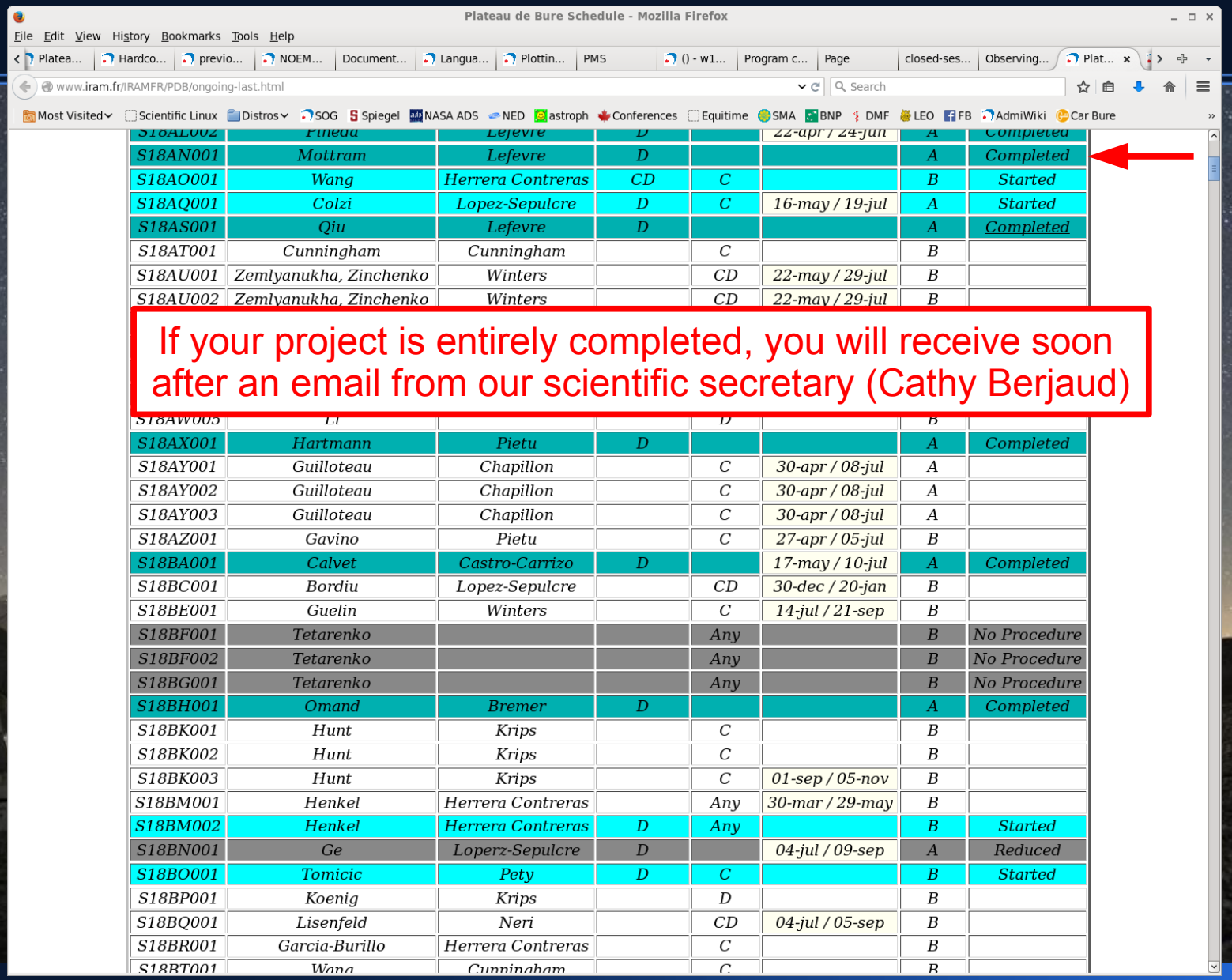
How to use NOEMA

How to use NOEMA



S18AL002	Pineda	Lefevre	D		22-apr / 24-jun	A	Completed
S18AN001	Mottram	Lefevre	D			A	Completed
S18AO001	Wang	Herrera Contreras	CD	C		B	Started
S18AQ001	Colzi	Lopez-Sepulcre	D	C	16-may / 19-jul	A	Started
S18AS001	Qiu	Lefevre	D			A	Completed
S18AT001	Cunningham	Cunningham		C		B	
S18AU001	Zemlyanukha, Zinchenko	Winters		CD	22-may / 29-jul	B	
S18AU002	Zemlyanukha, Zinchenko	Winters		CD	22-may / 29-jul	B	
S18AV001	Wyrowski	Cunningham		D	24-nov / 27-jan	B	
S18AW002	Li	Krips		D	09-dec / 28-jan	B	
S18AW003	Li	Krips		D	21-nov / 25-jan	B	
S18AW004	Li	Krips		D	02-dec / 29-jan	B	
S18AW005	Li			D		B	
S18AX001	Hartmann	Pietu	D			A	Completed
S18AY001	Guilloteau	Chapillon		C	30-apr / 08-jul	A	
S18AY002	Guilloteau	Chapillon		C	30-apr / 08-jul	A	
S18AY003	Guilloteau	Chapillon		C	30-apr / 08-jul	A	
S18AZ001	Gavino	Pietu		C	27-apr / 05-jul	B	
S18BA001	Calvet	Castro-Carrizo	D		17-may / 10-jul	A	Completed
S18BC001	Bordiu	Lopez-Sepulcre		CD	30-dec / 20-jan	B	
S18BE001	Guelin	Winters		C	14-jul / 21-sep	B	
S18BF001	Tetarenko			Any		B	No Procedure
S18BF002	Tetarenko			Any		B	No Procedure
S18BG001	Tetarenko			Any		B	No Procedure
S18BH001	Omand	Bremer	D			A	Completed
S18BK001	Hunt	Krips		C		B	
S18BK002	Hunt	Krips		C		B	
S18BK003	Hunt	Krips		C	01-sep / 05-nov	B	
S18BM001	Henkel	Herrera Contreras		Any	30-mar / 29-may	B	
S18BM002	Henkel	Herrera Contreras	D	Any		B	Started
S18BN001	Ge	Loperz-Sepulcre	D		04-jul / 09-sep	A	Reduced
S18BO001	Tomicic	Pety	D	C		B	Started
S18BP001	Koenig	Krips		D		B	
S18BQ001	Lisenfeld	Neri		CD	04-jul / 05-sep	B	
S18BR001	Garcia-Burillo	Herrera Contreras		C		B	
S18BT001	Wang	Cunningham		C		B	

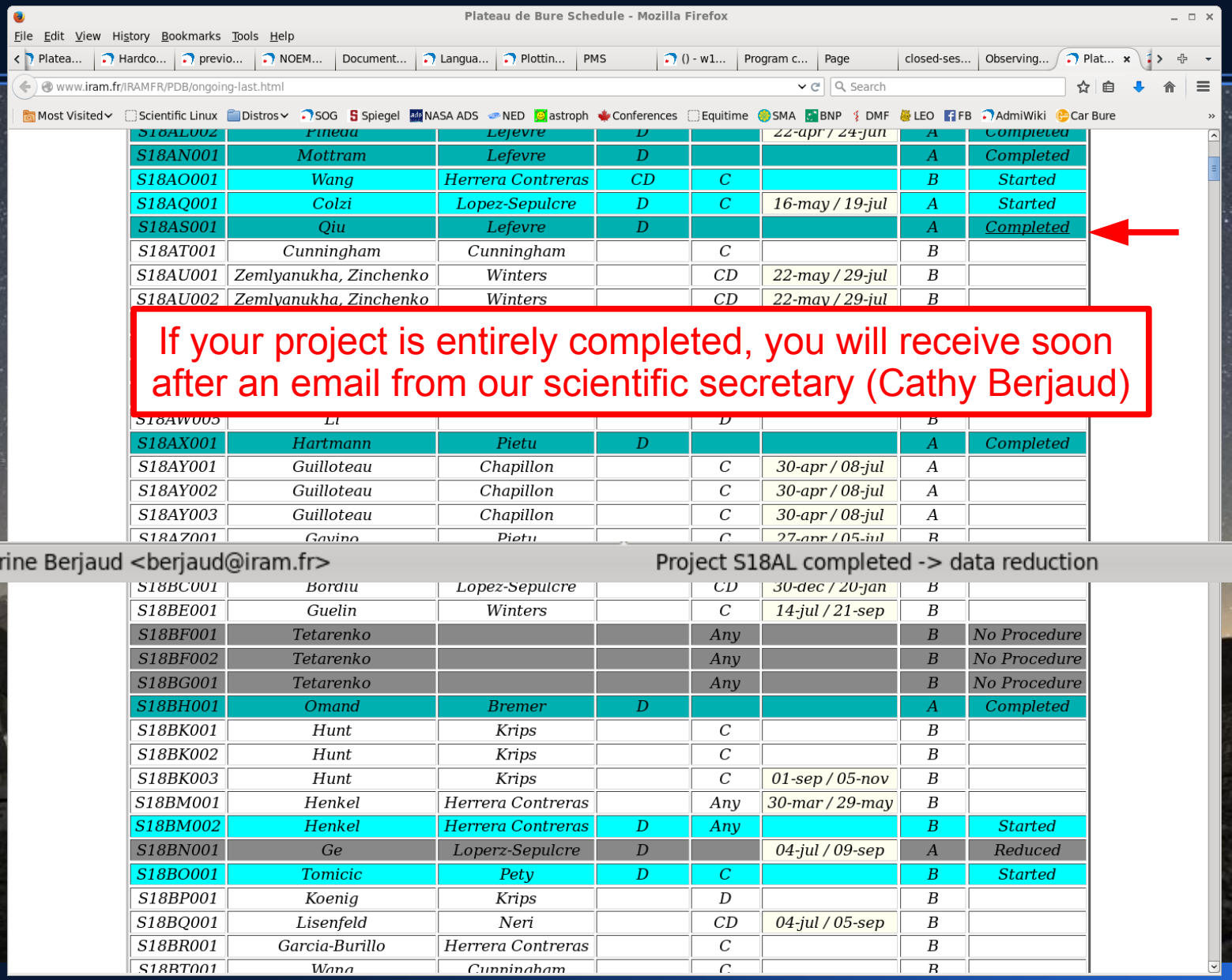
How to use NOEMA



S18AN001	Mottram	Lefevre	D			A	Completed
S18AO001	Wang	Herrera Contreras	CD	C		B	Started
S18AQ001	Colzi	Lopez-Sepulcre	D	C	16-may / 19-jul	A	Started
S18AS001	Qiu	Lefevre	D			A	Completed
S18AT001	Cunningham	Cunningham		C		B	
S18AU001	Zemlyanukha, Zinchenko	Winters		CD	22-may / 29-jul	B	
S18AU002	Zemlyanukha, Zinchenko	Winters		CD	22-may / 29-jul	B	
S18AX001	Hartmann	Pietu	D			A	Completed
S18AY001	Guilloteau	Chapillon		C	30-apr / 08-jul	A	
S18AY002	Guilloteau	Chapillon		C	30-apr / 08-jul	A	
S18AY003	Guilloteau	Chapillon		C	30-apr / 08-jul	A	
S18AZ001	Gavino	Pietu		C	27-apr / 05-jul	B	
S18BA001	Calvet	Castro-Carrizo	D		17-may / 10-jul	A	Completed
S18BC001	Bordiu	Lopez-Sepulcre		CD	30-dec / 20-jan	B	
S18BE001	Guelin	Winters		C	14-jul / 21-sep	B	
S18BF001	Tetarenko			Any		B	No Procedure
S18BF002	Tetarenko			Any		B	No Procedure
S18BG001	Tetarenko			Any		B	No Procedure
S18BH001	Omand	Bremer	D			A	Completed
S18BK001	Hunt	Krips		C		B	
S18BK002	Hunt	Krips		C		B	
S18BK003	Hunt	Krips		C	01-sep / 05-nov	B	
S18BM001	Henkel	Herrera Contreras		Any	30-mar / 29-may	B	
S18BM002	Henkel	Herrera Contreras	D	Any		B	Started
S18BN001	Ge	Loperz-Sepulcre	D		04-jul / 09-sep	A	Reduced
S18BO001	Tomicic	Pety	D	C		B	Started
S18BP001	Koenig	Krips		D		B	
S18BQ001	Lisenfeld	Neri		CD	04-jul / 05-sep	B	
S18BR001	Garcia-Burillo	Herrera Contreras		C		B	

If your project is entirely completed, you will receive soon after an email from our scientific secretary (Cathy Berjaud)

How to use NOEMA



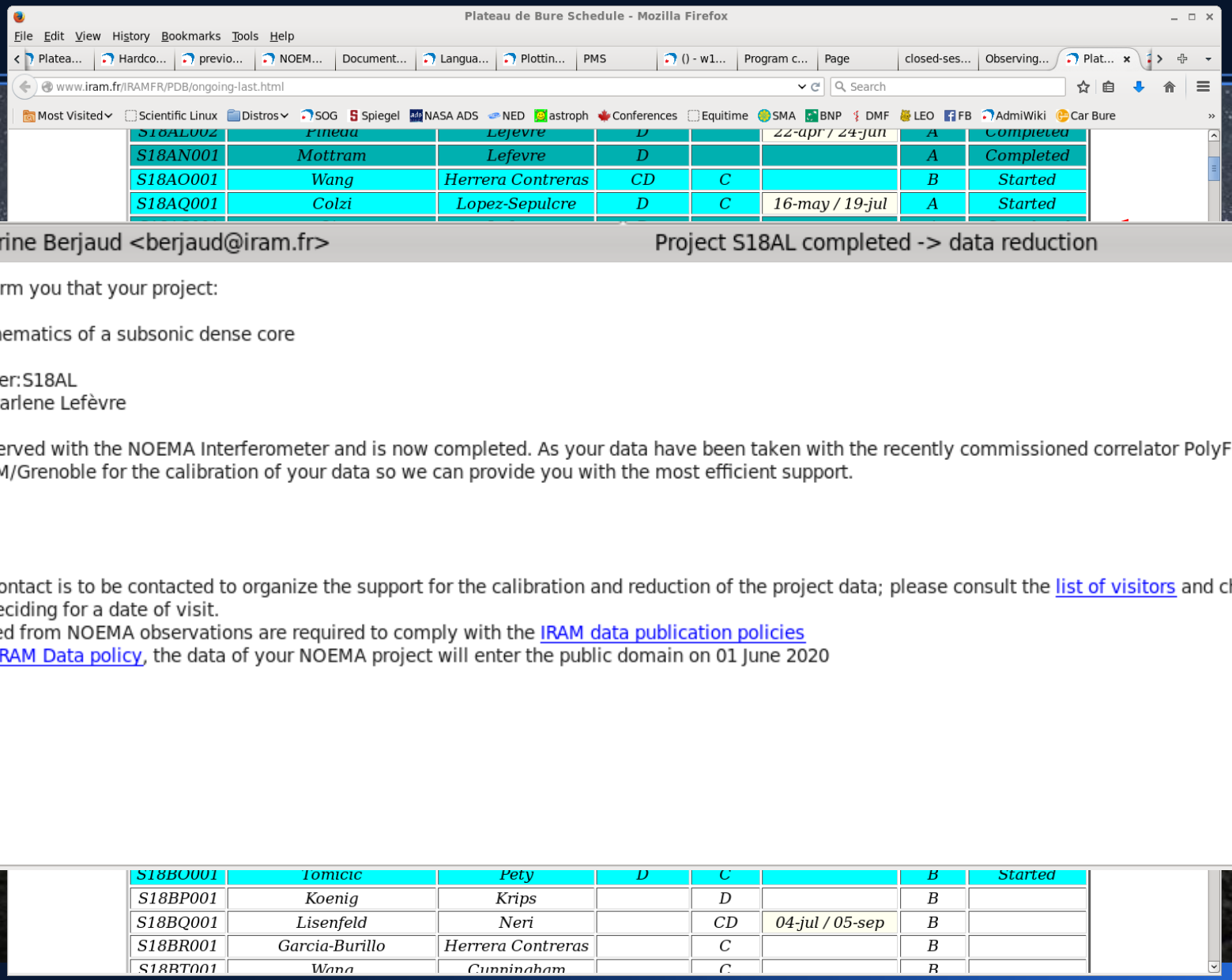
S18AN001	Mottram	Lefevre	D			A	Completed
S18AO001	Wang	Herrera Contreras	CD	C		B	Started
S18AQ001	Colzi	Lopez-Sepulcre	D	C	16-may / 19-jul	A	Started
S18AS001	Qiu	Lefevre	D			A	Completed
S18AT001	Cunningham	Cunningham		C		B	
S18AU001	Zemlyanukha, Zinchenko	Winters		CD	22-may / 29-jul	B	
S18AU002	Zemlyanukha, Zinchenko	Winters		CD	22-may / 29-jul	B	

If your project is entirely completed, you will receive soon after an email from our scientific secretary (Cathy Berjaud)

S18AX001	Hartmann	Pietu	D			A	Completed
S18AY001	Guilloteau	Chapillon		C	30-apr / 08-jul	A	
S18AY002	Guilloteau	Chapillon		C	30-apr / 08-jul	A	
S18AY003	Guilloteau	Chapillon		C	30-apr / 08-jul	A	
S18BE001	Guelin	Winters		C	14-jul / 21-sep	B	
S18BF001	Tetarenko			Any		B	No Procedure
S18BF002	Tetarenko			Any		B	No Procedure
S18BG001	Tetarenko			Any		B	No Procedure
S18BH001	Omand	Bremer	D			A	Completed
S18BK001	Hunt	Krips		C		B	
S18BK002	Hunt	Krips		C		B	
S18BK003	Hunt	Krips		C	01-sep / 05-nov	B	
S18BM001	Henkel	Herrera Contreras		Any	30-mar / 29-may	B	
S18BM002	Henkel	Herrera Contreras	D	Any		B	Started
S18BN001	Ge	Loperz-Sepulcre	D		04-jul / 09-sep	A	Reduced
S18BO001	Tomicic	Pety	D	C		B	Started
S18BP001	Koenig	Krips		D		B	
S18BQ001	Lisenfeld	Neri		CD	04-jul / 05-sep	B	
S18BR001	Garcia-Burillo	Herrera Contreras		C		B	

Catherine Berjaud <berjaud@iram.fr> Project S18AL completed -> data reduction

How to use NOEMA



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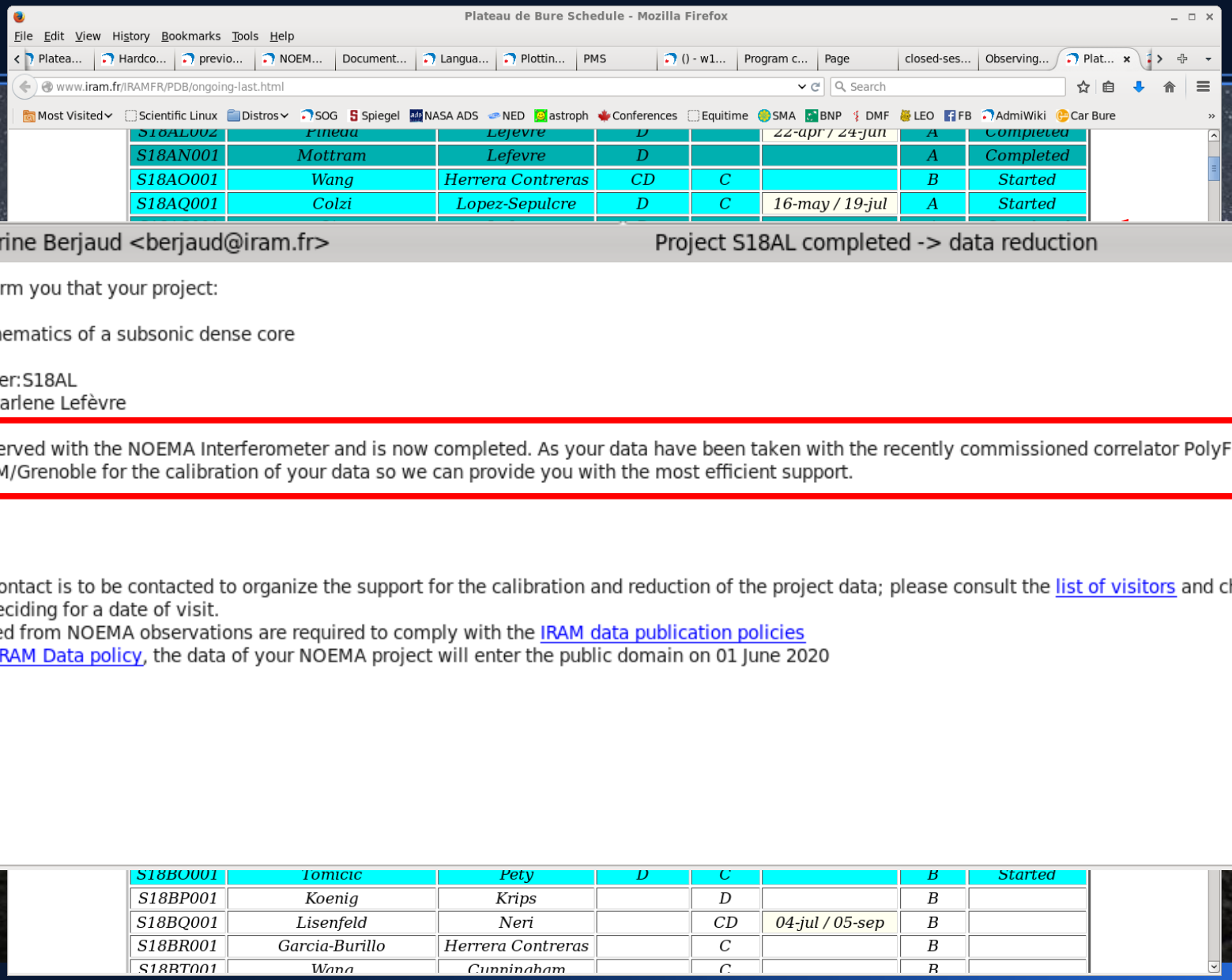

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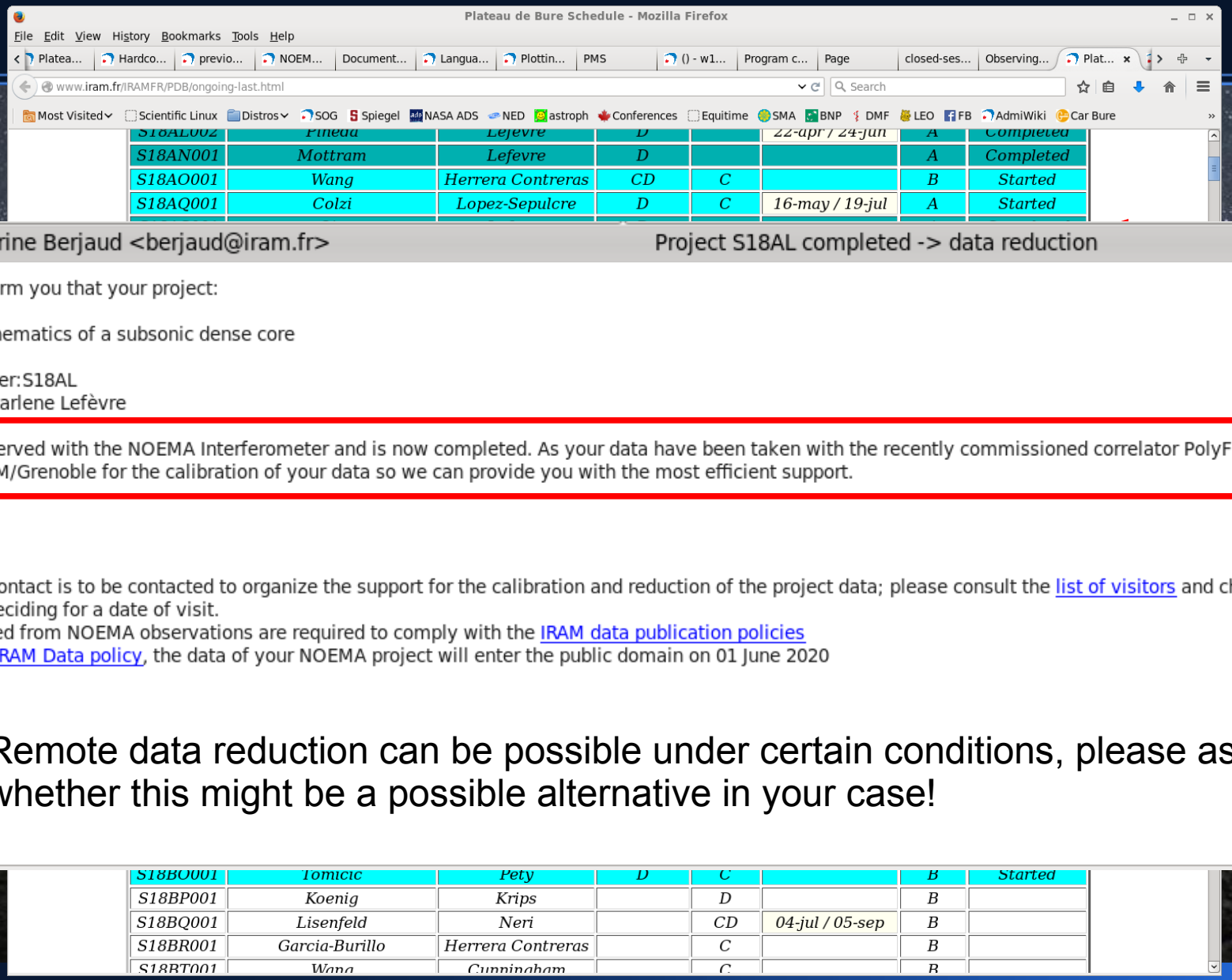

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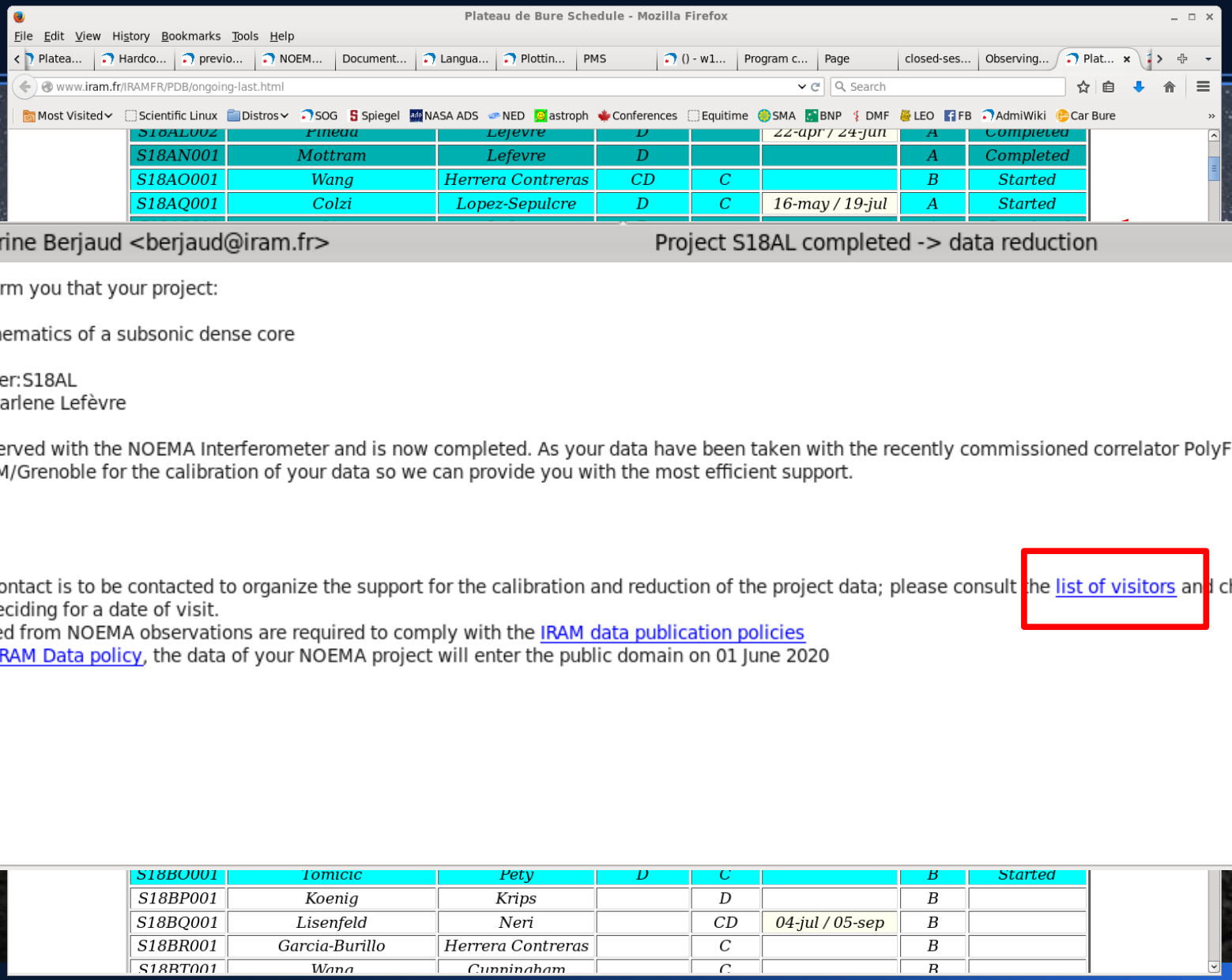

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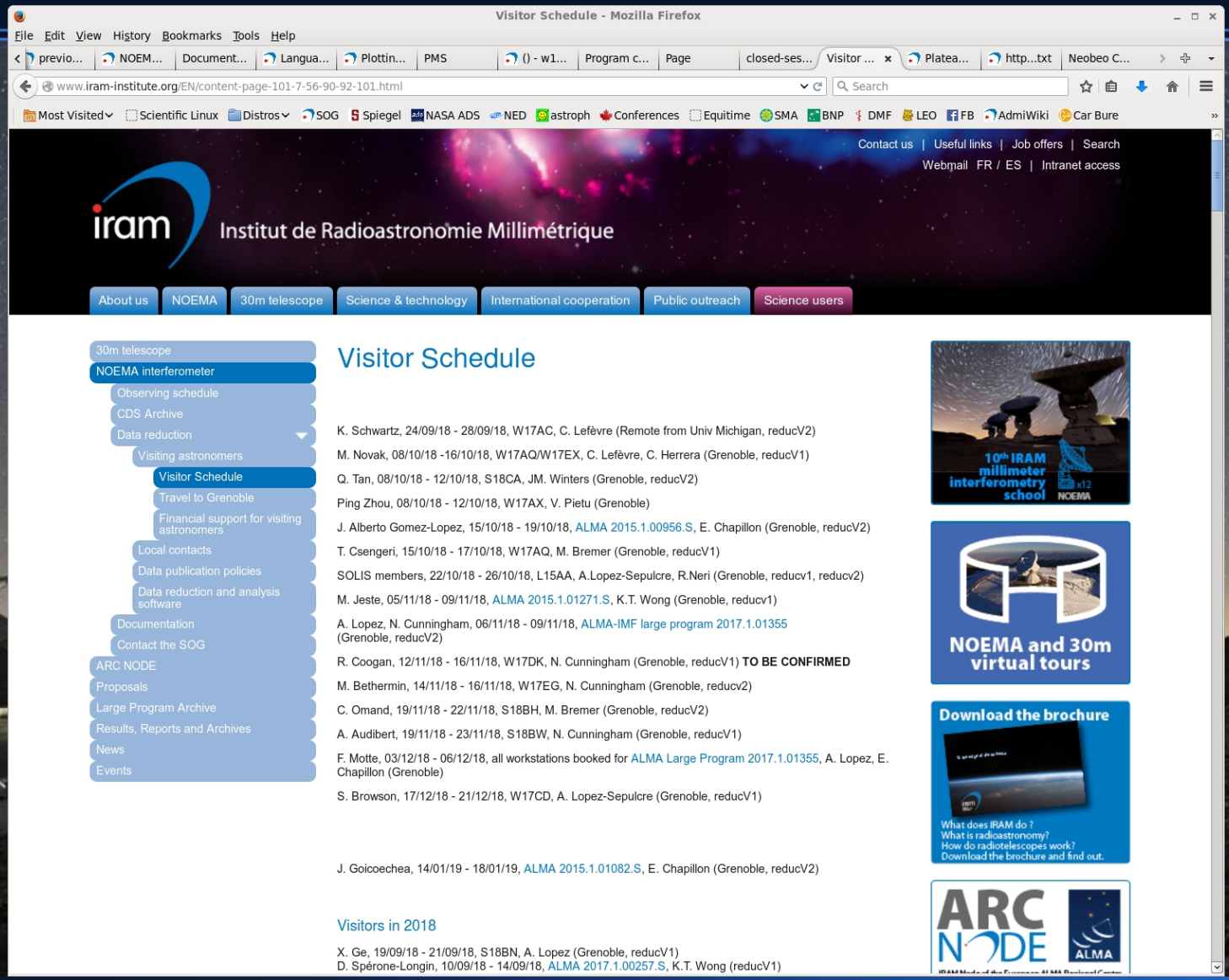

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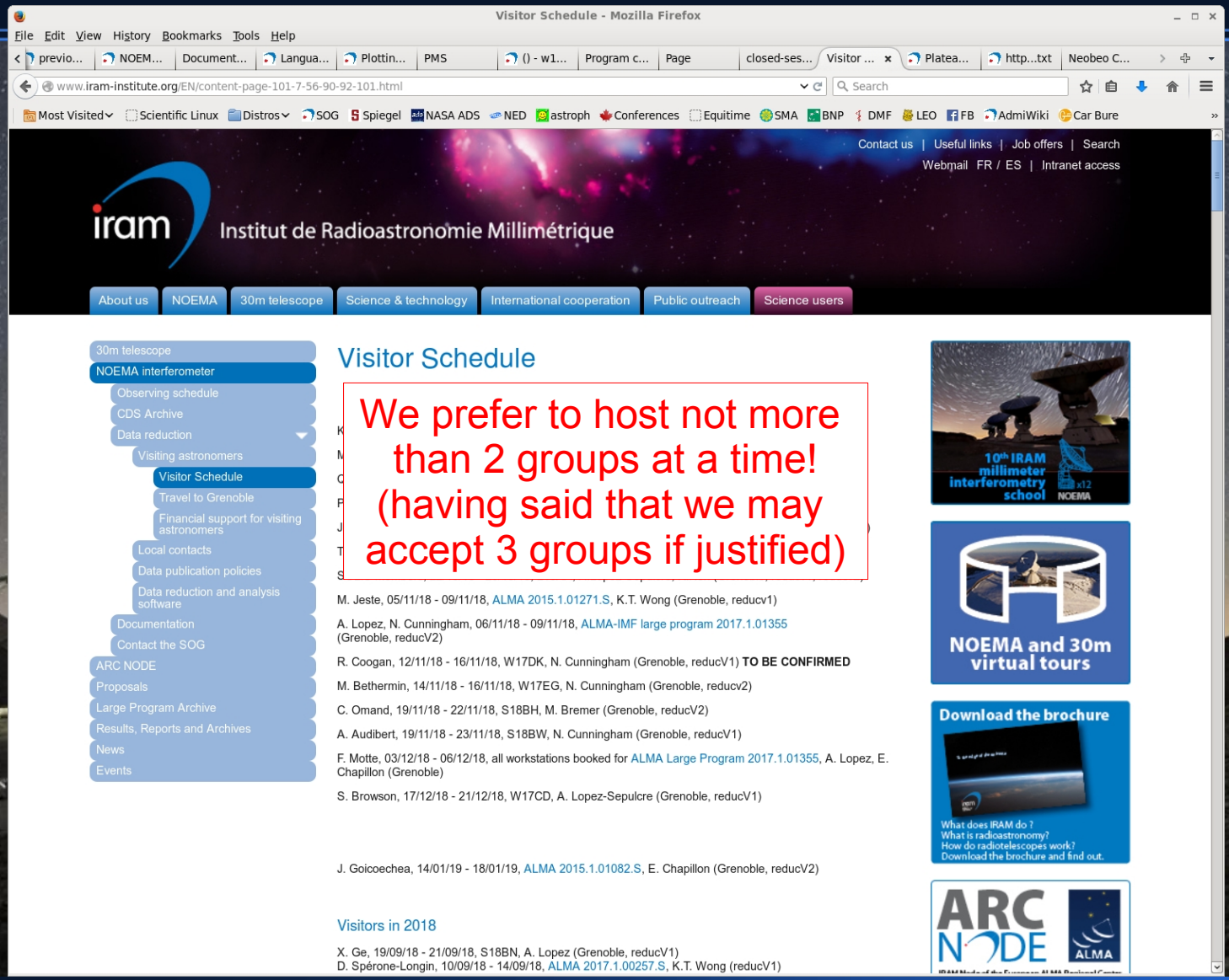
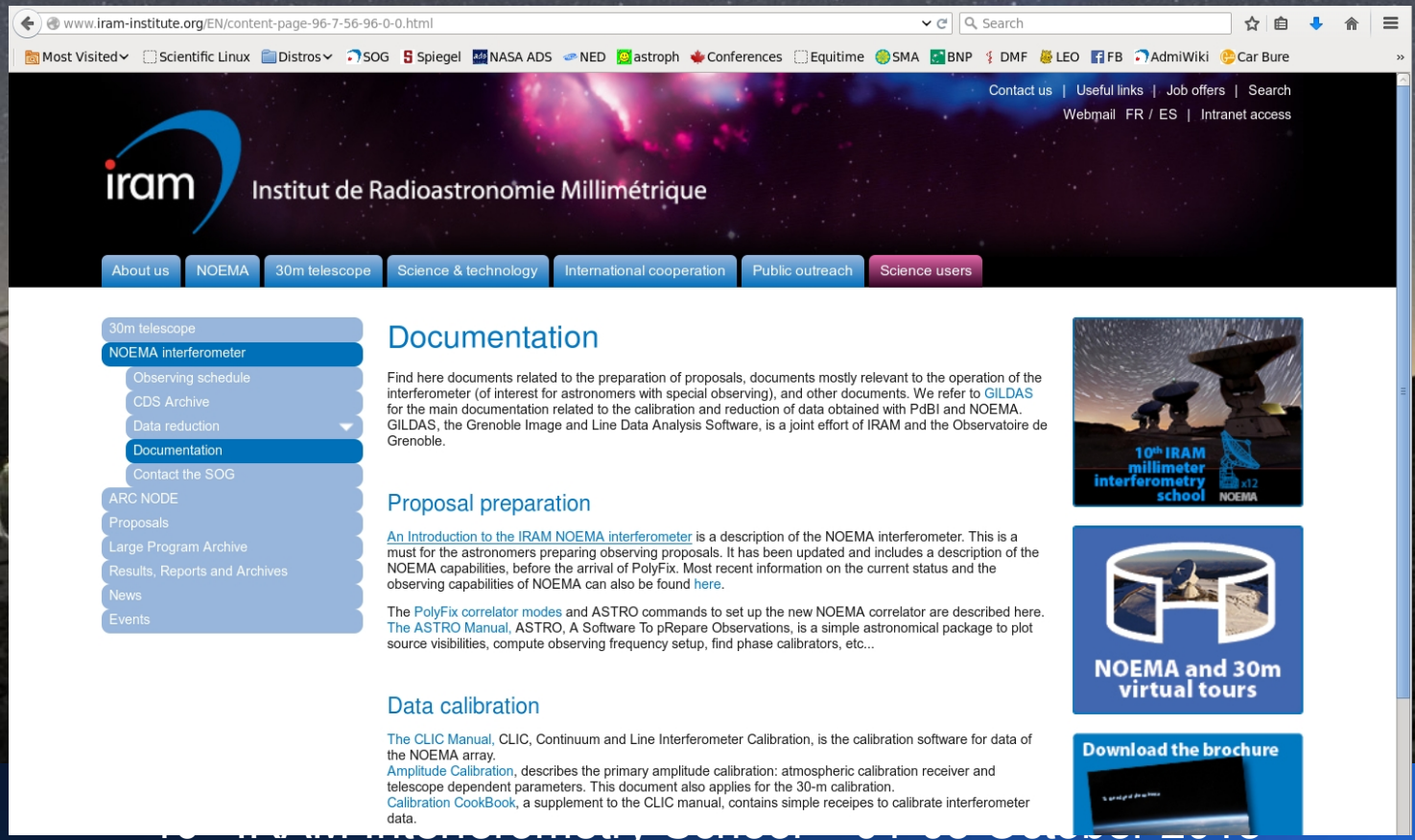


How to use NOEMA

DATA Reduction:

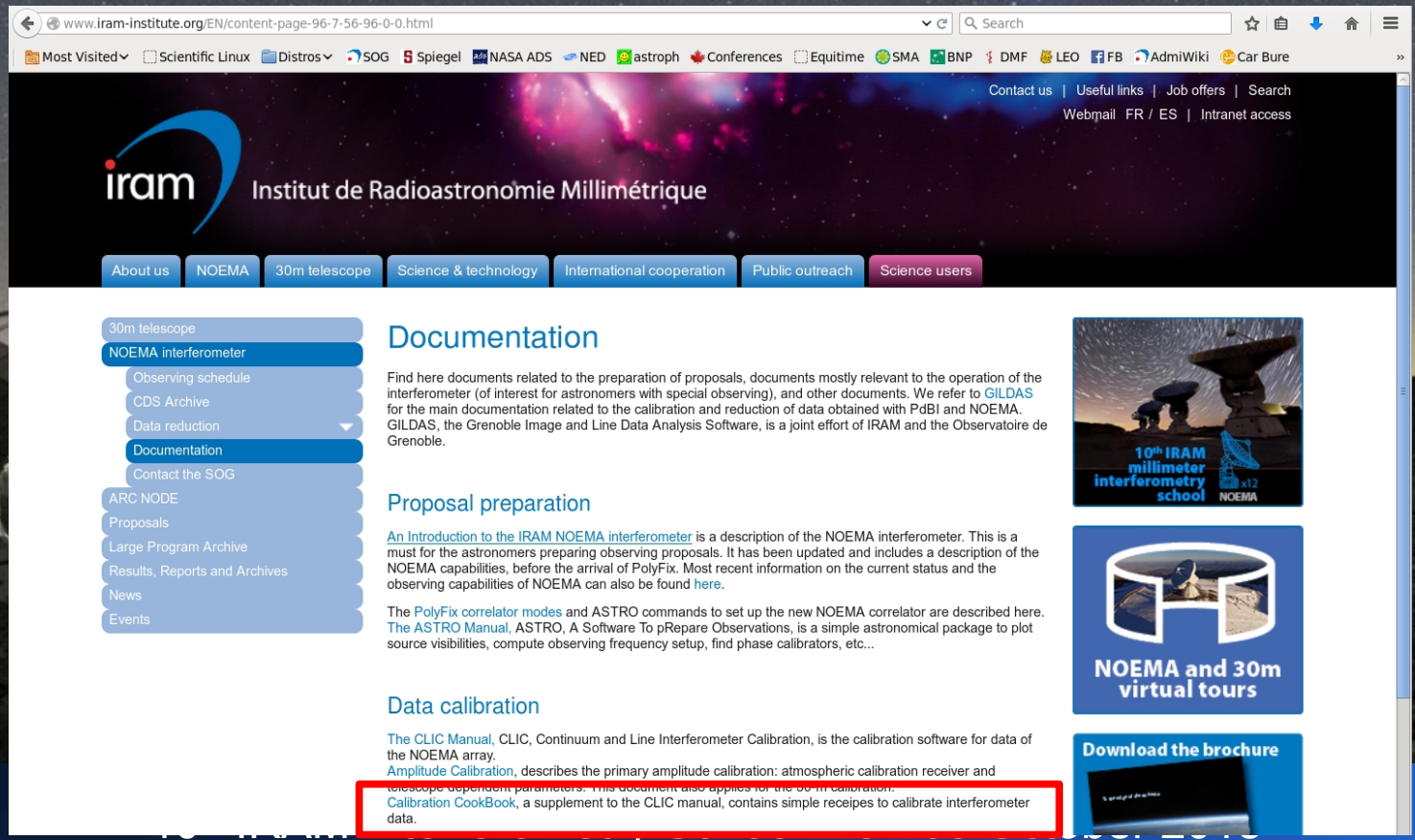
* Please, read the documentation before your visit (especially if your a first time visitor):



How to use NOEMA

DATA Reduction:

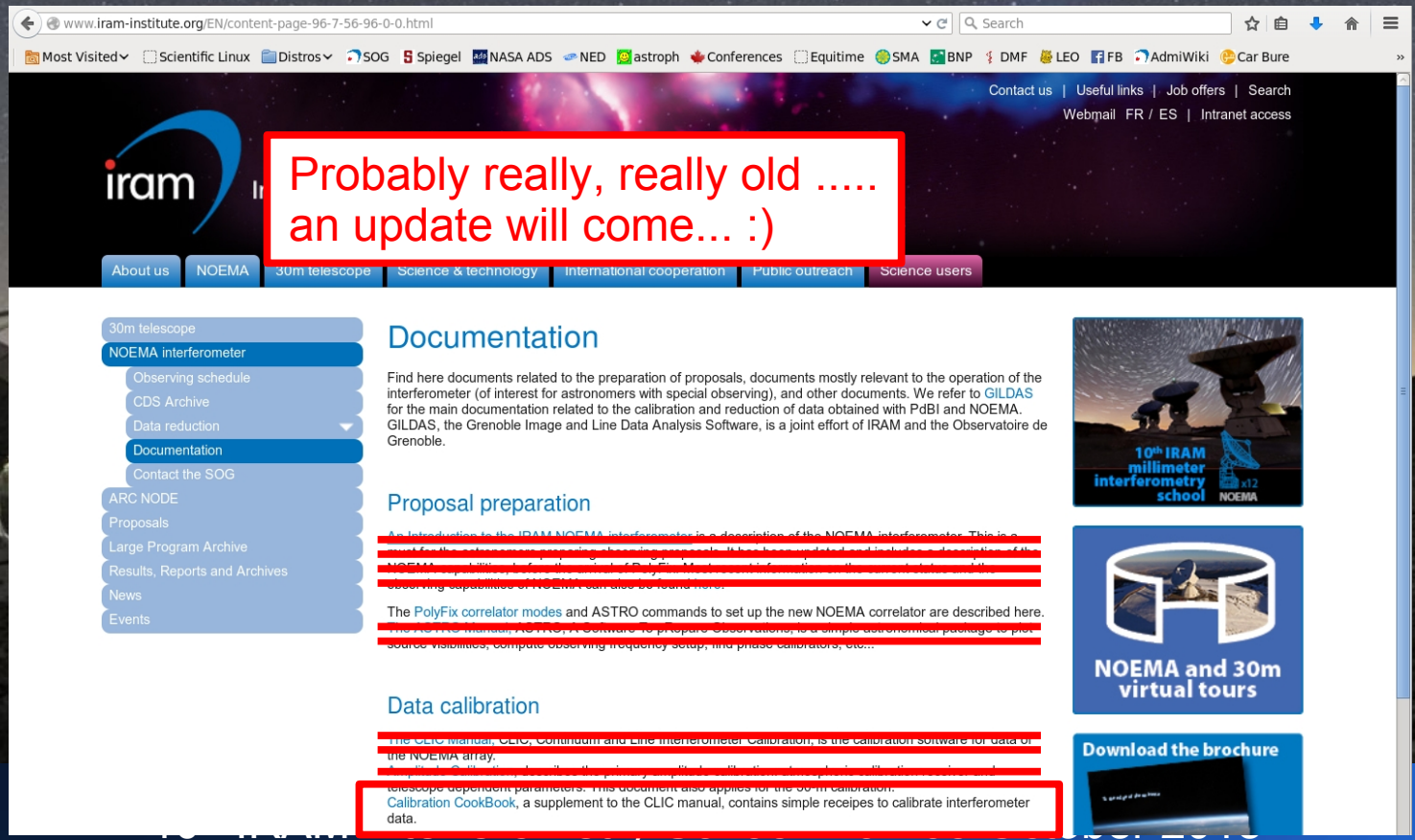
* Please, read the documentation before your visit (especially if your a first time visitor):



How to use NOEMA

DATA Reduction:

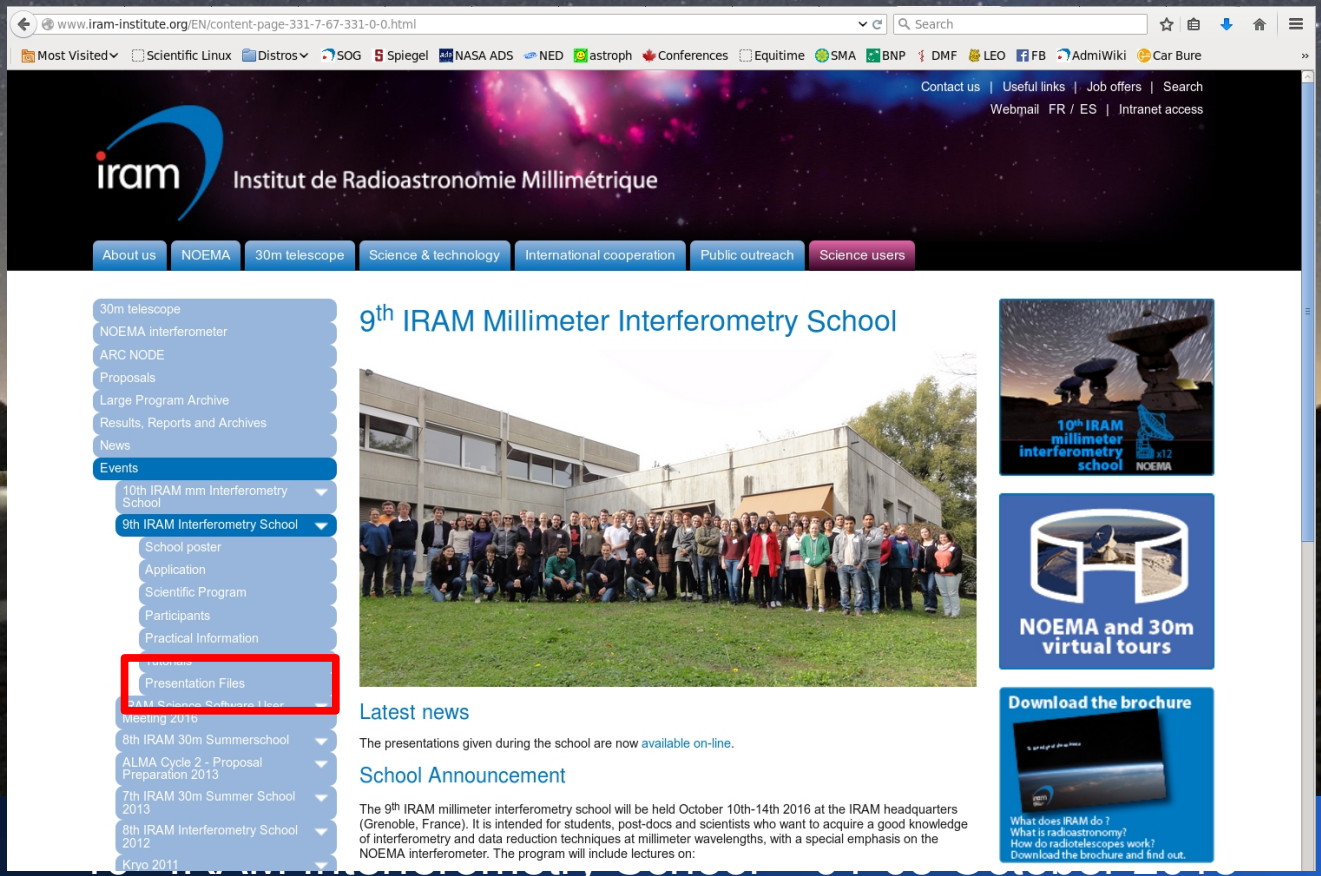
* Please, read the documentation before your visit (especially if your a first time visitor):



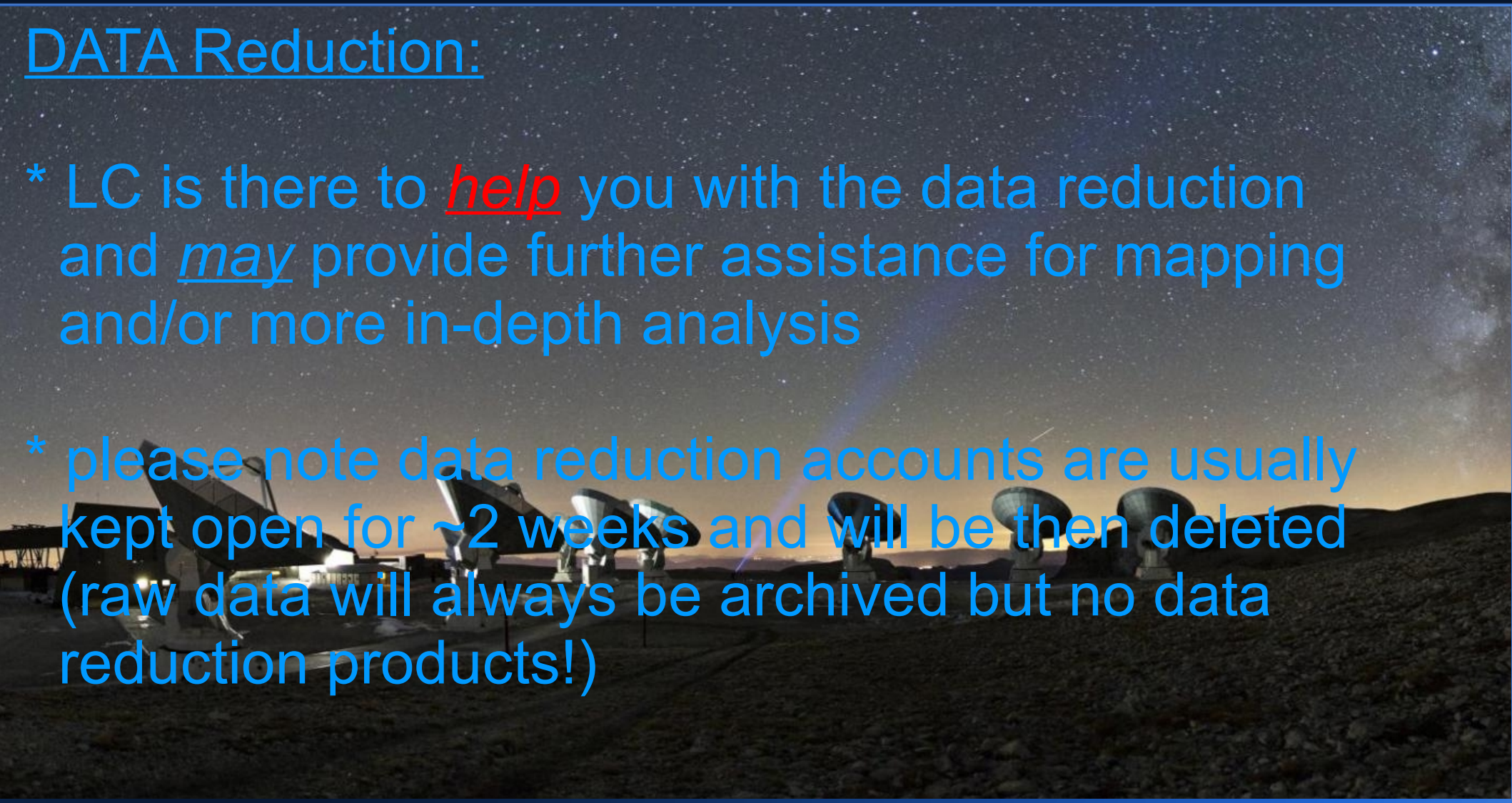
How to use NOEMA

DATA Reduction:

* Please, read the documentation before your visit (especially if your a first time visitor):



How to use NOEMA



DATA Reduction:

* LC is there to ***help*** you with the data reduction and *may* provide further assistance for mapping and/or more in-depth analysis

* please note data reduction accounts are usually kept open for ~2 weeks and will be then deleted (raw data will always be archived but no data reduction products!)

How to use NOEMA

DATA Reduction:

* Travel funds may be available:
  - either for Pis from IRAM partner countries (travel will be paid by IRAM)
  - through RADIONET TA program funding if Certain criteria are fullfilled; contact:

winters@iram.fr

or read more here:

http://www.iram-institute.org/EN/content-page-114-7-56-90-92-114.html

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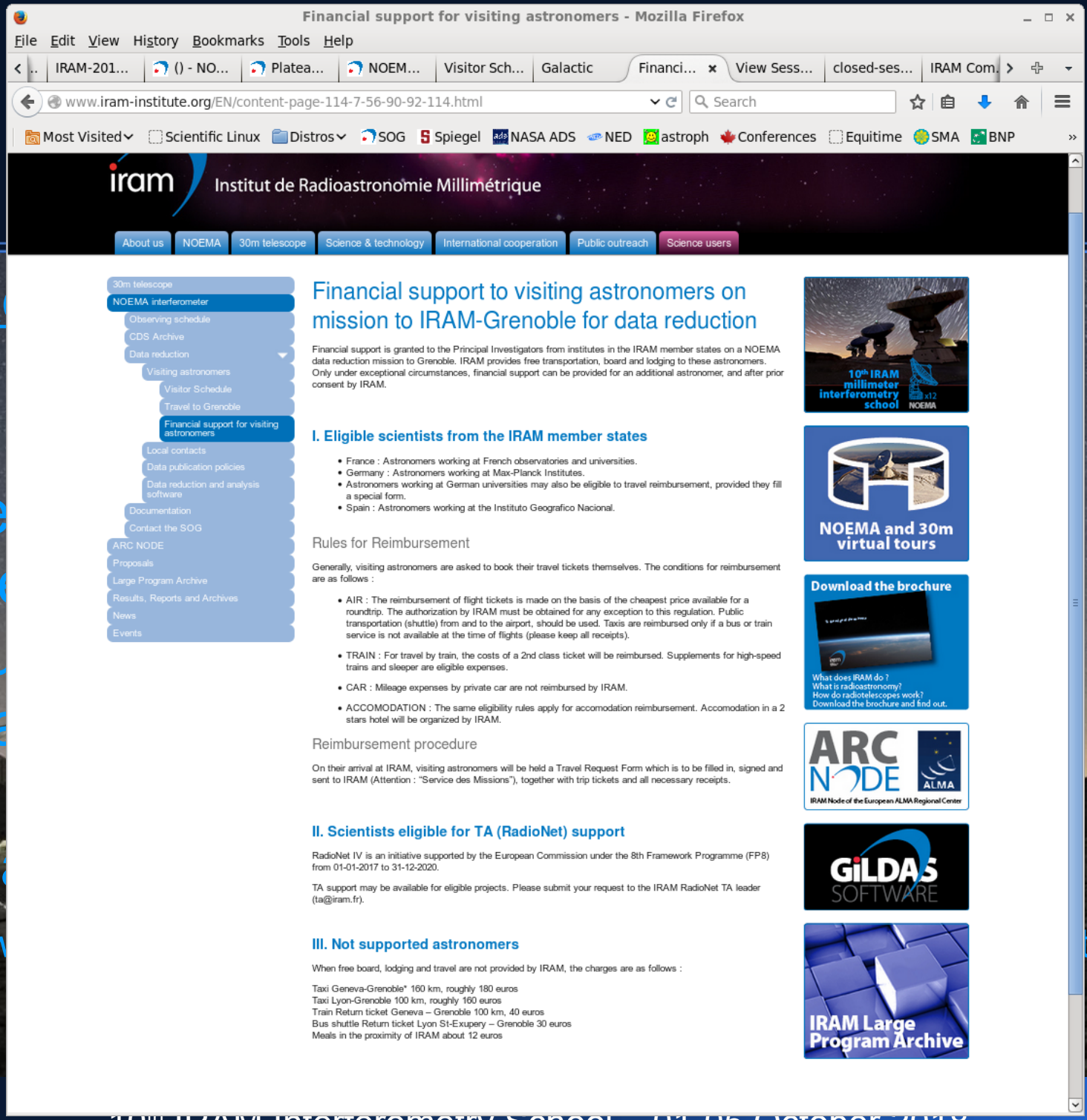
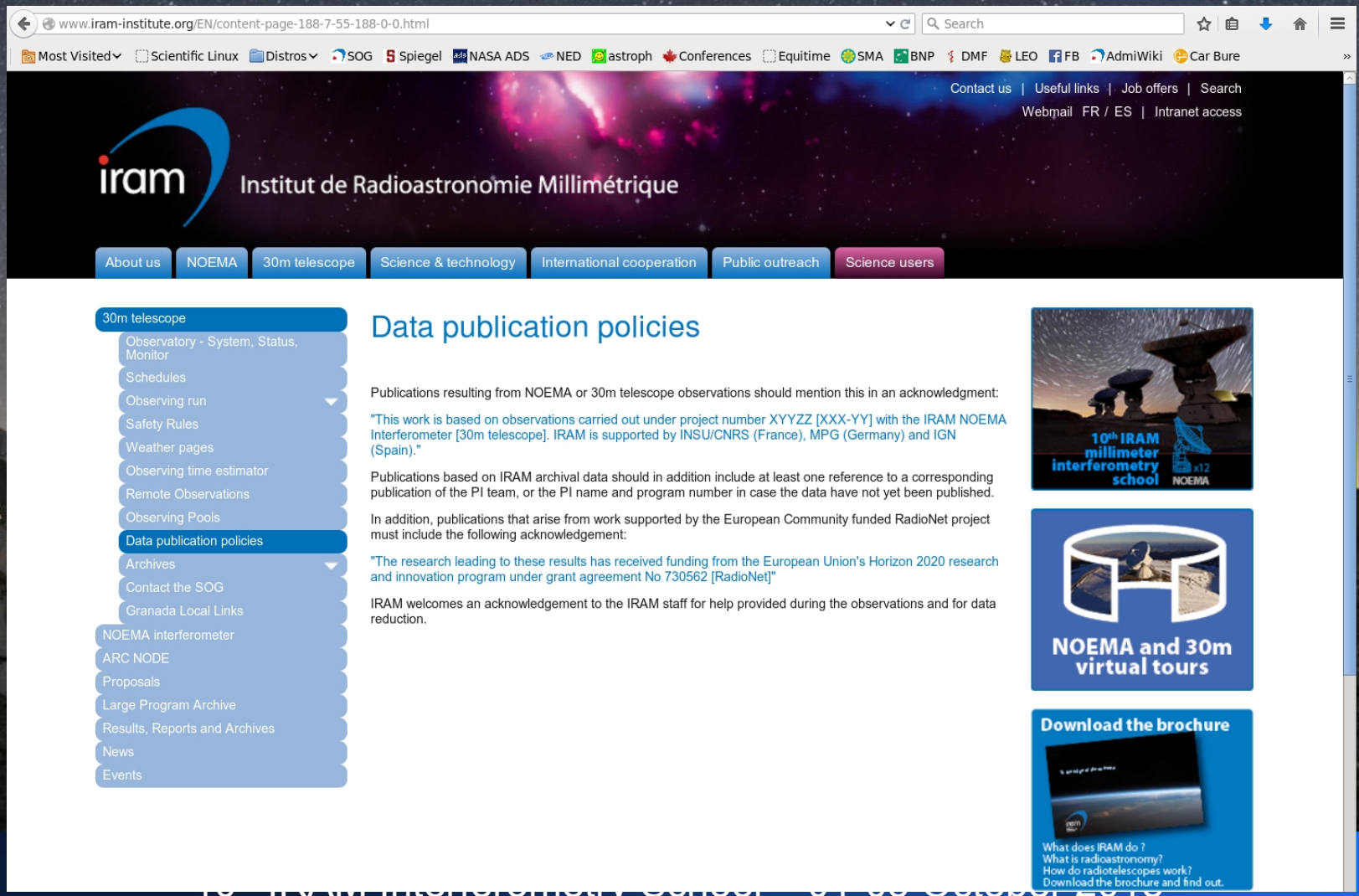


How to use NOEMA

DATA Publication Policy:



How to use NOEMA